

MEASURING SURFACE DISPLACEMENT WITH SYNTHETIC APERTURE
RADAR INTERFEROMETRY (INSAR) AND GAMMA SOFTWARE: A CASE
STUDY FROM KAPOHO, ISLAND OF HAWAII

MASTERS OF SCIENCE
IN
GEOLOGY AND GEOPHYSICS

DECEMBER 2006

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Conventions

Arial bold	Commands and scripts
<i>Arial italic</i>	Filenames and paths
<u>Arial underlined</u>	Hyperlinks to external websites, internal text files, figures etc.
<code>Courier new</code>	Snapshots from actual processing flow showing commands and scripts
<...>	mandatory input parameter
[...]	optional input parameter

This document will be available in html and as a printable version (PDF). The html version of this manual can be accessed using a web browser (e.g., Mozilla Firefox) either online through the PGF website or offline after copying the necessary files to the local machine (CD). Hyperlinks throughout the document can thus be used. The printable version of this manual is equipped with complimentary documents providing a list of figures and linked text files.

A glossary of acronyms and abbreviations can be found in **Appendix A**.

Acknowledgements

I would like to thank my thesis committee, consisting of the chair person Ben Brooks who acted as my academic advisor, my committee members Charles Werner, Steve Martel and Janet Becker, who gratefully replaced Neil Frazer in my committee. Charles is one of the founders of GAMMA Remote Sensing has been an essential part in the completion of this user's guide explaining the overall structure and usage of the software GAMMA. Steve became an important part of my academic and personal life about three years ago and provided me with a huge amount of professional and mental support throughout my academic career. Janet is responsible for the advancement in my math skills and gladly joined my committee to support me and my thesis greatly. Neil played a valuable role in the decision making for my Master's route, I received a lot of encouragement and human support from him. Hao Zhou, Matt Patrick and Arjun Aryal were essential in getting started and further improve the usage of GAMMA. Paco Gomez also helped in the development of routines to efficiently process radar data. Mike Poland provided support as well as additional data. Mark Munneke helped out a great bunch as our computer technician. James Foster has been endless source of geophysical, technical and mental help. And last but not least, Joseph Shacat, my lovely husband without whom I would have never gotten this far. He kept me on track and reminded me often what life is really all about.

1. Introduction

The principle product of this Plan B Masters is a user's guide for GAMMA InSAR software to produce deformation maps. We demonstrate the software in a case study of Kapoho, located in the Puna district on eastern tip of the island of Hawaii (from here on referred to as the Big Island). This document represents a customized, updated lab guide that should be used in conjunction with the GAMMA software documentation.

Land surface deformation may be associated with a variety of processes that include, for example, faulting, volcanism, glaciation and deglaciation, landsliding, ground water pumping, and mining (Figure 1: [Deformation PDF 9.6 MB](#)). Volcanic processes include seismic events, volcanic eruptions, magma-related subsidence and intrusion. These processes can be destructive to life and property. By measuring and modeling the displacement at the surface of the Earth we can learn more about the processes acting in the subsurface. Monitoring of these potentially hazardous processes can be accomplished using Interferometric Synthetic Aperture Radar (InSAR).

The Kilauea lower east rift zone extends subaerially about 60 km from the summit of Kilauea Volcano to the eastern tip of the Big Island [*Delaney et al.*, 1998]. This area has experienced gradual subsidence of about -1.4 cm/year since the M 7.2 Kalapana earthquake in 1975 increasing the tidal pond-area significantly. Coastal communities are particularly susceptible to hazards associated with sea-level change. Causes of sea-level change include or are a combination of global sea-level rise and relative land subsidence [*Caccamise et al.*, 2005]. The community of Kapoho is situated in relatively flat terrain south of Kapoho Bay (Figure 2a: [Kapoho PDF 5.4 MB](#)). Coastal subsidence is drastically affecting two residential subdivisions, the Kapoho beach lots and the Kapoho Vacationland. As a result, these properties are flooded during ordinary high water stands (Figure 2b: [High water](#)) and major health hazards related to the sewage system are developing. It has therefore become important to further investigate the ground motion in that area using currently available leveling and GPS data and incorporating InSAR.

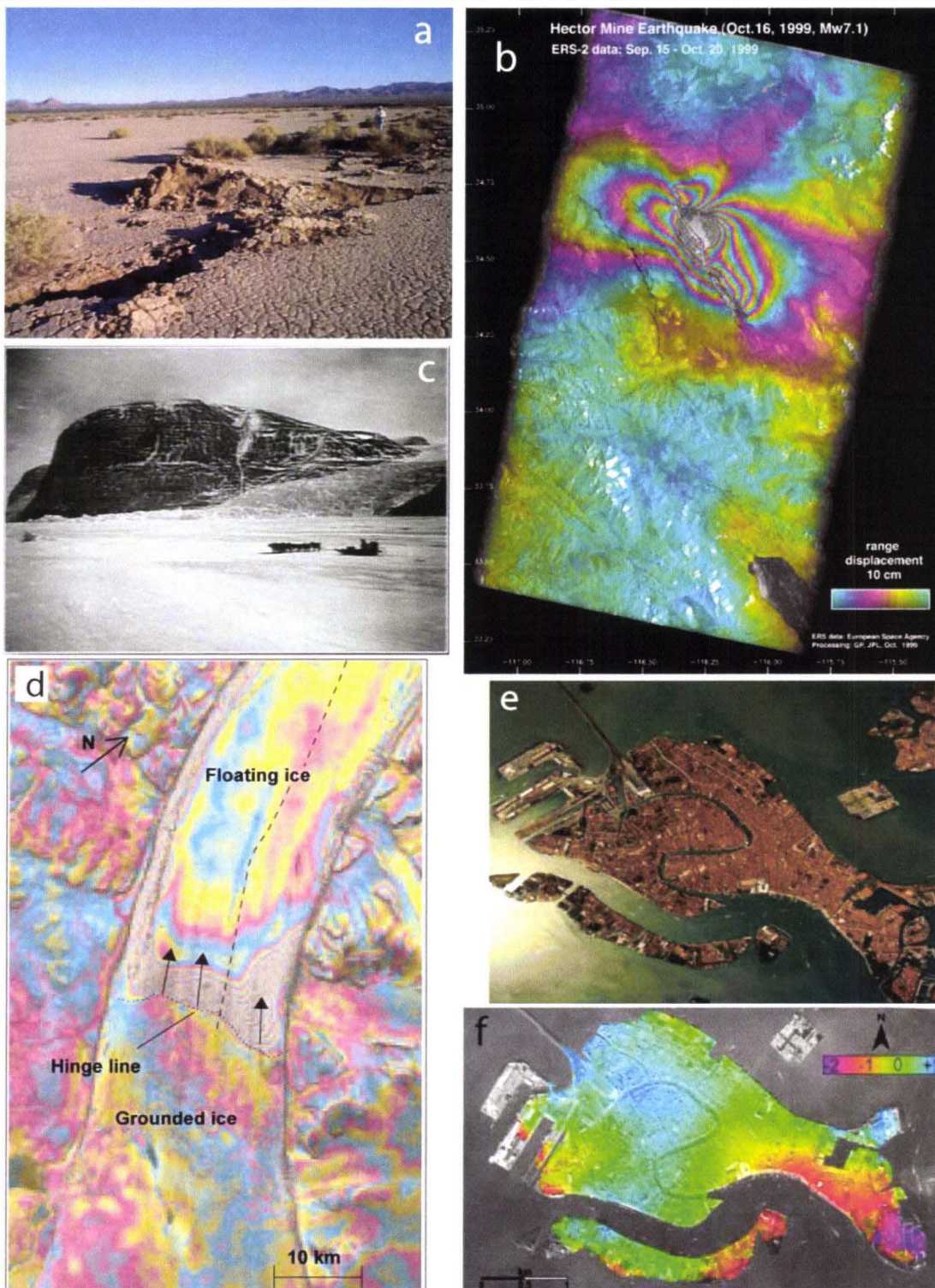


Figure 1: Examples of deformation and interferometric processing results. (a) Surface expression of the Hector Mine Earthquake (courtesy of Chris Walls), (b) Co-seismic deformation pattern associated with the Hector Mine Earthquake (courtesy of G. Peltzer, 2001), (c) Petermann glacier (courtesy of Lauge Koch), (d) Tidal displacements of Petermann glacier (one color-cycle corresponds to 28 mm of range displacement, courtesy of E. Rignot, 1997), (e) Aerial view of Venice, Italy (www.toon.heindl-internet.de), (f) Land subsidence rate of Venice between 1992 and 1996 in mm/year (www.gamma-rs.ch).

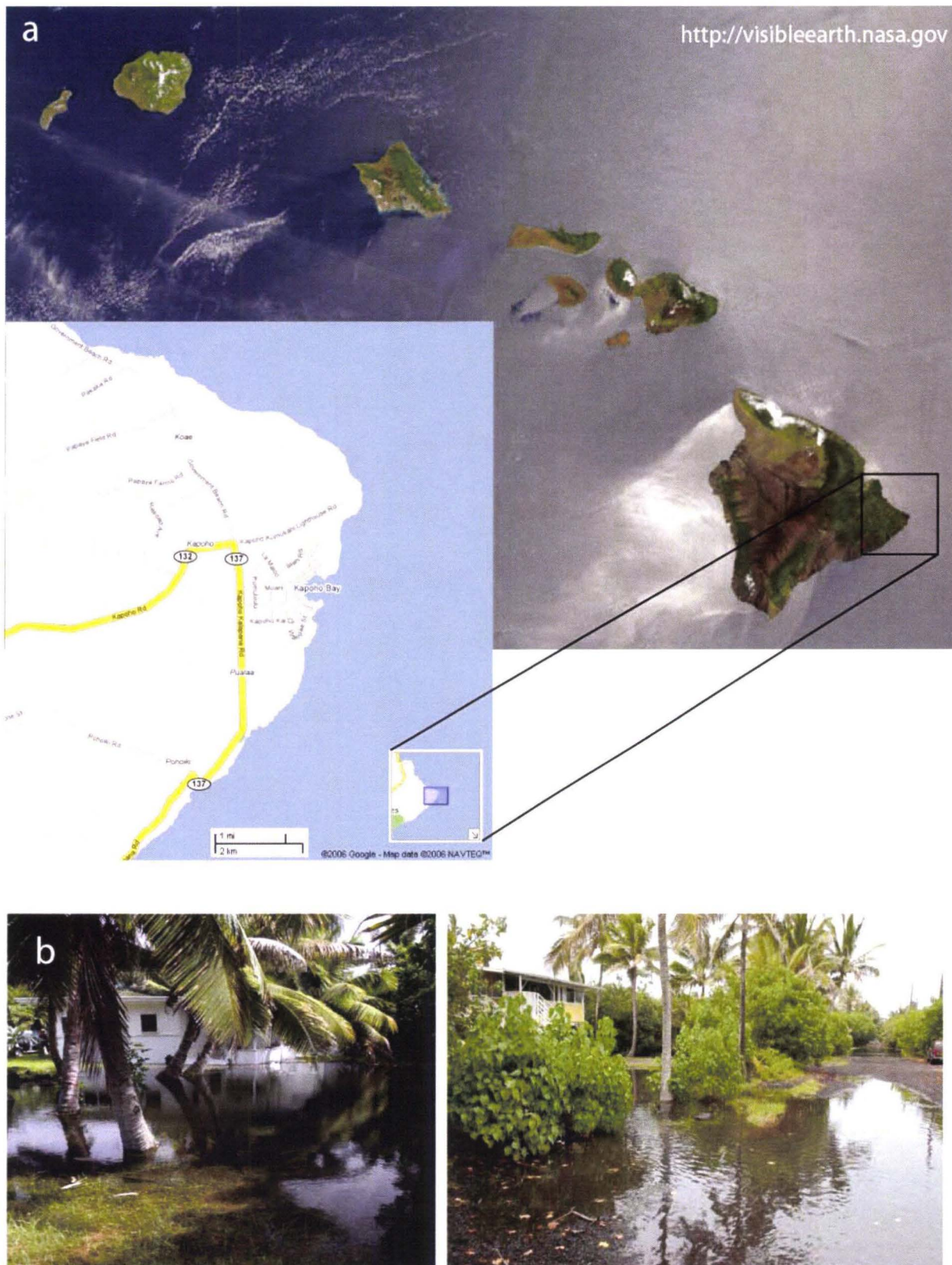
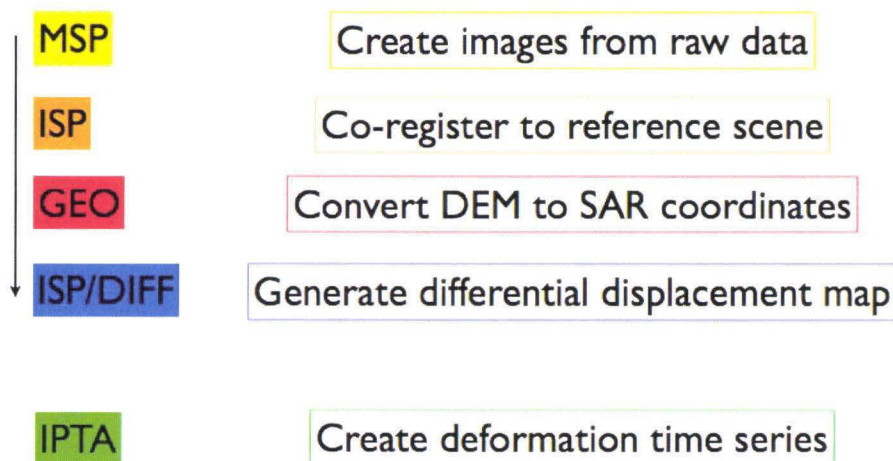


Figure 2: (a) Location of Kapoho, Puna District, Island and State of Hawaii, (b) Residential areas at high tide (courtesy of D. Hwang).

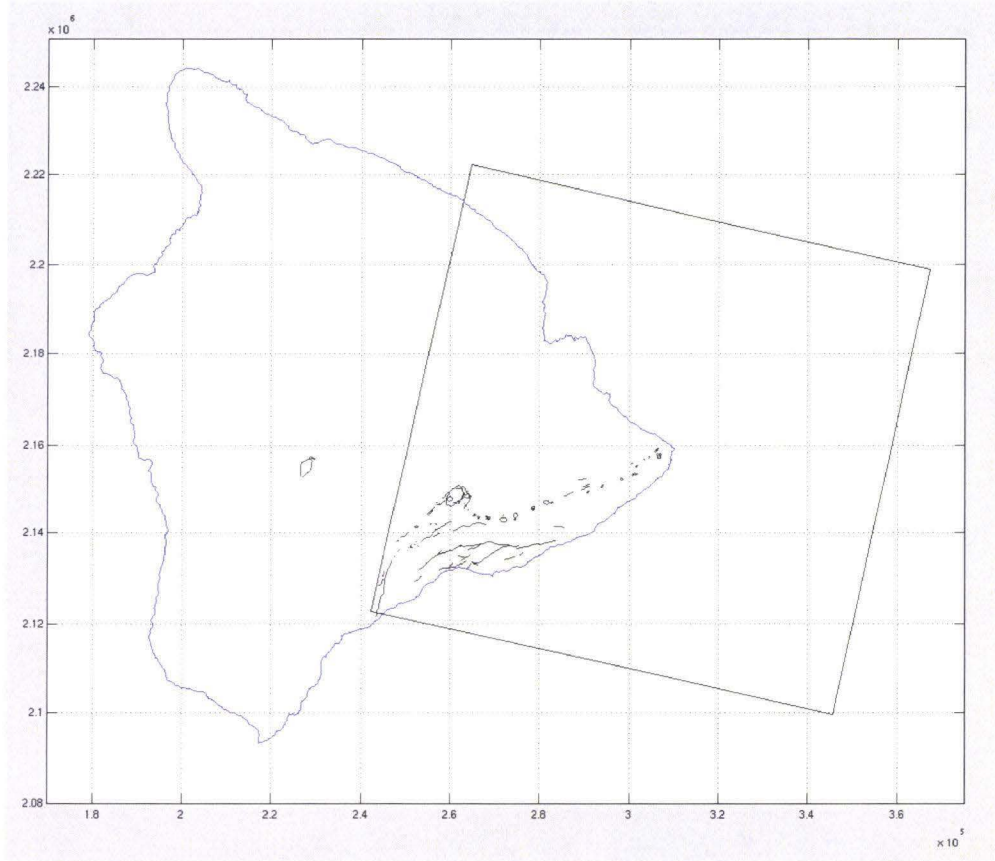
2. GAMMA Processing Work Flow

[GAMMA software](#) is capable of processing synthetic aperture radar (SAR) data from a variety of Earth orbiting satellite platforms. The software supports processing of raw SAR data and creation of final interferometric products such as digital elevation models (DEM), deformation maps, and land-use maps. GAMMA contains the following packages: Modular (MSP) and Interferometric SAR processing (ISP), Differential SAR Interferometry and Geocoding (DIFF&GEO), Interferometric Point Target Analysis (IPTA), and a Display toolbox (DISP). The flowchart below briefly illustrates the modules, whereas at the beginning of each section one can find a more detailed flowchart of the corresponding module. This manual will describe most of these toolboxes and demonstrate how to use them in the context of an example from Kapoho, Hawaii using data from the Envisat platform.

GAMMA Processing Flow



The Environmental Satellite Envisat, maintained by the European Space Agency (ESA @ <http://earth.esa.int/>), acquires radar data approximately once per month using an Advanced Synthetic Aperture Radar (ASAR) instrument (for information on other satellites refer to **Appendix B**). The following steps describe a typical, complete work flow in order to process raw Envisat ASAR image mode data to make a final map of surface displacement (see flowchart above). Radar images or scenes are recorded as a sequence of frames in a track, which corresponds to the satellite's ground coverage (for details see **Appendix C**). The frame 3213 from the descending track 429 covers the eastern portion of the Big Island of Hawaii. The 21 scenes for this specific track and frame were acquired in image swath 2, with incidence angles ranging from 19.2° to 26.7° (see **Appendix D** for data availability). The ground coverage of this specific frame is displayed in the figure below.



GAMMA provides different modules in order to fully support interferometric processing. The processing flows will go through most of the modules and show how certain techniques may be applied and how specific commands are run. The manual is accompanied by a processing log (README file) containing individual steps ([README link](#)) utilized throughout the processing. The README includes every command and script that is run in chronological order. The manual provides short explanations of scripts and commands as well as examples of various products such as images of interferograms, displacement maps, input parameter files etc. More information can be accessed on line through www.gamma-rs.ch. The existing [html software documentation](#) provided by GAMMA is used as the foundation. The GAMMA user guide is divided into the different modular packages and will be referred to in links.

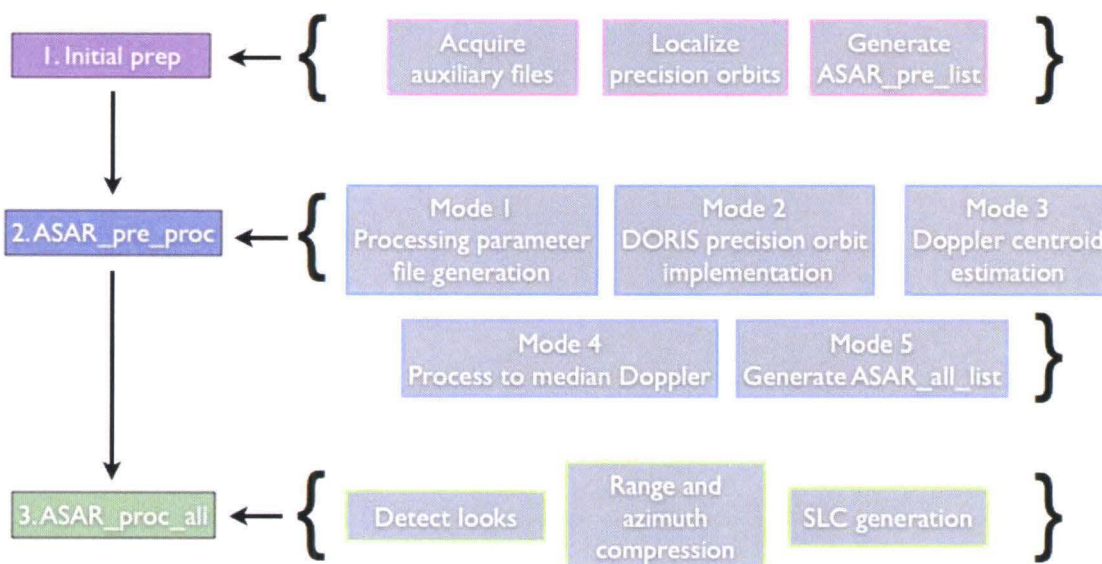
Please note that commands and scripts in GAMMA are all case-sensitive. Every command and script is run inside the processing directory, unless otherwise noted. Scripts are usually used to run commands on the entire list of available data scenes. Please refer to **Appendix E** for a list and syntax of commands and scripts used in this manual.

Radar acquisitions are referred to as scenes or layers in this manual. The entire list of acquired scenes is often referred to as the “data stack”. However, the “stacking” process (see ISP/DIFF module) estimates an average linear velocity for each point in the layer throughout the entire “data stack”. When this process has been carried out, we refer to the result as the “stacked data” containing of one single map of averaged velocities. Interferograms are also called ‘pairs’ and are layers of the interferometric “data stack”.

MSP – [Modular SAR Processor](#)

In MSP processing raw radar data from current space or airborne sensors is converted to single-look complex images (SLC). The data may be downloaded from [WInSAR](#) (online joint data archive) or ordered from ESA (**Appendix D** provides information about data acquisition). Firstly, a set of auxiliary data files such as instrumentation, calibration and precision orbits files is applied to the raw data. The GAMMA modular SAR processor then allows for accurate range-Doppler algorithms, radiometric calibration to perform basic data conditioning and thereby preserves the phase for interferometric processing.

The following flow chart describes the basic processing flow where each step is going to be illuminated in greater detail throughout this chapter. Processing scripts consist of a series of different commands and sometimes can be run in various modes (i.e., mode 0, mode 2 etc.), utilizing a different sequence of commands. Text displayed in the font `courier new` refers to either the content of a file (e.g., `ASAR_pre_list`) or the script/command that is run and their standard output. The input arguments are separated by spaces. For an explanation of the input/output parameters please refer to **Appendix E** (index of commands/scripts and their syntax). A command history can be found in the [README](#) file and in the example of the MSP application in **Appendix H**.



MSP-1 Initial Preparation

In order to run the initial commands a few points need to be checked off first:

- Current location of auxiliary files on processing machine is determined
- Auxiliary files are up to date
- Current location of precision orbits on processing machine is known
- Precision orbits are up to date
- Preprocessing list is generated (necessary for running next script, see MSP-1.2)

The following few paragraphs explain these files and provide download information as well as how to use them.

MSP-1.1 Auxiliary Files

Auxiliary files, (i.e., instrument characterization, external calibration and precise orbit files) are necessary input parameters for the MSP preprocessing and can be obtained through ESA (see **Appendix F**). It is important to determine exactly which auxiliary file is appropriate for a specific date since ESA generates updated versions while keeping previous ones. The instrument characterization files are only updated if the instrument operating parameters change. External calibration files, in contrast, are updated every six months. The most recent processing date (the date when the auxiliary file was created) must be used and the correct date range for the SAR scene must be determined.

Instrumentation characterization files (*ASA_INS_**) contain parameters that characterize the ASAR instrument, such as look-up tables and other ground processing factors. The files contain subsets of the instrument characterization database. The appropriate filename including the full path is specified in a list (*ASAR_pre_list*) during the MSP preprocessing. For example, the date of a SAR raw data file is **February 5th 2003** (20030205) and the following list of *ASA_INS_** files is available,

```
ASA_INS_AXVIEC20051219_161945_20030211_000000_20061231_000000
ASA_INS_AXVIEC20031212_105841_20021017_162400_20030204_110000
ASA_INS_AXVIEC20031209_113259_20021030_110000_20030211_000000
ASA_INS_AXVIEC20031202_122530_20020815_131000_20031017_162400
```

where for the first example file column three corresponds to the processing date (*20051219) and columns five and seven (20030211_*_20061231) represent the date range that the specific file is to be used for.

For a data scene from February 5th 2003 one would choose the following instrumentation characterization file: *ASA_INS_AXVIEC20031209_113259_20021030_110000_20030211_000000*, where 20031209 is the most recent processing date and 20030205 is in the date range between 20021030 and 20030211.

External calibration files (*ASA_XCA_**) contain calibration data such as the external calibration scaling factor and the in-flight elevation pattern estimates. These files are required to extract the antenna pattern as measured on the ground and they are used to compensate for the antenna gain across the swath. Example filename: *ASA_XCA_AXVIEC20050803_150715_20030211_000000_20030601_000000*, where 20050803 is the processing, 20030211 the start date and 20030601 the end date. The same selection procedure as for the instrumentation characterization files is applied to find the appropriate file if multiple processing dates or date ranges are available.

Precise orbit state vectors (*DOR_VOR_**) are obtained from the DORIS (Doppler Orbitography and Radiopositioning Integrated by Satellite) instrument. DORIS Doppler-shift data produces preliminary orbit estimates and precise orbit state vectors. Precision orbits often improve baseline estimates allowing a better simulation of the topographic phase, hence resulting in more precise differential interferograms.

Examples of DORIS vector files:

DOR_VOR_AXVF-P20030415_154900_20030203_215528_20030205_002328

DOR_VOR_AXVF-P20030415_155000_20030204_215528_20030206_002328

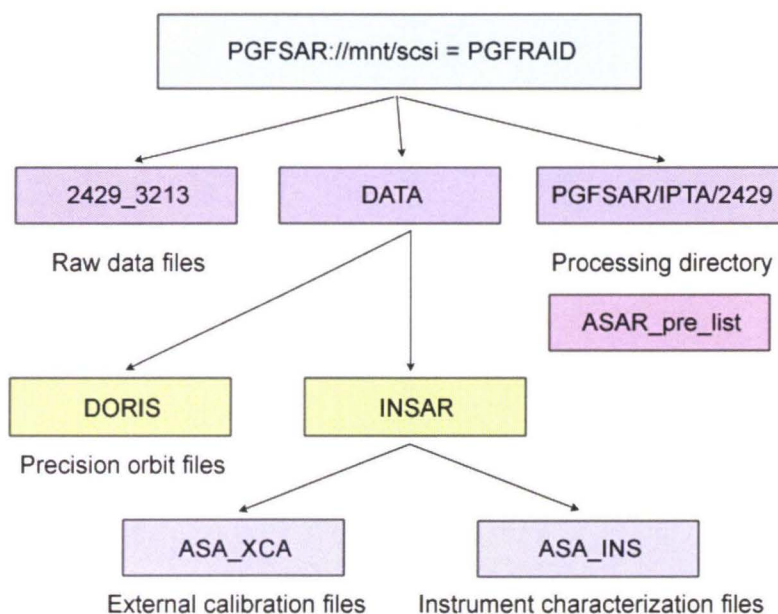
DOR_VOR_AXVF-P20030415_155100_20030205_215528_20030207_002328

For the **February 5th 2003** scene *DOR_VOR_AXVF-*

P20030415_155000_20030204_215528_20030206_002328 would be the appropriate file to use because 20030205 is in the range *20030204_*_20030206*.

An alternative set of precision orbits ([DELFT orbits](#)) is available through the University of Delft, where DORIS tracking data and SLR (Satellite Laser Ranging) are [combined](#).

The schematic below shows the location of the auxiliary files on the processing machine PGFSAR with respect to the local processing directory (2429_3213).



MSP-1.2 ASAR_pre_list

GAMMA's ASAR raw data preprocessing script requires an input list. This ASAR_pre_list consists of four columns per entry listing the ASAR raw data and associated calibration data:

Column 1: Level 0 (L0) ASAR raw data set, filename includes full path

Column 2: ASAR instrument characterization data file, filename includes full path

Column 3: ASAR external calibration data file, filename includes full path

Column 4: DORIS state vector file, filename does NOT include path

For every available data scene four columns have to be filled with (1) the correct raw data filename, (2) instrument characterization and (3) external calibration files, all with full path information, and (4) the DORIS filename without the path. The four columns must be in one line in a text editor.

Here are the first two entries for two data scenes in an [ASAR_pre_list](#):

```
/mnt/scsi/2429_3213/ASA_IM_OCNPDE20050112_202041_000000152033_00429_15012_5852.N1
/mnt/scsi/DATA/INSAR/ASA_INS/ASA_INS_AXVIEC20051219_161945_20030211_000000_20061231_000000
/mnt/scsi/DATA/INSAR/ASA_XCA/ASA_XCA_AXVIEC20060223_133247_20050101_000000_20050914_000000
DOR_VOR_AXVF-P20050303_101000_20050111_215528_20050113_002328
/mnt/scsi/2429_3213/ASA_IM_OCNPDE20050914_202045_000000152040_00429_18519_2347.N1
/mnt/scsi/DATA/INSAR/ASA_INS/ASA_INS_AXVIEC20051219_161945_20030211_000000_20061231_000000
/mnt/scsi/DATA/INSAR/ASA_XCA/ASA_XCA_AXVIEC20060223_133247_20050101_000000_20050914_000000
DOR_VOR_AXVF-P20051216_153300_20050913_215528_20050915_002328
```

Two lines each four columns long are shown above, appearing as eight lines because of the limited document width. It is essential that the end of every line after the fourth column is marked by a carriage return (ENTER) and that no empty lines exist. In the above example the carriage return is only behind the fourth column (*DOR_VOR*20050113_002328* and *DOR_VOR*20050915_002328*).

After generation of the *ASAR_pre_list*, stored inside the working directory, the script [ASAR_pre_proc](#) is ready to be run.

MSP-2 ASAR Raw Data Preprocessing

In the MSP raw data preprocessing the *ASAR_pre_list* is used in the GAMMA script [ASAR_pre_proc](#) that is run in modes 1 through 5.

The GAMMA script [ASAR_pre_proc](#) should be sequentially run in the following five modes:

- 1: Create processing parameter files and unpack raw data
- 2: Extract and interpolate DORIS state vectors and update MSP processing parameter files
- 3: Estimate Doppler centroid
- 4: Set value a (e.g., median Doppler) in processing parameter files for a keyword:value pair (optional)
- 5: Generate processing list for use by [ASAR_proc_all](#)

Mode 1: Create processing parameter files and unpack raw data ([log](#))

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_1.log ASAR_all_list 1
```

Within [ASAR_pre_proc](#) the following commands are run in mode 1 for the entire list of scenes specified in the *ASAR_pre_list*:

ASAR_XCA, ASAR_IM_proc, set_value, ORB_prop

- Interpretation of ASAR external calibration data file, used to extract appropriate antenna pattern (the [antenna](#) parameter indicated on command line)

```
ASAR_XCA /mnt/scsi/DATA/INSAR/ASA_XCA/ASA_XCA_AXVIEC20060223_133247_20050101_000000_20050914_000000 raw/ASAR_20050914_IS2_VV.gain IS2 VV
```

- MSP preprocessing for ASAR L0 image mode data

[ASAR_IM_proc](#) generates MSP [sensor parameter](#) ([SAR_par example](#)) and [processing parameter](#) files

([PROC par example](#)) from Level 0 (raw) ASAR image mode data. This program also reformats the [raw](#) SAR signal data to be compatible with the MSP. The entire level 0 data set is converted to 8-bit I/Q unsigned binary complex samples. I/Q refers to the real in-phase (I) component and the imaginary quadrature (Q) component of the complex radar return. An unsigned numeric variable can only represent positive numbers.

```
ASAR_IM_proc /mnt/scsi/2429_3213/ASA_IM__OCNPDE20050914_202045_000000152040_00429_18519_
2347.N1 /mnt/scsi/DATA/INSAR/ASA_INS/ASA_INS_AXVIEC20051219_161945_20030211_000000_20061231_
000000 raw/ASAR_20050914_IS2_VV.par raw/p20050914.slc.par raw/20050914.raw raw/ASAR_20050914_
_IS2_VV.gain
```

- Set title in the processing parameter file

```
set_value raw/p20050914.slc.par raw/p20050914.slc.par "title" "ASA_IM__OCNPDE20050914_202045_
000000152040_00429_18519_2347.N1 ASAR_20050914_IS2_VV.par"
```

```
keyword found: title      old value: ASA_IM__OCNPDE20050914_202045_000000152040_00429_18519_
2347.N1 new value: ASA_IM__OCNPDE20050914_202045_000000152040_00429_18519_2347.N1 ASAR_
20050914_IS2_VV.par
```

- Calculate additional state vectors using orbit propagation and interpolation

```
ORB_prop raw/p20050914.slc.par 7 10
```

Mode 2: Extract and interpolate DORIS state vectors and update MSP processing parameter files ([log](#))

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_2.log ASAR_all_list 2
```

The following command is run in mode 2 for the entire list of scenes: **DORIS_proc**

- Extract Envisat DORIS state vectors and write to a MSP processing parameter file

```
DORIS_proc raw/p20050112.slc.par /mnt/scsi/DATA/DORIS/vor/DOR_VOR_AXVF-P20050303_101000_
20050111_215528_20050113_002328
```

Mode 3: Estimate Doppler centroid ([log](#))

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_3.log ASAR_all_list 3
```

SAR takes advantage of the Doppler history of the radar echoes generated by the forward motion of the satellite to synthesize a large antenna. The Doppler effect in combination with the Earth's rotation introduce a frequency shift in the received data, also known as the Doppler shift [*Hanssen*, 2001]. Returns from features in the area ahead of the platform will have upshifted (higher) frequencies, whereas returns from features behind the platform will have downshifted (lower) frequencies. The Doppler shift has to be accounted for in order to combine the information of two scenes into an interferogram. The interferometric correlation is optimized if the scene is observed from the same aspect angle. For a set of scenes this is generally satisfied by processing to a common Doppler centroid.

Zero-Doppler denotes the direction in which the Doppler frequency shift is equal to zero. Returns from the centerline of the beamwidth will show no frequency shift. This direction is perpendicular to the velocity vector of the spacecraft. The actual antenna of the ASAR instrument is aimed almost exactly towards the zero-Doppler line. The Doppler centroid (frequency) is the center frequency of the recorded azimuth spectrum [*Fiedler*, 2005] and is separated from the zero-Doppler line by the squint-angle. The Doppler centroid should be as small as possible over the entire scene. Otherwise repeat-pass interferometry can not be carried through because of

errors related to azimuth compression, range migration and calibration caused by a too large Doppler centroid. If the Doppler centroid reaches a value above 1 pulse repetition frequency (PRF) the determination of the Doppler ambiguity becomes a necessary step. However, in modern systems such as ERS and Envisat yaw-steering is employed to maintain the Doppler centroid of the data within $\frac{1}{2}$ of the PRF of the SAR making the Doppler ambiguity determination an optional step ([GAMMA MSP documentation](#)) for Envisat data. After Wegmüller [1997] an estimate of the Doppler centroid is then unambiguous because the Nyquist criteria for sampling of band-limited signals is satisfied.

Estimation of the fractional part of the Doppler centroid may be obtained by incoherent summation of azimuth spectra [Li, 1985] or cross correlation [Madson, 1989]. Since ASAR's yaw-steering avoids large squint angles, it is not necessary to use the cross-correlation algorithms which determine the Doppler centroid across the swath as a function of range. It is sufficient to average the azimuth spectrum.

The following command is run in mode 3 for the entire list of scenes: **azsp_IQ**

- Azimuth [spectrum](#) and Doppler centroid estimation for IQ raw SAR data (file format: [text](#))

```
azsp_IQ raw/ASAR_20050112_IS2_VV.par raw/p20050112.slc.par raw/20050112.raw raw/20050112.azsp
```

Calculate the median Doppler centroid (optional)

In order to further improve correlation all available scenes to the same median Doppler centroid. MATLAB can read the output of the following UNIX `grep` command if saved appropriately to calculate the median etc.

```
grep doppler_polynomial raw/*par >doppler
```

```
-5.73889e+01
-7.20603e+01
-1.58726e+02
-1.81143e+02
-1.61043e+02
-1.40870e+02
-1.53005e+02
-1.50379e+02
-1.56168e+02
-1.43340e+02
-1.52255e+02
-1.51796e+02
-1.56925e+02
-1.26771e+02
-1.21063e+02
-1.56638e+02
-2.09223e+02
-1.10465e+02
-1.12540e+02
-1.12393e+02
-1.29532e+02
```

```
matlab&
mean -138.7488
median -150.3790
range 151.8341
```


Mode 4: Set a value in the processing parameter files for a keyword: value pair (optional) (log)

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_4.log ASAR_all_list 4
doppler_polynomial "-150.3790 0. 0. 0."
```

The following command is run in mode 4 for the entire list of scenes: **set_value**

- Replace individual Doppler polynomials by median value

```
set_value raw/p20050112.slc.par raw/p20050112.slc.par doppler_polynomial "-150.3790 0. 0. 0."
keyword found: doppler_polynomial      old value: -1.21063e+02  0.00000e+00  0.00000e+00
0.00000e+00  Hz          Hz/m          Hz/m^2          Hz/m^3  new value: -150.3790 0. 0. 0.
```

Mode 5: Generate processing list for use by ASAR_proc_all (log)

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_5.log ASAR_all_list 5
```

Output parameter: [ASAR_all_list](#)

MSP-3 ASAR Processing from RAW Data

In order to prepare for further processing the multi-look ratio has to be determined. One of the results from the [ASAR_proc_all](#) run are the multi-look images, created from the SLC.

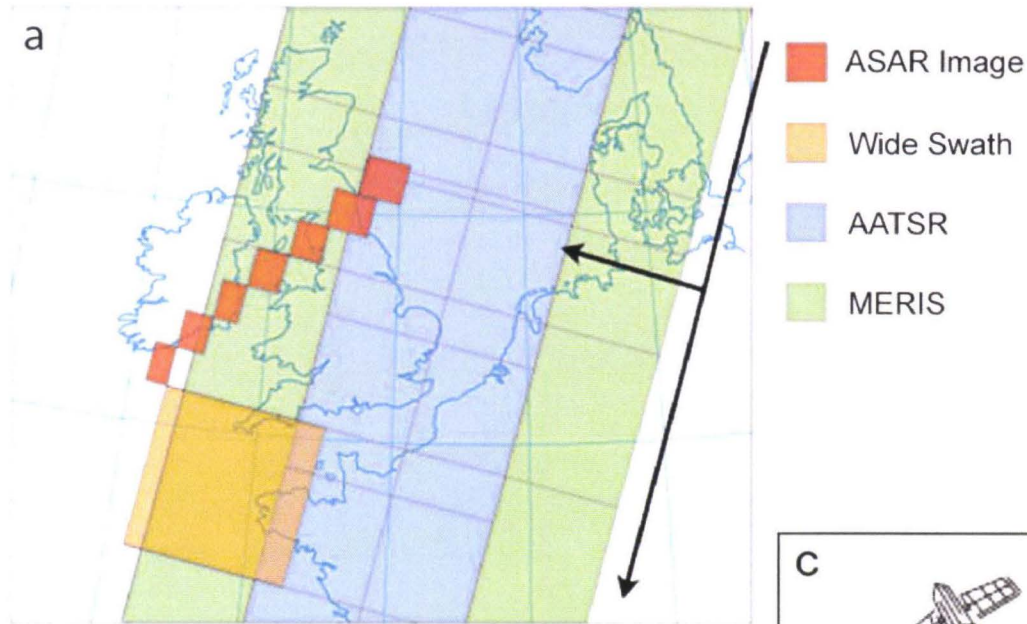
MSP-3.1 Detect Looks

The ASAR instrument aboard the Envisat satellite can maneuver in seven different image swaths when operating as a stripmap SAR (Figure 3a: [Envisat ground coverage](#) PDF 650 KB). The beam incidence angle increases with increasing image swath (IS1 - IS7), accordingly the swath width generally decreases. The slant-range pixel spacing or SRPS (slant-range: satellite's line-of-sight) depends only on the bandwidth of the sensor and is for Envisat approximately 7.8 m. Slant-range can be converted to ground-range by simply applying trigonometry, using the maximum and minimum beam incidence angles for the appropriate image swath.

The azimuth pixel spacing (APS) solely depends on the velocity of the satellite and is 4.1 m for IS2. The beam incidence angle ranges for IS2 from 19.2° to 26.7°, calculating a ground-range pixel spacing (GRPS) that ranges from 23.7 m for the lowest possible incidence angle to 17.4 m for the highest possible incidence angle (Figure 3b: [Table of ASAR image swaths & parameters](#)).

Imaging SLC data results in a distorted picture (Figure 4a: [SLC](#) PDF 10.1 MB) related to the slant-range projection. In order to show the radar data in a physically logical way, the SLC data needs to be projected to ground-range, i.e., it has to be multi-looked. The appropriate multi-look ratio (MLR) can be determined as follows:

For IS2 a GRPS of 23.7 – 17.4 m was calculated using a constant azimuth pixel spacing of 4.1 m (Figure 3c: [Slant-range to ground-range conversion](#)). It can be shown that the ratio between GRPS and APS ranges from 5.78 to 4.24. The average integer is 5, clearly the only possible integer in the middle of the spectrum between 5.78 and 4.24. The ratio therefore is 1:5, meaning the azimuth samples will be averaged by a factor of five while the range sample size remains the same.



b

Image Swath	Swath Width (km)	Ground position from nadir (km)	Incidence Angle Range ^o
IS1	105	187—292	15.0—22.9
IS2	105	242—347	19.2—26.7
IS3	82	337—419	26.0—31.4
IS4	88	412—500	31.0—36.3
IS5	64	490—555	35.8—39.4
IS6	70	550—620	39.1—42.8
IS7	56	615—671	42.5—45.2

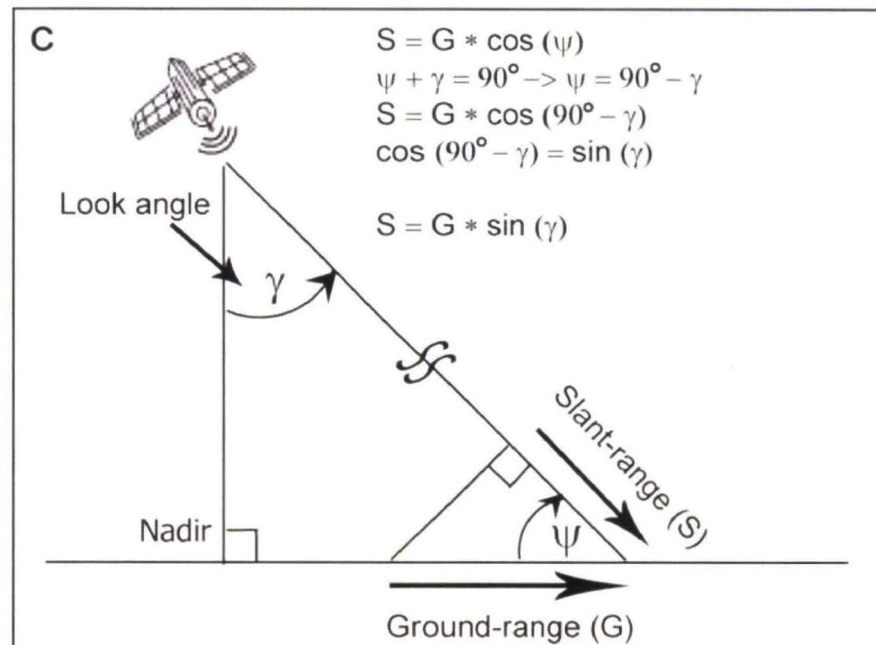


Figure 3: (a) Ground coverage of Envisat's ASAR image mode swaths 7 - 1 (from left to right) descending track, (b) Envisat's image swaths (IS) and corresponding parameters, red box refers to IS of dataset used in processing example (a, b: www.eurimage.com), (c) Slant-range (S) to ground-range (G) conversion (after ESA @ envisat.esa.int).

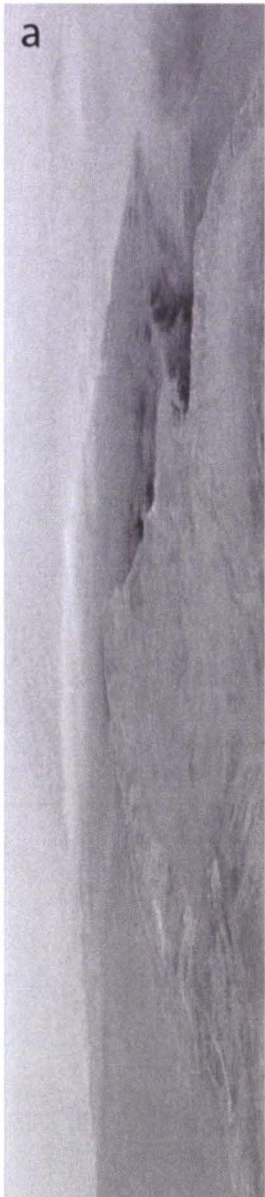


Figure 4: (a) Single-look raster image of the 20030212 single-look complex image (SLC) (MLR 1:1), (b) Multi-look intensity image (MLI) of the 20030212 SLC (MLR 1:5). Envisat ASAR IS2 track 429 frame 3213 (arrows indicate flight and look direction).

IS2 APS = 4.1 m SRPS = 7.8 m	Minimum value	Maximum value
γ	19.2°	26.7°
GRPS = SRPS/sin γ	23.7 m	17.4 m
MLR = GRPS/APS	5.78	4.24

For IS6 the multi-look ratio turns out to be 1:3.

IS6 APS = 3.9 m SRPS = 7.8 m	Minimum value	Maximum value
γ	39.1°	42.8°
GRPS = SRPS/sin γ	12.37 m	11.48 m
MLR = GRPS/APS	3.17	2.94

Multi-lookung reduces the number of azimuth samples by increasing the APS. For example, IS2 data originally has an APS of 4.1 m. Multi-lookung it 1:5 increases the APS by a factor of five to 20.5 m. This value is similar to the average GRPS which ranges from 17.4 m for the maximum incidence angle to 23.7 m for the lowest. Five pixels are now joined in one, increasing the signal-to-noise ratio (SNR) because more data (signal) is available for a specific area.

[ASAR_proc_all](#) is the final script to run in the MSP processing. Depending on the area covered in the scene there are two ways to run this script.

Firstly, these directories need to be generated in the processing directory (2429_3213/):

```
mkdir tmp
mkdir slc
mkdir mli_1_5
```

MSP-3.2 ASAR_proc_all

Scenes that are mostly covered by land close to the center of the SAR image will have no problem running this script. Scenes in which the center is largely surrounded by water however, are more problematic. One of the commands that is run in this script (**autof**) tries to auto focus over this area of water (no real data values over water) and does not determine a high enough SNR, hence it crashes (ERROR: autofocus rejected, correlation SNR below threshold: 10.00). Therefore the script [ASAR_proc_all_no_autof](#) has the two **autof** commands commented out. The main purpose of the script is to perform range and azimuth compression to recover full resolution. To follow is a sequence of commands run by the script ASAR_proc_all that generates the Single-Look Complex image (SLC) and associated parameter files.

```
ASAR_proc_all_no_autof ASAR_all_list raw tmp slc mli_1_5 1 5 0
```

The script runs the following commands for the entire list of scenes:

set_value, set_value, pre_rc, az_proc_dop2d, multi_SLC, raspwr, par_MSP, par_MSP

- Uses information in *ASAR_all_list* to update Doppler-polynomial (round median Doppler to first decimal)

```
set_value raw/p20050112.slc.par raw/p20050112.slc.par doppler_polynomial "-150.3790 0. 0. 0."
keyword found: doppler_polynomial    old value: -1.21063e+02  0.00000e+00  0.00000e+00
0.00000e+00  Hz      Hz/m      Hz/m^2      Hz/m^3  new value: -150.3790 0. 0. 0.
```

- Set the azimuth bandwidth fraction in the processing parameter file

```
set_value raw/p20030212.slc.par raw/p20030212.slc.par azimuth_bandwidth_fraction 0.8
keyword found: azimuth_bandwidth_fraction    old value: 0.8000  new value: 0.8
```

- SAR data prefilter, decimation and range compression for complex IQ data

A matched filter is applied to the data in range resulting in range compression. The option of decimation in azimuth (prefiltering) exists. Output file format: [rc_data](#)

```
pre_rc raw/ASAR_20030212_IS2_VV.par raw/p20030212.slc.par raw/20030212.raw tmp/20030212.rc
```

- SAR range-Doppler azimuth compression with along-track Doppler centroid update

This command reads the sensor and processing parameter files and the SAR range-compressed data. It performs azimuth compression using the range-Doppler algorithm. Radiometric compensation and calibration is applied. Corrections are performed for slant-range, antenna pattern and receiver gain. **Az_proc** creates the SLC which is visualized in [Figure 4](#). The two images appear “flipped” horizontally according to the viewing geometry of the radar on the satellite’s descending path (~9° east of north). The reference frame is also referred to as the range-Doppler coordinate system (RDC) describing the SAR geometry. To transform RDC to an Equal-Angle projection (EQA) in latitude/longitude the images would have to be “flipped” horizontally and “rotated” so north is up (see Geocoding chapter for more details). Figure 4a shows the SLC image with a look ratio of 1:1 (single-look), clearly indicating a “stretched” view. Multi-look adjusts the ground range pixel spacing to create “square” pixels (see section MSP-3.1). The raster image of the multi-look intensity image is shown in Figure 4b.

```
az_proc_dop2d raw/ASAR_20030212_IS2_VV.par raw/p20030212.slc.par tmp/20030212.rc
slc/20030212.slc 6144 0 -30 0 2.12
```

- Calculate a Multi-Look Intensity image ([MLI](#)) and [MLI_PROC_par](#) from a SLC using the previously determined MLR 1:5

```
multi_SLC raw/p20030212.slc.par mli_1_5/p20030212.mli.par slc/20030212.slc mli_1_5/20030212.mli 1 5
```

- Generate raster image of the multi-look intensity image (Figure 4b: [MLI](#))

```
raspwr mli_1_5/20030212.mli 5158
```

- Create [ISP image parameter](#) file from MSP processing parameter and sensor files

```
par_MSP raw/ASAR_20030212_IS2_VV.par raw/p20030212.slc.par slc/20030212.slc.par
```

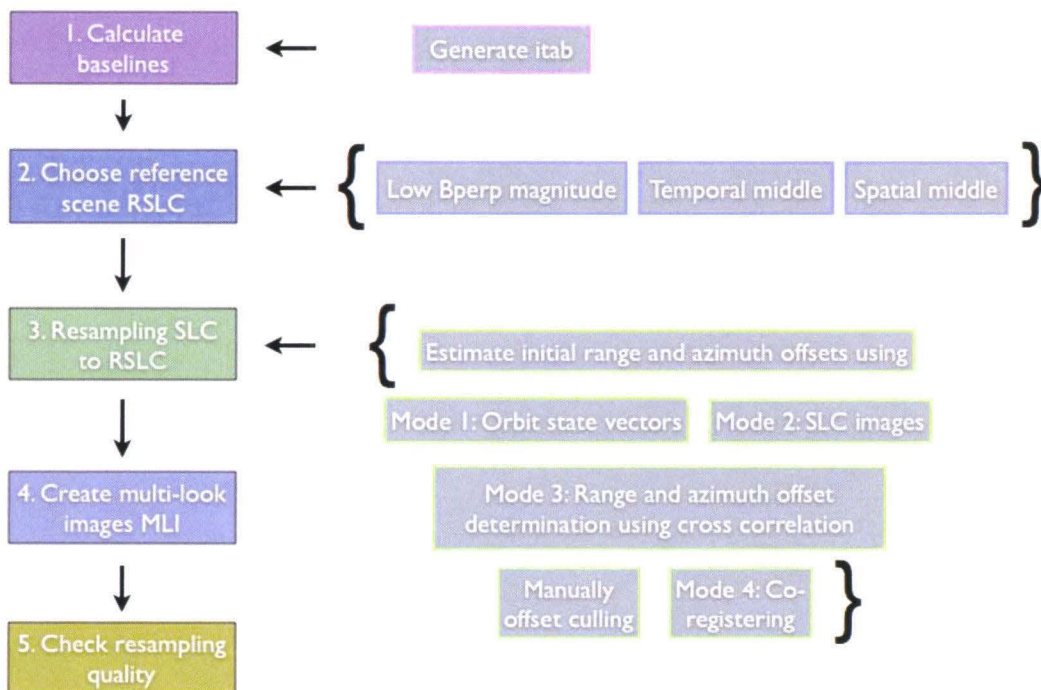
- Create [MLI ISP image parameter](#) file from MSP processing parameter and sensor files

```
par_MSP raw/ASAR_20030212_IS2_VV.par mli_1_5/p20030212.mli.par mli_1_5/20030212.mli.par
```

ISP – [Interferometric SAR processor](#) – Resampling

After generating of the single-look complex (SLC) images the resampling process can be started. One scene from the available data set is chosen as the reference scene. All other scenes need to be co-registered to the geometry of this reference scene. This permits the creation of interferograms at a later stage (ISP/DIFF).

The following flow chart describes the basic processing flow where each step is going to be illuminated in greater detail throughout this chapter. Processing scripts consist of a series of different commands and sometimes can be run in various modes, utilizing a different sequence of commands. Text displayed in the font `courier new` refers to either the content of a file or the script/command that is run and their standard output. The input arguments are separated by spaces. For an explanation of the input/output parameters please refer to **Appendix E** (index of commands/scripts and their syntax). A command history can be found in the [README](#) file and in the example of the MSP application in **Appendix H**.



In order to prepare for further processing, a list of all the SLC and the corresponding parameter files (two column list), referred to as the [SLC tab](#), can be created as follows in the processing directory (2429_3213/):

```

/bin/ls -l slc/*.slc > q1
/bin/ls -l slc/*.slc.par > q2
paste q1 q2 > SLC_tab
  
```

ISP-1 Calculate Baselines and Generate Itab File

In order to choose an appropriate reference scene, the baselines need to be calculated. The baseline refers to the distance between the two satellites at the time of the acquisitions and can be decomposed into a

perpendicular and a parallel baseline component. We calculate the perpendicular baselines from every scene with respect to all the other scenes to determine which scene is most appropriate as a reference. The GAMMA command `base_calc` calculates the baselines; `base_calc` also creates the `itab` file, a control file for further processing (example `itab`). The `itab` is an interferometric table that associates interferogram stack record numbers with pairs of SLC stack records (list from `SLC_tab`), enabling selective processing by “turning on or off” individual pairs. The first column in the `itab` corresponds to the master scene whereas the second column denotes the slave. The number refers to the position in the `SLC_tab`. The third column in the `itab` is the pair or record number. The fourth column is used to indicate if that specific pair is to be processed (1) or not (0). The example below shows excerpts from an `itab` (left) and a `SLC_tab` (right).

1	2	1	1		
1	5	2	1		
1	8	3	1	<code>slc/20030212.slc</code>	<code>slc/20030212.slc.par</code>
1	9	4	1	<code>slc/20030702.slc</code>	<code>slc/20030702.slc.par</code>
1	12	5	1	<code>slc/20030910.slc</code>	<code>slc/20030910.slc.par</code>
1	14	6	1	<code>slc/20031015.slc</code>	<code>slc/20031015.slc.par</code>
1	15	7	1	<code>slc/20031119.slc</code>	<code>slc/20031119.slc.par</code>
1	20	8	1	<code>slc/20031224.slc</code>	<code>slc/20031224.slc.par</code>
1	21	9	1	<code>slc/20040128.slc</code>	<code>slc/20040128.slc.par</code>
2	3	10	1	<code>slc/20040303.slc</code>	<code>slc/20040303.slc.par</code>
2	4	11	1	<code>slc/20040512.slc</code>	<code>slc/20040512.slc.par</code>

For example, the first pair denoted in the `itab` (1 2 1 1) has 20030212 as master scene (first entry in `SLC_tab`) and 20030702 as slave (second entry in `SLC_tab`). The pair (1 8 3 1) can be identified as the third record because of the number 3 in the third column. This pair has the same master scene but the eight entry from the `SLC_tab` as slave (20040303). All `itab` pairs shown here are “turned on” (1) and will be considered in the processing.

The number of pairs is limited in the `itab` by applying baseline and time span restrictions. The `itab` can either be a “single-reference” list, where the master (reference) scene in all pairs is the same or it can be an “all-pair” list allowing different master scenes in all interferometric pairs.

The standard output of every run is saved by piping it to a file using the operator >

```
base_calc SLC_tab slc/20030212.slc.par 20030212.bperp.gr 20030212.bperp itab 0 >1
base_calc SLC_tab slc/20030702.slc.par 20030702.bperp.gr 20030702.bperp itab 0 >2
```

ISP-2 Reference Scene: Selection-criteria and Algorithms

A “good” reference scene has to meet a few criteria. Its perpendicular baseline magnitude with respect to the other scenes should be low (< 200 m), the acquisition date should resemble the approximate temporal middle of all the available scenes and the center lat/long should represent a mean spatial center (see below).

The standard output of the command `base_calc` saved into separate files can be searched for the lowest perpendicular baseline magnitude. Perpendicular baselines can range from a few meters to about 1000 m and can also have negative values (relative positions of satellites).

The “average bperp magnitude” determines the average of the absolute values of the perpendicular baselines, whereas the “average bperp” averages the actual values of the perpendicular baselines which can be positive or negative. The average of the absolute values is used to establish a reference scene.

```
grep average 1
average bperp magnitude (m): 770.556
average bperp (m): 755.511
```

```
grep average 2
average bperp magnitude (m): 401.879
average bperp (m): 222.413
```

It is also important to find a reference scene that is temporally in the center of the time span of the set of available scenes. This is done by manually evaluating the acquisition dates.

Another important feature to take into account is the spatial middle of the stack. The satellite usually acquires over the same area, however, for some scenes the center lat/long can be very different. By choosing a scene whose area coverage is close to the areas covered by most of the other scenes in the stack, the area processed is maximized. The MATLAB program [scenelocation.m](#) (written by Matt Patrick) can be utilized for this purpose.

A few points need to be checked off before the program `scenelocation.m` can be run correctly:

- Copy the script `scenelocation.m` (see **Appendix F** for location) to the `raw` directory
- Adjust lines 3 and 29 in the script according to the current path of the `raw` directory
- Check the line number of the center latitude and center longitude in the `p*.slc.par` and change line 18 appropriately
- Choose a scene for the mid date (line 32), e.g., lowest Bperp, or temporal middle of the stack

Figures 5a and b ([MATLAB graphs](#) PDF 6.6 MB) show the MATLAB figure for the list of scenes used in this example. As you can see, the selected scene 20041208 is approximately in the middle in latitude (upper plot) and in longitude (lower plot). There are scenes that might be a little bit better suited geographically, but their Bperp magnitudes are larger.

ISP-3 Resample Set of SLC to a Common Reference SLC

The GAMMA script [SLC_resamp_all](#) is modified to accommodate a few user specific features. Increased number of estimated offsets and a manual offset culling mechanism are implemented. The chosen reference scene is `20041208.slc`. Before the resampling process can be started, the reference scene must be copied into the new directory `2429_3213/rslc`.

```
mkdir rslc
```

ISP-3.1 Copy Reference SLC to RSLC

In this case the entire SLC area is going to be processed. A smaller area might be chosen if processing times have to be kept low.

```
SLC_copy slc/20041208.slc slc/20041208.slc.par rslc/20041208.rslc rslc/20041208.rslc.par 1
```

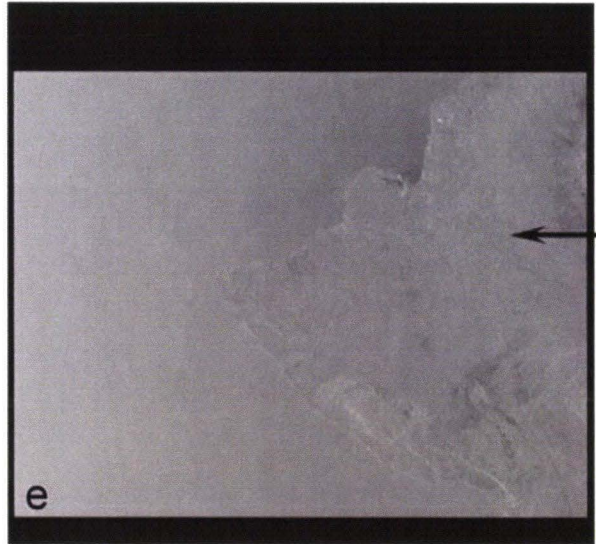
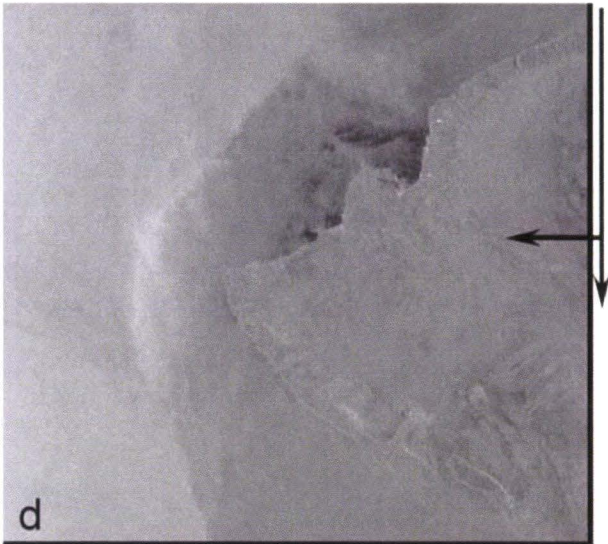
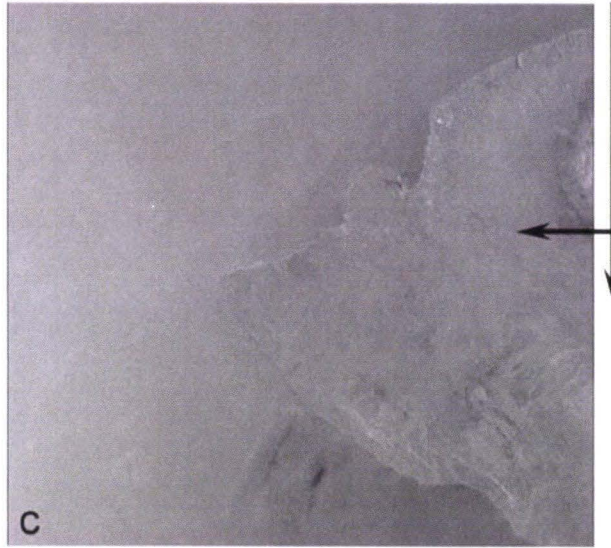
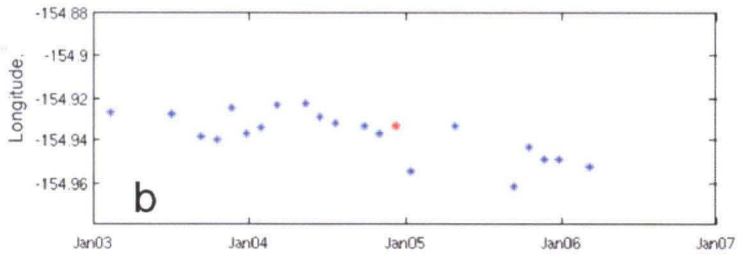
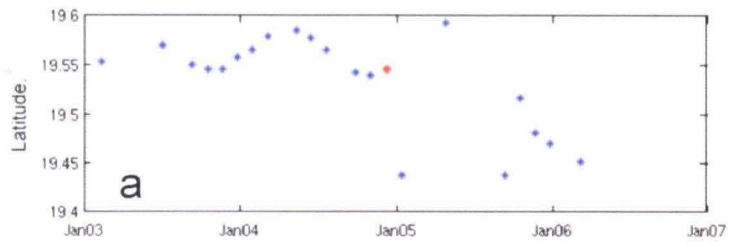



Figure 5: (a), (b) MATLAB figure generated by scenelocation.m, red star indicates reference scene 20041208. (a) Center latitude and (b) center longitude values for all 21 scenes, (c) Multi-look intensity (MLI) image of reference scene 20041208, (d) MLI image of 20030212, notice black stripes around the edges as remnants of the co-registration process to the reference scene, (e) average MLI image, note the small area representing the average covered area and the smoothness as a result of a high SNR.

In the next step multi-look images are generated. Again, the previously estimated multi-look ratio (MLR) is used.

```
mkdir rml_i_1_5
```

ISP-3.2 Calculate a Multi-Look Intensity (MLI) Image from the Reference RSLC

```
multi_look rslc/20041208.rslc rslc/20041208.rslc.par rml_i_1_5/20041208.rml_i_1_5/20041208.rml_i_1_5.par 1 5
```

ISP-3.3 Generate Raster Image of RMLI (Figure 5c: [RMLI](#))

```
raspwr rml_i_1_5/20041208.rml_i_1_5 5174
```

ISP-3.4 Resample all SLC to Reference RSLC ([RSLC_tab](#))

The modified script [SLC_resamp_all_cs](#) should be run in the following modes to resample a set of SLC:

- 0: Create offset parameter files
- 1: Estimate initial range and azimuth offsets using orbit state vectors
- 2: Measure initial range and azimuth offsets using SLC images
- 3: Estimate range and azimuth offset models using correlation of image intensities
- 4: Resample SLC images using offset models

Mode 0: Create offset parameter files (log)

```
SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 0 1 1 5
```

The script runs the following command in mode 0 for the entire list of SLC: **create_offset**

- Create offset parameter files for SLC – RSLC offsets

```
create_offset rslc/20041208.rslc.par slc/20030212.slc.par rslc/20041208_20030212.off <
create_offset.in
```

[create_offset.in](#) is created in SLC_resamp_all_cs mode 0 and includes the input parameters for create_offset.

Without this file the parameters would have to be specified manually on the command line for each scene.

Mode 1: Estimate initial range and azimuth offsets using orbit state vectors (log)

```
SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 1 1 1 5
```

The script runs the following command in mode 1 for the entire list of SLC: **init_offset_orbit**

- Initial SLC image offset estimation from orbit state vector data

```
init_offset_orbit rslc/20041208.rslc.par slc/20030212.slc.par rslc/20041208_20030212.off
```

It is recommended to save the standard output in order to compare the orbit-derived offset estimation to the SLC-derived offset estimation conducted in mode 2.

orbit-derived initial range and azimuth offsets:

```
init_offset_orbit rslc/20041208.rslc.par slc/20030212.slc.par rslc/20041208_20030212.off
initial_range_offset:          16
initial_azimuth_offset:       221

init_offset_orbit rslc/20041208.rslc.par slc/20030702.slc.par rslc/20041208_20030702.off
initial_range_offset:          -9
initial_azimuth_offset:       674
```

Mode 2: Measure initial range and azimuth offsets using SLC images (log)

```
SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 2 1 1 5
```

The script runs the following command in mode 2 for the entire list of SLC: **init_offset**

- Determine initial offset between SLC images using correlation of image intensity

In order to avoid ambiguity problems and still achieve an accurate offset estimate, **init_offset** can first be run with multi-looking followed by a second run at single-look resolution. The initial estimates from the first run are used as an initial guess in the second run. In the case of using precise orbits as illustrated in this example, **init_offset_orbit** is run prior to **init_offset**, providing the initial guess and accounting for large registration offsets.

```
init_offset rslc/20041208.rslc slc/20030212.slc rslc/20041208.rslc.par slc/20030212.slc.par
rslc/20041208_20030212.off 1 5 2587 11908
```

It is recommended to save the standard output in order to compare the SLC-derived offset estimation to the orbit-derived offset estimation (see mode 1).

```
SLC-derived initial range and azimuth offsets:
-8 *** (-9) means that -8 is SLC-derived, whereas -9 was orbit-derived

init_offset rslc/20041208.rslc slc/20030212.slc rslc/20041208.rslc.par slc/20030212.slc.par
rslc/20041208_20030212.off 1 5 2587 11908
initial_range_offset:          16
initial_azimuth_offset:       221

init_offset rslc/20041208.rslc slc/20030702.slc rslc/20041208.rslc.par slc/20030702.slc.par
rslc/20041208_20030702.off 1 5 2587 11908
initial_range_offset:          -8*** (-9)
initial_azimuth_offset:       674
```

Mode 3: Estimate range and azimuth offset models using correlation of image intensities (log)

```
SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 3 1 1 5
```

The script runs the following command in mode 3 for the entire list of SLC: **offset_pwr**

- Measure range and azimuth registration offset field between reference SLC and RSLC

Cross correlation optimization of detected SLC data is used rather than fringe visibility algorithm (**offset_SLC**)

```
offset_pwr rslc/20041208.rslc slc/20030212.slc rslc/20041208.rslc.par slc/20030212.slc.par
rslc/20041208_20030212.off rslc/20041208_20030212.off rslc/20041208_20030212.snr - - - 2 100
100
```

Note: The script **SLC_resamp_all_cs** does not run the last **offset_pwr** command completely; it needs to be run manually after mode 3 finishes (the command is printed in std out on the screen).

Manual culling of offsets

The script `SLC_resamp_all` was changed in order to interactively cull the estimated offsets. The range and azimuth offset polynomials are computed from the offsets estimated by `offset_pwr`. The program creates polynomial models of range and azimuth offsets using linear least-squares. Particular offsets can be identified as erroneous depending on an indicated SNR (in this case 7.0) and on the deviation of the offset measurement to the current model value. Culling the offsets manually allows the user to observe how many offsets are “thrown out” by setting lower azimuth and range standard deviation thresholds and how that affects the offset-polynomial (here changed to a 4th order polynomial). The commands for all the scenes have to be created manually. In general, a co-registration accuracy better than 0.2 pixels allows to limit the loss of correlation by registration errors to < 10%. It is recommended to save the solution of the offset fit (in the processing [README](#)), i.e., the number of offset estimates accepted out of the total number of samples (calculated by `offset_pwr` in mode 3).

```
offset_fit rslc/20041208_20030212.off rslc/20041208_20030212.sn timer/20041208_20030212.off
rslc/20041208_20030212.coffs - 6.5 4 1
solution: 27569 offset estimates accepted out of 250000 samples
range offset poly. coeff.:      11.42545  1.74877e-03  5.73562e-05  -1.05131e-08
azimuth offset poly. coeff.:    221.03254  -2.32978e-04  1.81379e-06  -8.40027e-11
model fit std. dev. (samples) range:  0.0994  azimuth:  0.0983
```

```
offset_fit rslc/20041208_20030702.off rslc/20041208_20030702.sn timer/20041208_20030702.off
rslc/20041208_20030702.coffs - 6.5 4 1
solution: 42089 offset estimates accepted out of 250000 samples
range offset poly. coeff.:      -8.71272  2.12797e-04  -1.91231e-05  -7.35684e-11
azimuth offset poly. coeff.:    672.52513  6.60038e-04  -1.68263e-06  -4.94542e-09
model fit std. dev. (samples) range:  0.0748  azimuth:  0.0997
```

Method of interactively culling the offsets:

(1) Run the command `offset_fit` from above:

```
solution: 51845 offset estimates accepted out of 250000 samples
range offset poly. coeff.: 11.41332 1.74030e-03 7.09690e-05 -1.30885e-08
azimuth offset poly. coeff.: 221.60961 -3.28656e-04 -2.07447e-04 4.72914e-08
model fit std. dev. (samples) range : 1.4857 azimuth: 3.2809
```

The model fit standard deviation (std dev) of the range and azimuth samples should be < 0.1 pixel in our case to absolutely minimize the errors due to misregistration. Usually the azimuth std dev is higher because of the multi-looking in this direction (larger pixel spacing).

(2) Enter the minimum SNR threshold: 6.5

(3) Enter the range and azimuth error thresholds: 10 10

Initial thresholds should be high in order to eliminate outliers.

(4) std out:

```
*** improved least-squares polynomial coefficients 1 ***
solution: 51473 offset estimates accepted out of 250000 samples
range offset poly. coeff.: 11.71386 1.67188e-03 3.97096e-05 -5.85950e-09
azimuth offset poly. coeff.: 220.94872 -2.17446e-04 8.21322e-06 -1.36307e-09
model fit std. dev. (samples) range: 0.2090 azimuth: 0.8535
```

Without outliers the std dev's came down from 1.5 to 0.2 pixel in range and from 3.3 to 0.9 pixel in azimuth.

(5) Iterate and further refine the offset fit? (1 = yes, 0 = no): 1

(6) Enter the minimum SNR threshold: 6.5

(7) Enter the range and azimuth error thresholds: 5 5

Slowly decrease the thresholds.

(8) std out:

```
solution: 51229 offset estimates accepted out of 250000 samples
range offset poly. coeff.: 11.72166 1.67031e-03 3.89725e-05 -5.70626e-09
azimuth offset poly. coeff.: 220.93766 -2.13889e-04 8.41821e-06 -1.46604e-09
model fit std. dev. (samples) range: 0.1754 azimuth: 0.7181
```

This decreased the std dev's only by a small amount, i.e., smaller thresholds can be entered.

(9) Repeat steps 6 and 7, slowly decreasing the thresholds, until the std dev's get close to 0.1 pixel.

Always watch the number of offsets decreasing and how the polynomials change. It is acceptable for the polynomials to only change by a couple of decimals, while the number of accepted offsets falls by thousands. Depending on the dataset the co-registered images will look fine even with few accepted offsets (~1000 out of 250000).

(10) std out:

```
enter the range and azimuth error thresholds: .3 .3
range, azimuth error thresholds: 0.3000 0.3000
SNR threshold: 6.5000
*** improved least-squares polynomial coefficients 1 ***
solution: 33285 offset estimates accepted out of 250000 samples
range offset poly. coeff.: 11.55531 1.71481e-03 4.90487e-05 -8.39746e-09
azimuth offset poly. coeff.: 220.98648 -2.20710e-04 4.05980e-06 -6.85199e-10
model fit std. dev. (samples) range: 0.1191 azimuth: 0.1258
```

Now decrease the thresholds in steps of 0.01 to be sure not to throw out the "good" offsets while approaching values just below 0.1 pixel. Stop iterating when both std dev's are below 0.1 pixel.

(11) std out:

```
solution: 15623 offset estimates accepted out of 250000 samples
range offset poly. coeff.: 11.24656 1.77427e-03 8.84671e-05 -1.65948e-08
azimuth offset poly. coeff.: 221.24406 -2.56562e-04 -7.46270e-05 1.62692e-08
model fit std. dev. (samples) range: 0.0998 azimuth: 0.0997
```

```
Iterate and further refine the offset fit? (1=yes, 0=no): 0
```

Summary:

250,000 samples	# Accepted offset estimates
Initial	51845
Final	15623

Polynomials coefficients	Range	Azimuth
Initial	11.41 1.74e-03 7.10e-05 -1.31e-08	221.61 -3.29e-04 -2.07e-04 4.73e-08
Final	11.25 1.77e-03 8.85e-05 -1.66e-08	221.24 -2.57e-04 -7.46e-05 1.62e-08

Model fit std dev	Range	Azimuth
Initial	1.4857	3.2809
Final	0.0998	0.0997

Mode 4: Resample SLC images using offset models (log)

```
SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 4 1 1 5
```

Note that the rflag is turned on (1) for measuring offsets to the resampled SLC to confirm the offset model

The script runs the following commands in mode 4 for the entire list of SLC:

SLC_interp, create_offset, offset_pwr, offset_fit, set_value, set_value, SLC_interp, offset_pwr, offset_fit

- SLC complex image resampling using 2-D SINC interpolation and the range/azimuth polynomials

```
SLC_interp slc/20030212.slc rslc/20041208.rslc.par slc/20030212.slc.par rslc/20041208_20030212.off rslc/20030212.rslc rslc/20030212.rslc.par
```

- Create offset file for offset measurement between the reference SLC and the resampled SLC

```
create_offset rslc/20041208.rslc.par rslc/20030212.rslc.par rslc/20041208_20030212_2.off < create_offset.in
```

- Measure offsets between reference SLC and RSLC

```
offset_pwr rslc/20041208.rslc rslc/20030212.rslc rslc/20041208.rslc.par rslc/20030212.rslc.par rslc/20041208_20030212_2.off rslc/20041208_20030212_2.offsets rslc/20041208_20030212_2.snr - - - 2 50 50
```

- Estimate the offset polynomial

```
offset_fit rslc/20041208_20030212_2.offsets rslc/20041208_20030212_2.snr rslc/20041208_20030212_2.off rslc/20041208_20030212_2.coffs - - 3
```

Show below is the standard output printed on the screen after running this command.

```
*** improved least-squares polynomial coefficients 6 ***
solution: 448 offset estimates accepted out of 2500 samples
range offset poly. coeff.:          -0.00545  -7.99216e-06  2.36481e-06
azimuth offset poly. coeff.:        -0.06388   1.17525e-05  1.30676e-06
model fit std. dev. (samples) range:  0.1082  azimuth:  0.2766

total number of culling iterations: 6
final solution: 448 offset estimates accepted out of 2500 samples

final range offset poly. coeff.:          -0.00545  -7.99216e-06  2.36481e-06
final range offset poly. coeff. errors:  2.46161e-03  5.75577e-07  7.11318e-08

final azimuth offset poly. coeff.:        -0.06388   1.17525e-05  1.30676e-06
final azimuth offset poly. coeff. errors:  6.29377e-03  1.47162e-06  1.81868e-07

final model fit std. dev. (samples) range:  0.1082  azimuth:  0.2766
```

The just estimated offset polynomial is added to the previously estimated offset polynomial, performed by the script SLC_resamp_all_cs.

- The sum of the range offset polynomials is written in the offset parameter file

```
set_value rslc/20041208_20030212.off rslc/20041208_20030212.off "range_offset_polynomial" "
1.1420000e+01 1.74081e-03 5.97208e-05 -1.05130e-08"
*** update keyword:value based parameter files ***
*** Copyright 2004, Gamma Remote Sensing, v1.2 12-Feb-2004 clw ***

keyword found: range_offset_polynomial    old value: 11.42545  1.7488e-03  5.7356e-05  -
1.0513e-08  new value:  1.1420000e+01  1.74081e-03  5.97208e-05  -1.05130e-08
```

- The sum of the azimuth offset polynomials is written in the offset parameter

```
set_value rslc/20041208_20030212.off rslc/20041208_20030212.off "azimuth_offset_polynomial" "
2.2096866e+02 -2.21227e-04 3.12060e-06 -8.40030e-11"
*** update keyword:value based parameter files ***
*** Copyright 2004, Gamma Remote Sensing, v1.2 12-Feb-2004 clw ***

keyword found: azimuth_offset_polynomial  old value: 221.03254  -2.3298e-04  1.8138e-06  -
8.4003e-11  new value:  2.2096866e+02  -2.21227e-04  3.12060e-06  -8.40030e-11
```

- Resample the SLC again using the refined offset polynomial stored in the offset parameter file

```
SLC_interp slc/20030212.slc rslc/20041208.rslc.par slc/20030212.slc.par rslc/20041208_
20030212.off rslc/20030212.rslc rslc/20030212.rslc.par
```

- Measure offsets between refined RSLC and reference SLC

```
offset_pwr rslc/20041208.rslc rslc/20030212.rslc rslc/20041208.rslc.par rslc/20030212.rslc.par
rslc/20041208_20030212_2.off rslc/20041208_20030212_2.off rslc/20041208_20030212_2.snr - - - 2
50 50
```

- Fit polynomial to the offsets

```
offset_fit rslc/20041208_20030212_2.off rslc/20041208_20030212_2.snr rslc/20041208_20030212_
2.off rslc/20041208_20030212_2.coffs - - 3
```

```
solution: 444 offset estimates accepted out of 2500 samples
range offset poly. coeff.:          -0.00592  4.21391e-07  2.67473e-07
azimuth offset poly. coeff.:        -0.00636  -6.67370e-07  5.54602e-07
model fit std. dev. (samples) range: 0.1059  azimuth: 0.2719
```

```
total number of culling iterations: 6
final solution: 444 offset estimates accepted out of 2500 samples
```

```
final range offset poly. coeff.:          -0.00592  4.21391e-07  2.67473e-07
final range offset poly. coeff. errors:  2.41795e-03  5.66593e-07  6.99042e-08

final azimuth offset poly. coeff.:        -0.00636  -6.67370e-07  5.54602e-07
final azimuth offset poly. coeff. errors:  6.20595e-03  1.45423e-06  1.79418e-07

final model fit std. dev. (samples) range: 0.1059  azimuth: 0.2719
```

Compared to the earlier results these std dev's are smaller and therefore the polynomials are accepted as the

more precise offset polynomials. One iteration is usually sufficient to improve the models.

ISP-4 Create Multi-look Images for Individual RSLC and Calculate Average MLI

```
mk_mli_all RSLC_tab rml_i_1_5 1 5 1
```

The script runs the following commands for the entire list of RSLC:

multi_look, raspwr, ave_image, raspwr, disras

- Multi-look the co-registered RSLC using a 1:5 ratio

```
multi_look rslc/20030212.rslc rslc/20030212.rslc.par rml_i_1_5/20030212.rml_i_1_5/20030212.rml_i_1_5.par 1 5
```

- Create a raster image of the RMLI (Figure 5d: [RMLI](#))

```
raspwr rml_i_1_5/20030212.rml_i_1_5 5174 1 4763 1 1 0.9 0.35
```

- Calculate average of input 2D data files (RMLI)

```
ave_image rml_i_1_5/mli_list 5174 rml_i_1_5/rml_i_1_5.ave
```

- Generate raster of average image (Figure 5e: [Average RMLI](#))

```
raspwr rml_i_1_5/rml_i_1_5.ave 5174 1 0 1 1 .85
```

- Display of Sun raster images

```
disras rml_i_1_5/rml_i_1_5.ave.ras&
```

- Create parameter file for the average image

```
cp rml_i_1_5/20041208.rml_i_1_5.par rml_i_1_5/rml_i_1_5.ave.par
```

ISP-5 Check Quality of the Co-registered SLC

Display the co-registered RSLC (as MLI) and the reference SLC on top of each other (middle mouse button switches between images)

- Alternating display of 2 SUN raster or BMP format images

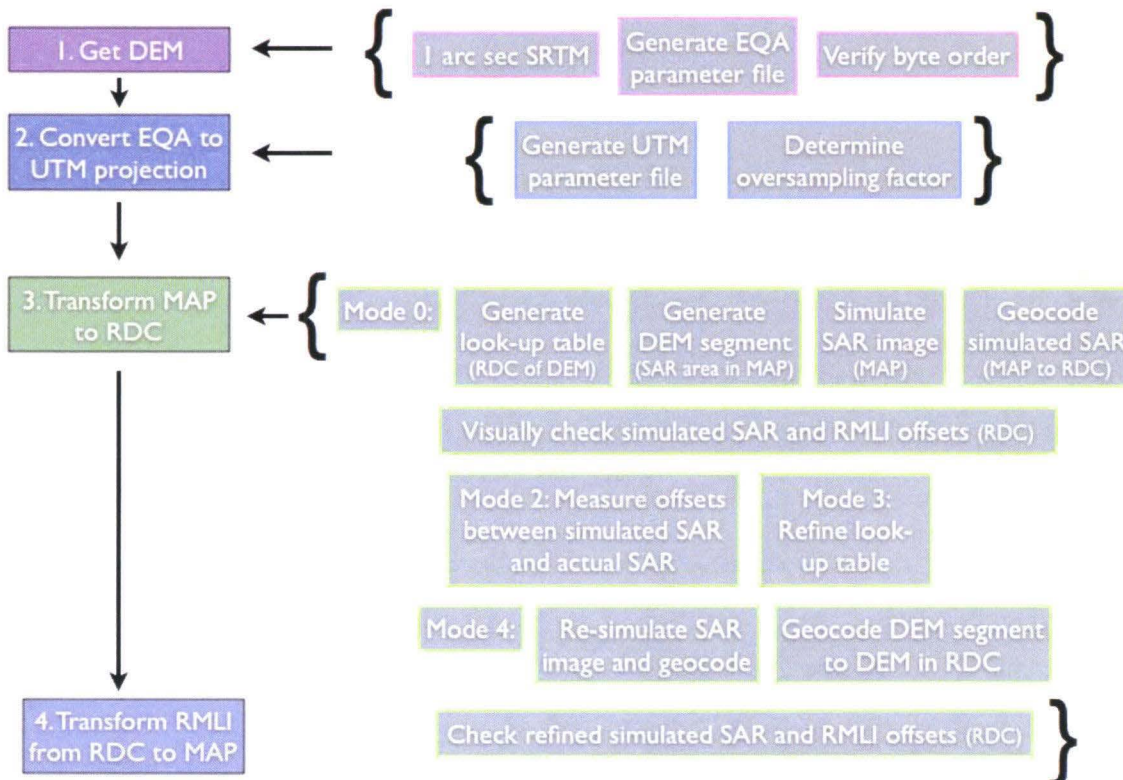
```
dis2ras rml_i_1_5/20030212.rml_i_1_5.ras rml_i_1_5/20041208.rml_i_1_5.ras
dis2ras rml_i_1_5/20030702.rml_i_1_5.ras rml_i_1_5/20041208.rml_i_1_5.ras
```

By using the middle mouse button and focusing on the zoom window one can switch between image 1 and image 2 and verify the co-registration. If no obvious differences across the entire scene exist, the resampling process is done. However, for certain datasets there are noticeable offsets of 1 pixel or above. In that case the script `SLC_resamp_lt_all` needs to be run, which uses a look-up table (map) of offsets instead of a polynomial. The scenes need to all have the same size (width x height) as the reference scene.

GEO – [SAR Geocoding and Image Registration](#)

The idea of interferometric processing is to gather information about the phase after combining two complex SAR images. The interferometric phase consist of signals from surface topography, displacement along the look-vector and noise terms such as atmospheric path delay and errors related to baselines. In order to solve for the deformation signal, the topographic phase contribution has to be simulated and subtracted from the interferogram. As part of the geocoding process described in this chapter, an existing digital elevation model (DEM) will be used to simulate the topographic phase. GAMMA's geocoding module is capable of converting between two different coordinate systems. DEMs in lat/long projection are in MAP coordinates, whereas SAR products such as SAR images and interferograms are in range-Doppler coordinates (RDC). To be able to subtract the topographic phase from the complex interferogram in SAR projection, the DEM needs to be converted to RDC. The result is a differential interferogram, consisting of only deformation, atmospheric distortion and baseline related errors. The term forward geocoding is synonymous for the transformation from MAP to RDC, whereas backward geocoding stands for RDC to MAP conversion.

The following flow chart describes the basic processing flow where each step is going to be illuminated in greater detail throughout this chapter. Processing scripts consist of a series of different commands and sometimes can be run in various modes, utilizing a different sequence of commands. Text displayed in the font `courier new` refers to either the content of a file (e.g., `ASAR_pre_list`) or the script/command that is run and their standard output. The input arguments are separated by spaces. For an explanation of the input/output parameters please refer to **Appendix E** (Index of commands/scripts and their syntax). A command history can be found in the [README](#) file and in the example of the MSP application in **Appendix H**.



GEO-1 Type and Location of Raw DEM Data

The DEM used in the following processing is a product of the Shuttle Radar Topography Mission (SRTM). It was downloaded as the finished data product, delineating and flattening water bodies, better defining coastlines, removing "spikes" and "wells", and filling small voids (steep slopes still may have voids). The data are available at one or three arc sec resolution. One degree of arc is subdivided into 60 arc minutes. One minute of arc consists of 60 seconds. One arc minute of longitude is equivalent to 1853 m at the equator. In order to calculate the resolution of 1 arc second of longitude at latitude 20, the distance needs to be multiplied by the cosine (lat) which results in an approximate resolution of 29.02 m (2.77e-04 decimal degree) in longitude ([USGS SRTM DEM info](#)).

[USGS seamless server](#) also provides a [meta](#) file, including information on identification, data quality, spatial data organization etc, a [header](#) file, giving the dimensions as well as the byte order, a [*.blw](#) file including the pixel spacing in lat and long as well as the upper left lat/long coordinates for the downloaded data. The SRTM DEMs are only available in equal angle projection (EQA) and need to be converted to UTM if desired.

GEO-2 Generation of DEM Parameter File

Parameters from the corresponding files mentioned above are used as the input on the command line.

```
create_dem_par SRTM_BI_1arc_eqa.dem_par
```

GEO-3 Byte Order Verification

The Byte order of the SRTM DEM is Intel (little-endian, most significant byte (MSB) last). Since GAMMA operations are run on a big-endian machine, the bytes need to be swapped to Motorola (big-endian, MSB first) to continue processing.

- Swap bytes for binary data

```
swap_bytes SRTM_BI_1arc_eqa.dem SRTM_BI_1arc_eqa_swp.dem 2
```

- Display DEM to check area coverage (Figure 6a: [DEM in EQA](#) PDF 3.4 MB)

Display DEM with DEM/MAP parameter file as shaded relief to see whether SAR area is fully covered

```
disdem_par SRTM_BI_1arc_eqa_swp.dem SRTM_BI_1arc_eqa.dem_par
```

GEO-4 Conversion from EQA to UTM

The SAR image has an approximate ground-range pixel spacing of 20 m. The EQA DEM posting is at about 30 m. It is recommended to "adjust" the UTM DEM spacing to the SAR pixel spacing, utilizing one of two different approaches. The first approach, also presented in the following example, chooses 15 m as the UTM DEM spacing, i.e., an approximate oversampling factor of two would be applied during the DEM transformation. The original EQA DEM dimensions are 5492 x 5520, hence the UTM DEM dimensions should be around 10,000 x 10,000 (during [UTM DEM parameter file](#) generation). The input parameter <post> in the DEM geocoding script [mk_geo_cs](#) is therefore going to be set to 15 m. The script then determines the interpolation factor by dividing the UTM spacing (found in the parameter file) by the input parameter <post>. In this case the interpolation factor for northing and easting is one. The alternative approach leaves the original EQA spacing of 30 m in the

generation process of the UTM DEM parameter file. The script `mk_geo_cs` then determines an interpolation of 2, dividing 30 m by the input posting of 15 m. Another important detail to check for is the sign of the latitude/longitude of the first sample as well as for the latitude/longitude posting. By displaying the DEM in GAMMA the correct sign can be determined. Latitude and longitude have to behave in a certain way (increase/decrease) depending on where you are located on the globe. For example, the Big Island of Hawaii is located in the northern hemisphere, therefore latitude has to increase towards north. The latitude of the upper left corner has to be a positive value and the latitude posting value is negative, accounting for increasing latitude with decreasing azimuth. The same concept applies to the longitude. The longitude value is negative, indicating that Hawaii is west of 0° Greenwich. The longitude posting value is positive, accounting for the increasing value with increasing range.

```
create_dem_par SRTM_BI_15m_utm.dem_par rqli_1_5/rqli_1_5.ave.par
```

- Transform EQA DEM to UTM

```
dem_trans SRTM_BI_1arc_eqa.dem_par SRTM_BI_1arc_eqa_swp.dem SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm.dem
```

- Display the UTM DEM (Figure 6b: [UTM DEM](#)) and the SAR image (Figure 5e: [Average RMLI](#)) to check area coverage

```
disdem_par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm.dem_par
gimp rqli_1_5/rqli_1_5.ave.ras
```

GEO-5 Transforming DEM MAP Coordinates to RDC Geometry

GAMMA's geocoding script `mk_geo` was slightly modified, as described below. It can be run in the following seven modes.

0: Generate initial lookup table, simulated SAR image, and DEM segment parameters

1: Generate initial lookup table and simulated SAR image using existing DEM segment to determine image bounds

2: Measure initial offset between simulated SAR image and actual SAR image

3: Perform refinement of lookup table by offset measurement with respect to the simulated SAR image

4: Update lookup table and produce terrain geocoded SAR image and DEM in SAR range-Doppler coordinates (RDC)

5: Ellipsoid geocoding of the SAR image without a DEM: generate new DEM segment parameters

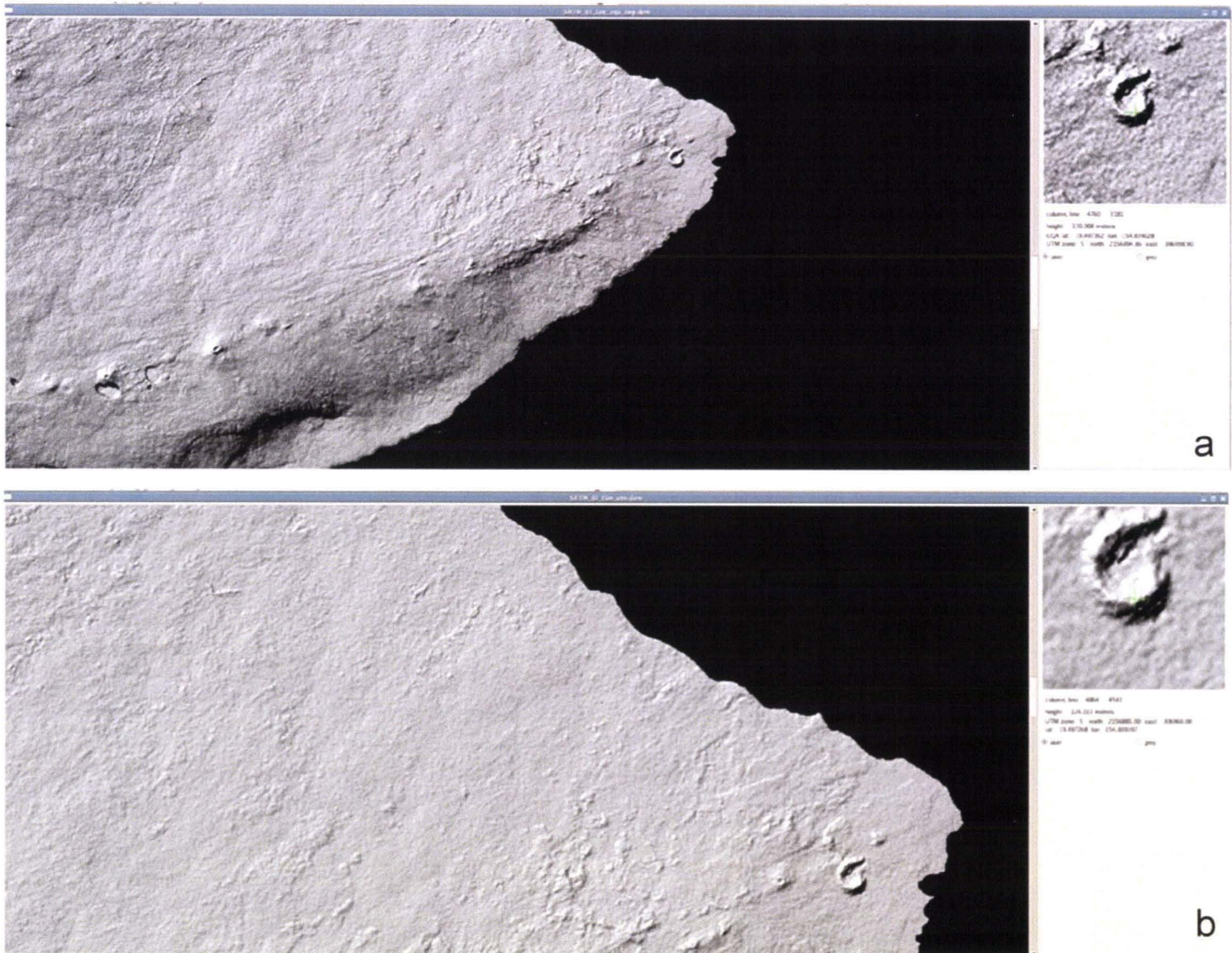
6: Ellipsoid geocoding of the SAR image without a DEM: use existing DEM segment parameters

Terrain geocoding of SAR images with lookup table refinement and transform DEM → SAR range-Doppler Coordinates (RDC). The following example will run `mk_geo_cs` in modes 0, 2, 3 and 4.

Mode 0: Generate initial lookup table, simulated SAR image, and DEM segment parameters ([log](#))

```
mk_geo_cs rqli_1_5/rqli_1_5.ave rqli_1_5/rqli_1_5.ave.par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI 15 0 2 2 2 8 1 5 6.5
- - - - 256 256
```

The script runs the following commands in mode 0: `gc_map`, `geocode`, `raspwr`, `dis2ras`



In the following processing a segment file will be generated. It represents the section the SAR image would cover on the SRTM DEM. In other words, the script will cut out the radar image section from the DEM. The first step is to remove any existing segment files.

```
rm -f SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par
```

- Derive the geocoding look-up table and simulate a SAR backscatter image (in DEM geometry)

```
gc_map rml1_1_5/rml1_1_5.ave.par - SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm_seg.dem_par SRTM_BI_15m_utm_seg.dem geo/BI_0.map_to_rdc 1 1 geo/BI_0.sim - - - - - 8 2 2
```

- Forward geocoding transformation using the look-up table

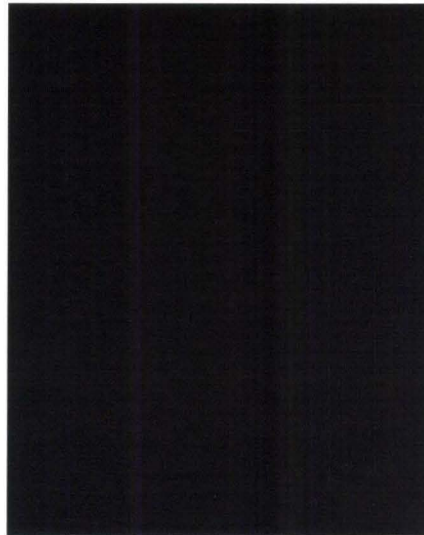
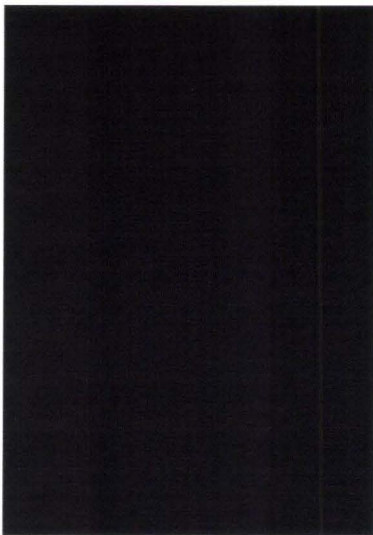
Transform the simulated SAR image (in MAP geometry) to RDC

```
geocode geo/BI_0.map_to_rdc geo/BI_0.sim 8334 geo/BI_0.sim_rdc 5174 4763 2 0 1 1 2 8
```

- Create raster image of the initial simulated SAR image in RDC (Figure 7a: [Initial simulated SAR image](#) PDF 540 KB)

```
raspwr geo/BI_0.sim_rdc 5174 1 0 1 1
```

The image appears to be “flipped” horizontally arising to the fact that the satellite is on a descending (north to south) track which is slightly rotated to the east by about 9° (see schematic below). The right-looking radar acquires data starting at proximal points on the ground after which more distal ground features are recorded. In the descending example, the SAR at first writes B, the closest features, into the data file, then A. While the satellite progresses along its track, ground features close to C are recorded before D. The imaged ground features (the SAR images) therefore seem flipped horizontally. On an ascending track, the SAR records ground features closest to D before it records C. Further along the flight the SAR writes A before it can “see” B. The ascending data therefore seem flipped vertically. Images in this geometry are in SAR or range-Doppler Coordinates (RDC) and will have the flight and look direction. Images in MAP geometry like EQA or UTM projection will show a north arrow.



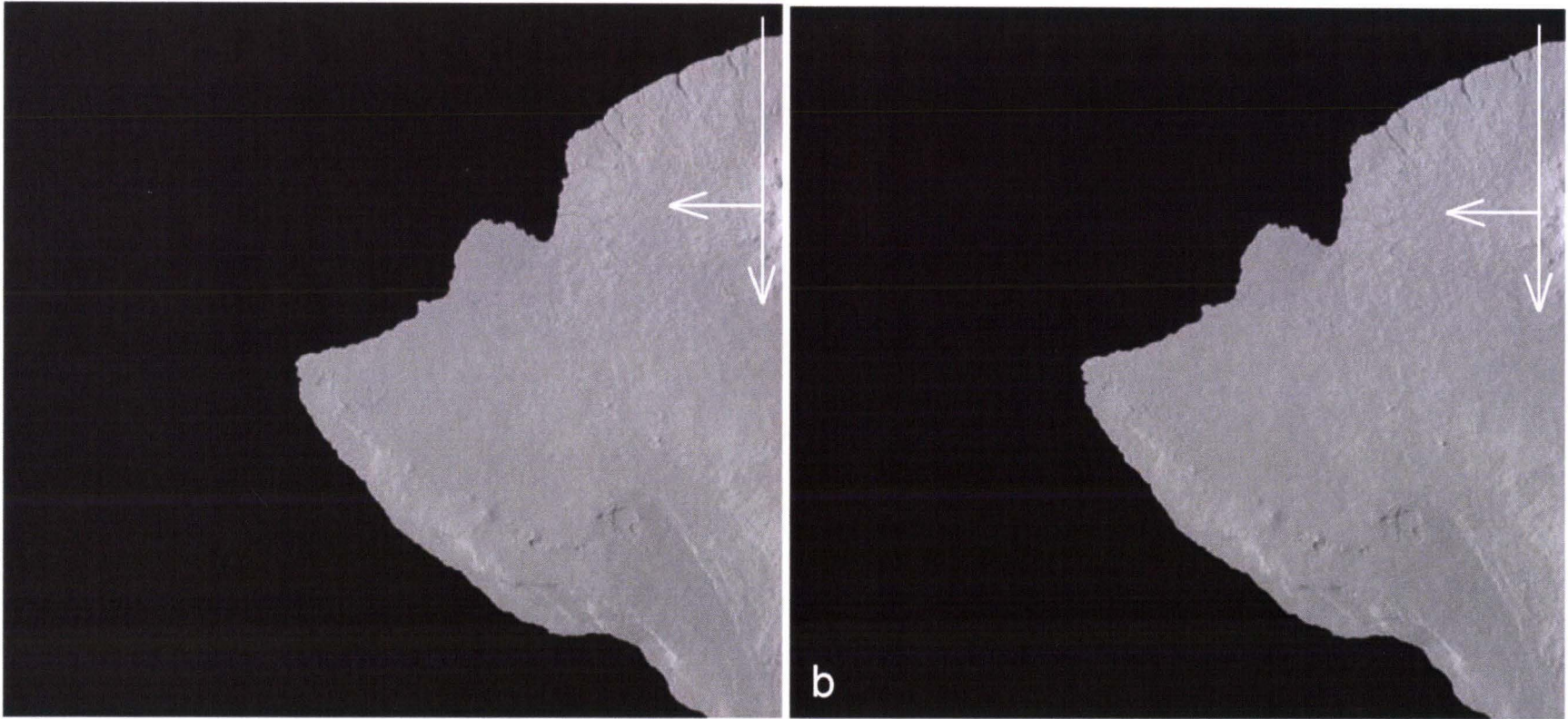


Figure 7: Simulated SAR image in range-Doppler coordinates (RDC). (a) Initial simulation, (b) refined simulation. Differences are in the 1-2 pixel range, hence hardly visible at this scale. The arrow pointing down indicates the satellites flight track whereas the smaller arrow points into the look-direction of the Synthetic Aperture Radar (SAR).

- Display the simulated SAR and the RMLI to detect offsets ([Figure 8](#) PDF 3.2 MB)

```
dis2ras geo/BI_0.sim_rdc.ras rml_i_1_5/rml_i_1_5.ave.ras&
```

Mode 2: Measure initial offset between simulated SAR image and actual SAR image ([log](#))

```
mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI 15 2 2 2 2 8 1 5 6.5 - - - - 256 256
```

The script runs the following commands in mode 2: **create_diff_par**, **init_offsetm**

Depending on the track and frame, the center of the image might be mostly surrounded by water. **init_offsetm** determines a SNR close to the center of the scene and crashes because it cannot find a high enough value to continue with the initial offset estimation. The command can be run manually to find a better region for the offset estimation. In the case shown below, the region was moved more on land (3000 x 3000 instead of the scene center 2587 x 2381.5).

- Create DIFF_par parameter file for image registration and geocoding

```
create_diff_par rml_i_1_5/rml_i_1_5.ave.par - geo/BI.diff 1 < geo/BI.diff.in
```

- Initial offset estimation for multi-look intensity images

```
init_offsetm geo/BI_0.sim_rdc rml_i_1_5/rml_i_1_5.ave geo/BI.diff 1 5 3000 3000 0 0 6.5 128 1
```

Mode 3: Perform refinement of lookup table by offset measurement with respect to the simulated SAR image ([log](#))

```
mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI 15 3 2 2 2 8 1 5 6.5 - - - - 256 256
```

The script was changed to increase the number of estimated offset and to manually cull the offsets.

It runs the following commands in mode 3: **offset_pwrn**, **offset_fitm**

- Measure offsets between MLI images using intensity cross-correlation

```
offset_pwrn geo/BI_0.sim_rdc rml_i_1_5/rml_i_1_5.ave geo/BI.diff geo/BI.off s geo/BI.snr - - - 2 100 100
```

- Range and azimuth offset polynomial estimation

Offsets will be culled manually according to the earlier described method (ISP-3.4).

```
offset_fitm geo/BI.off s geo/BI.snr geo/BI.diff geo/BI.coffs - 6.5 4 1
solution: 1725 offset estimates accepted out of 10000 samples
range offset poly. coeff.:          7.78618    1.68145e-04    6.77191e-04   -2.17621e-07
azimuth offset poly. coeff.:         1.40152   -1.47719e-04   -3.40710e-04    3.55617e-08
model fit std. dev. (samples) range: 0.0964 azimuth:    0.0986
```

Mode 4: Update lookup table and produce terrain geocoded SAR image and DEM in SAR range-Doppler Coordinates (RDC) ([log](#))

```
mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI 15 4 2 2 2 8 1 5 6.5 - - - - 256 256
```

The script runs the following commands in mode 4:

gc_map_fine, geocode, raspwr, dis2ras, geocode, rasshd, disras, geocode_back, raspwr, disras_dem_par

- Geocoding look-up table refinement (with bi-linear fine registration polynomial)

```
gc_map_fine geo/BI_0.map_to_rdc 8334 geo/BI.diff geo/BI_1.map_to_rdc 1
```

- Geocode the simulated SAR image using the refined look-up table

```
geocode geo/BI_1.map_to_rdc geo/BI_0.sim 8334 geo/BI_1.sim_rdc 5174 4763 2 0 1 1 2 8
```

- Create a raster image of the refined simulated SAR in RDC (Figure 7b: [Refined simulated SAR](#))

```
raspwr geo/BI_1.sim_rdc 5174 1 0 1 1
```

- Display the refined simulated SAR and the RMLI to detect offsets ([Figure 9](#) PDF 3.2 MB)

```
dis2ras geo/BI_1.sim_rdc.ras rml_i_1_5/rml_i_1_5.ave.ras&
```

- Use look-up table to transform the MAP DEM to RDC

```
geocode geo/BI_1.map_to_rdc SRTM_BI_15m_utm_seg.dem 8334 geo/BI_dem.rdc 5174 4763 2 0 1 1 2 8
```

- Create a raster image of the DEM in RDC as a shaded relief (Figure 10a: [DEM in RDC](#) PDF 300 KB)

The GAMMA command rashgt can also be used to create a raster image of the DEM (option of choosing m/color-cycle, see [Figure 30b](#) PDF 2.5 MB)

```
rasshd geo/BI_dem.rdc 5174 15 15 1 0 1 1
```

- Display the DEM in RDC

```
disras geo/BI_dem.rdc.ras&
```

- Backward geocoding transformation using the geocoding look-up table

Geocode back the RMLI to MAP geometry (UTM)

```
geocode_back rml_i_1_5/rml_i_1_5.ave 5174 geo/BI_1.map_to_rdc geo/BI_map.mli 8334 7825 1 0
```

- Generate raster image of MLI in MAP geometry (Figure 10b: [RMLI in MAP](#))

```
raspwr geo/BI_map.mli 8334 1 0 1 1
```

- Display the MLI DEM using the SRTM DEM UTM segment parameter file (Figure 11: [RMLI GAMMA display](#) PDF 1 MB)

```
disras_dem_par geo/BI_map.mli.ras SRTM_BI_15m_utm_seg.dem_par&
```

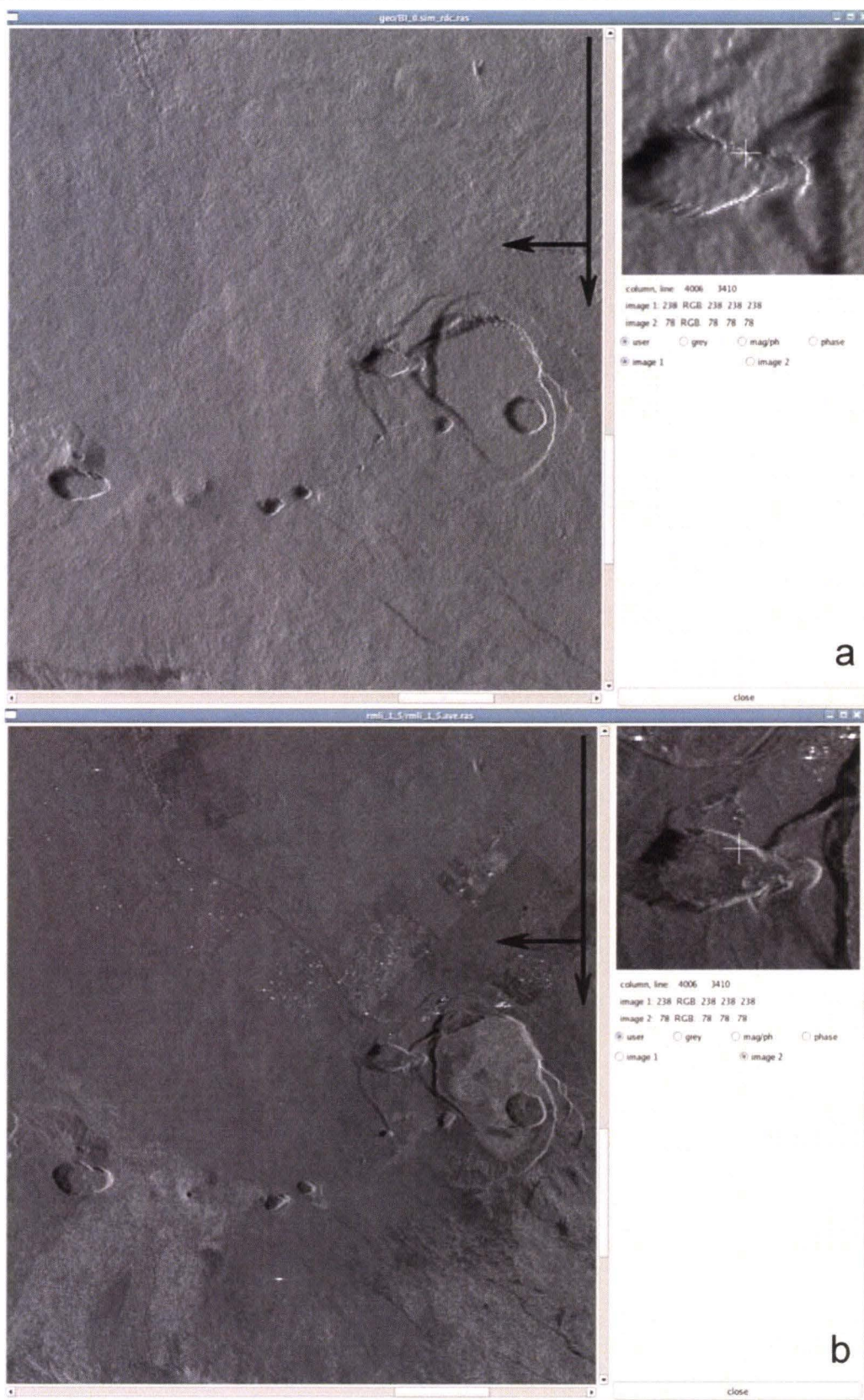



Figure 8: Display of (a) initial simulated SAR image and (b) RMLI using GAMMA dis2ras command. Registration offsets are easiest to detect when focusing on certain features in the zoom window while switching between (a) and (b) using the middle mouse button. Arrows indicate satellite's flight- and the SAR's look-direction.

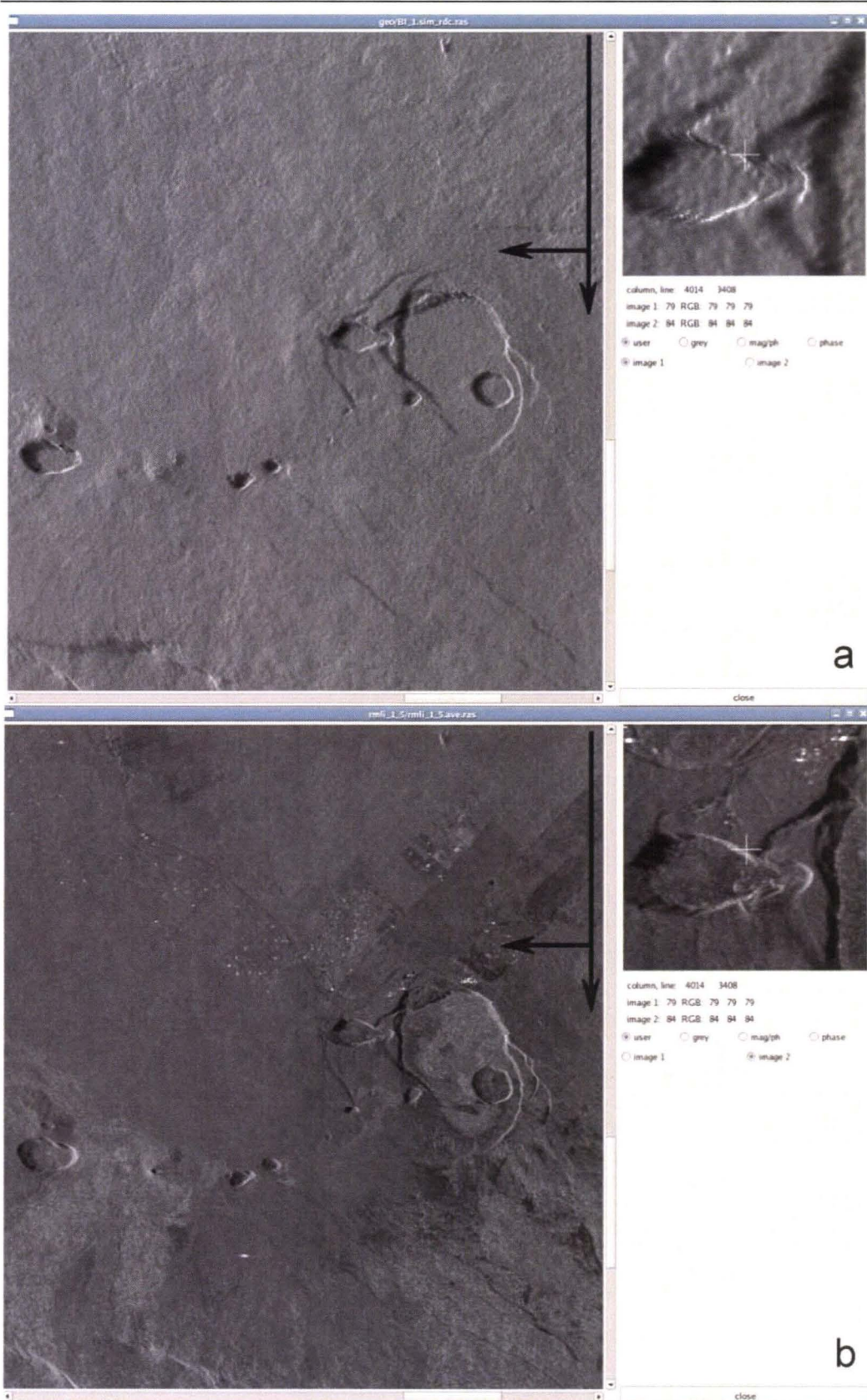


Figure 9: Display of (a) refined simulated SAR image and (b) RMLI using GAMMA dis2ras command. Registration offsets are easiest to detect when focusing on certain features in the zoom window while switching between (a) and (b) using the middle mouse button. Arrows indicate satellite's flight- and the SAR's look-direction.

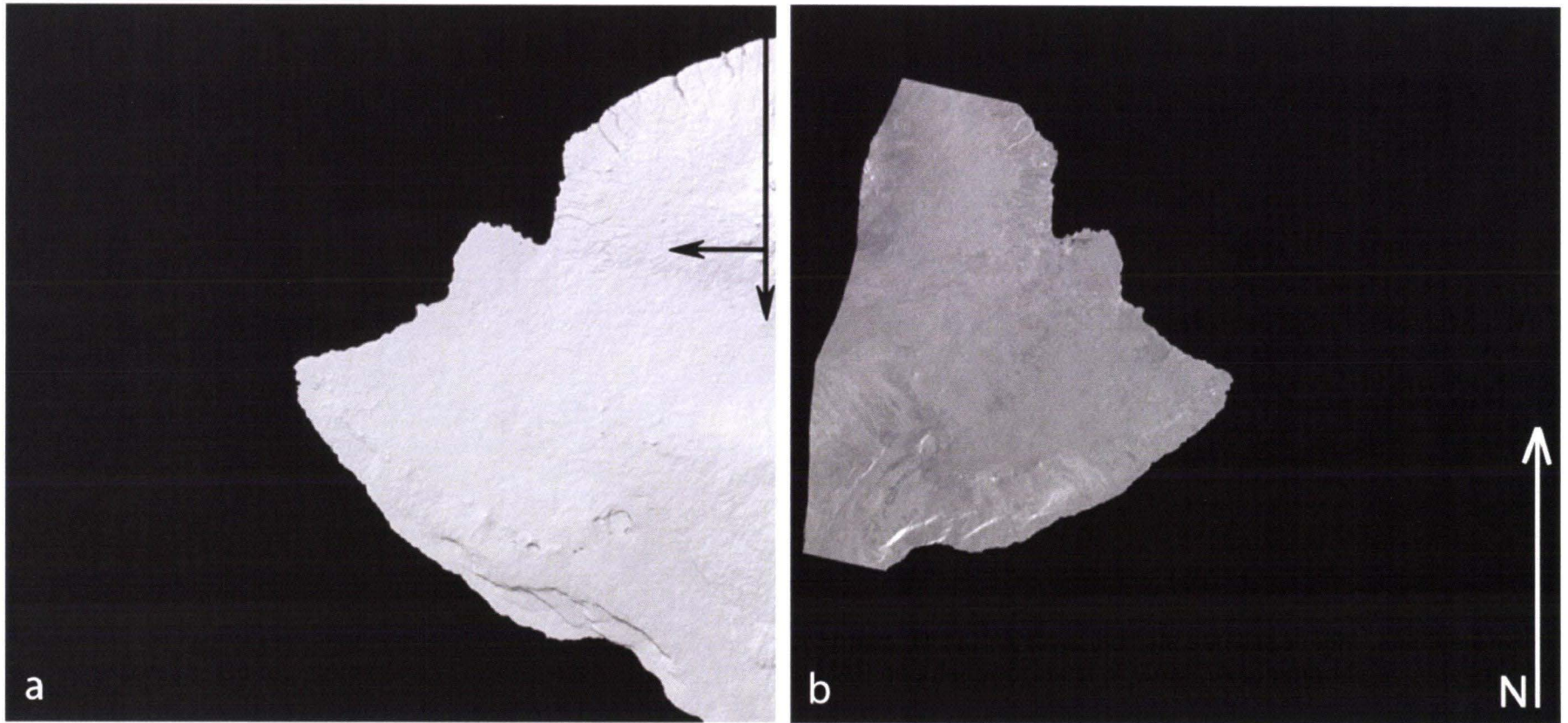


Figure 10: (a) Digital Elevation Model (DEM) in Range-Doppler Coordinates (RDC). Arrows indicate satellite's flight- and the SAR's look-direction, (b) Co-registered Multi-Look Intensity (RMLI) image in MAP geometry (product of GAMMA command geocode_back).

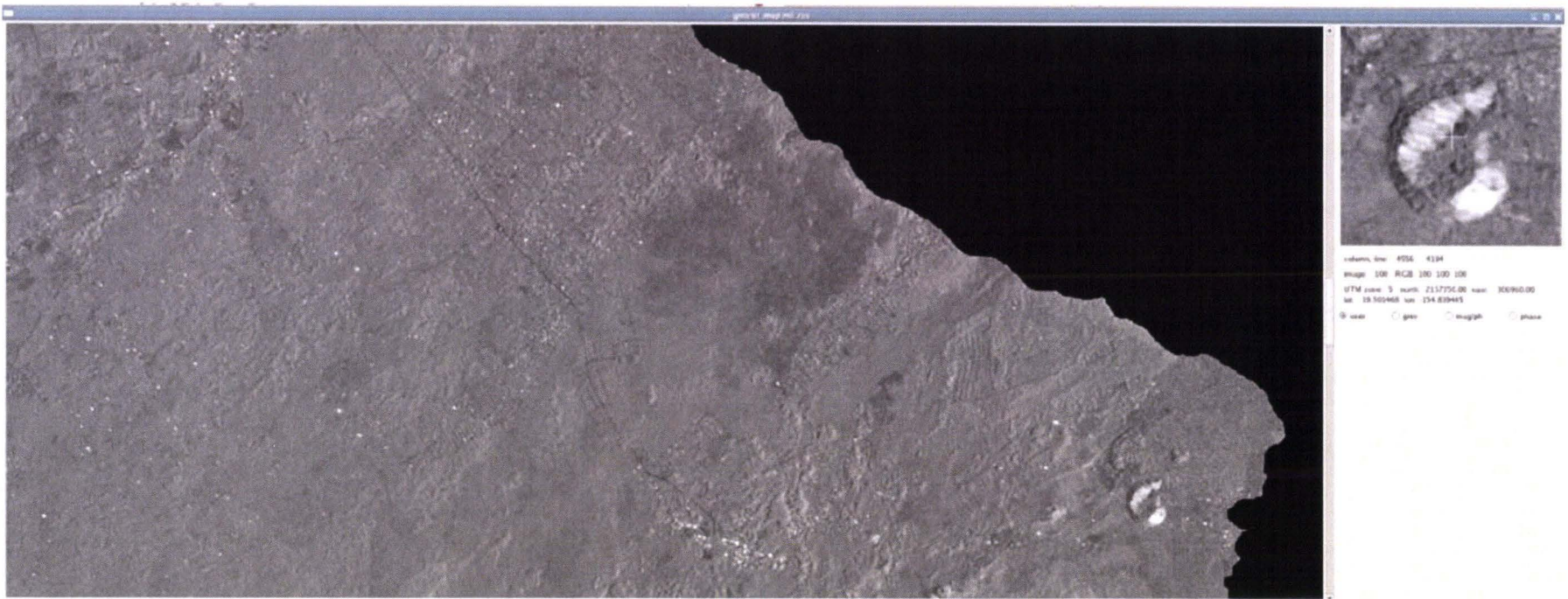
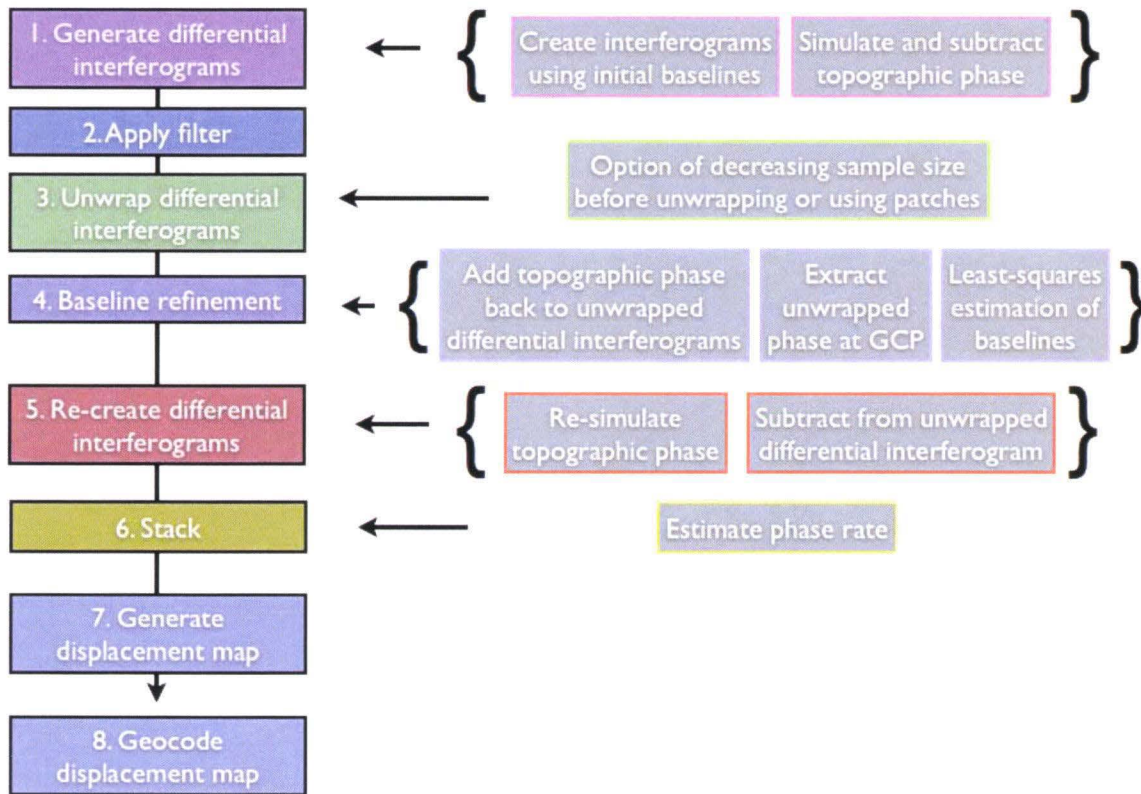


Figure 11: Average co-registered Multi-Look Intensity (RMLI) image converted to MAP geometry (north is up) displayed using GAMMA command `disras_dem_par`.

ISP/DIFF – Interferograms and Differential Interferometry

The following chapter will show how to generate interferograms, to simulate topographic phase and to create differential interferograms. The process of unwrapping will be covered as well as filtering and baseline refinement. The chapter will close with the explanation of stacking and how to produce displacement maps. Please note that it is essential to save your data because scripts are used multiple times in this sections overwriting previous results.

The following flow chart describes the basic processing flow where each step is going to be illuminated in greater detail throughout this chapter. Processing scripts consist of a series of different commands and sometimes can be run in various modes, utilizing a different sequence of commands. Text displayed in the font *courier new* refers to either the content of a file or the script/command that is run and their standard output. The input arguments are separated by spaces. For an explanation of the input/output parameters please refer to **Appendix E** (Index of commands/scripts and their syntax). A command history can be found in the [README](#) file and in the example of the MSP application in **Appendix H**.



ISP/DIFF-1 Generate Itab with All Possible Pair Combinations

The “all combination” itab ([itab_all](#)) is only used in the first [mk_diff_2d](#) run in order to evaluate the quality of long baseline pairs.

```
base_calc SLC_tab slc/20041208.slc.par 20041208.bperp.gr 20041208.bperp itab_all 1
average bperp magnitude (m): 500.162
number of SLC pairs that meet the bperp_max and delta_T_max criteria: 210
```

ISP/DIFF-2 Generate Differential Interferograms

The GAMMA script `mk_diff_2d` is used to generate initial differential interferograms using the orbit derived baselines ([log](#) – 62 MB)

```
mk_diff_2d RSLC_tab itab_all 0 geo/BI_dem.rdc - rml1_1_5/rml1_1_5.ave rml1_1_5 diff0_2d 1 5 3 1
0
```

The script runs the following commands for all scenes specified in the RSLC_tab:

create_offset, **base_orbit**, **SLC_intf**, **rasmph_pwr**, **create_diff_par**, **phase_sim**, **sub_phase**, **rasmph_pwr**, **cc_wave**, **rascc**

- Create interferogram offset files using script generated [off_par.in](#) parameters

```
create_offset rslc/20030212.rslc.par rslc/20030702.rslc.par diff0_2d/20030212_20030702.off <
diff0_2d/off_par.in
```

- Estimate baselines from orbit state vectors (DORIS)

```
base_orbit rslc/20030212.rslc.par rslc/20030702.rslc.par diff0_2d/20030212_20030702.base
```

- Interferogram generation from co-registered SLC data

```
SLC_intf rslc/20030212.rslc rslc/20030702.rslc rslc/20030212.rslc.par rslc/20030702.rslc.par
diff0_2d/20030212_20030702.off diff0_2d/20030212_20030702.int 1 5 0 - 1
```

- Create an image of the interferogram (command not run by script)

Interferometric fringes are visible in close-up (Figure 12: [Interferogram](#) PDF 5.8 MB).

```
rasmph_pwr diff0_2d_1_5/20030212_20030702.int rml1_1_5/20030212.rml1 5174
```

- Create `*.diff_par` using the script generated [diff_par.in](#)

```
create_diff_par diff0_2d/20030212_20030702.off - diff0_2d/20030212_20030702.diff_par <diff0_
2d/diff_par.in
```

- Simulate unwrapped topographic phase using DEM heights

```
phase_sim rslc/20030212.rslc.par diff0_2d/20030212_20030702.off diff0_2d/20030212_20030702.base
geo/BI_dem.rdc diff0_2d/20030212_20030702.sim_unw 0 0 - 140
```

- Create image of simulated unwrapped topography (not part of the script) (Figure 13a,b: [Simulated topography](#))

```
rasrmg diff0_2d/20030212_20030702.sim_unw rml1_1_5/rml1_1_5.ave 5174
```

- Subtract the simulated topographic phase from the interferogram to create differential phase

```
sub_phase diff0_2d/20030212_20030702.int diff0_2d/20030212_20030702.sim_unw diff0_2d/20030212_
20030702.diff_par diff0_2d/20030212_20030702.diff 1
```

- Create image of differential interferogram (Figure 13c,d: [Differential interferogram](#) PDF 4.7 MB)

```
rasmph_pwr diff0_2d/20030212_20030702.diff rml1_1_5/rml1_1_5.ave 5174
```

- Estimate interferogram coherence

```
cc_wave diff0_2d/20030212_20030702.diff rml1_1_5/20030212.rml1 rml1_1_5/20030702.rml1 diff0_
2d/20030212_20030702.cc 5174 3 3 1
```

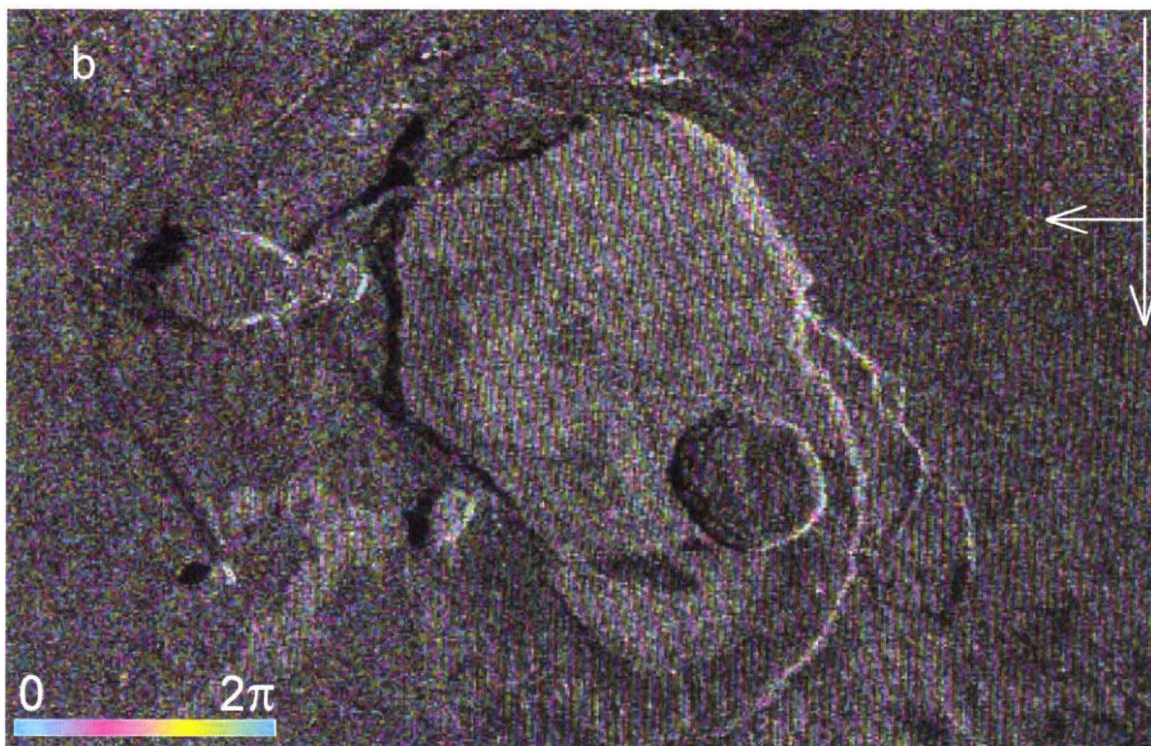


Figure 12: (a) Raster image of wrapped interferogram 20030212_20030702 generated by GAMMA command `rasmph_pwr`, (b) Fringes of interferometric phase visible in close-up of Kilauea Caldera. One color-cycle corresponds to 2π radians.

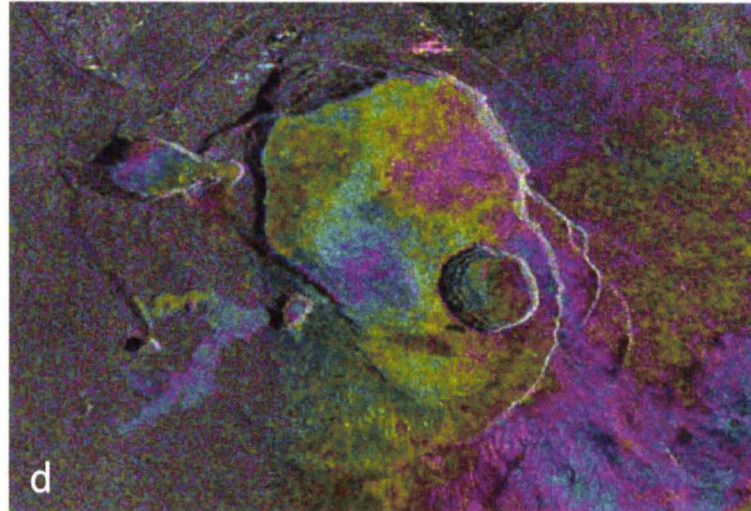
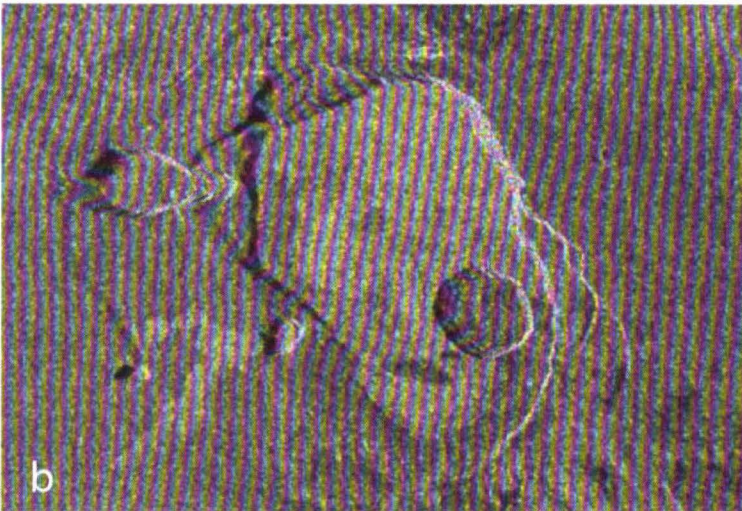
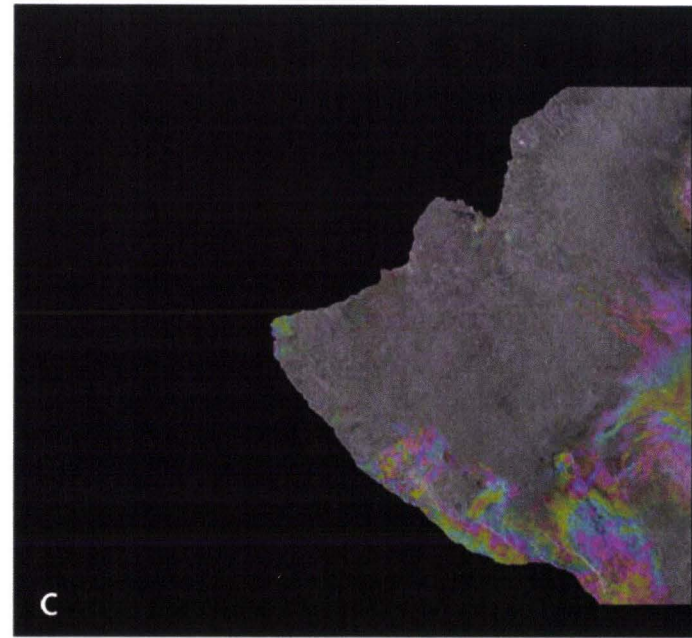


Figure 13: Subtract (a) simulated unwrapped topography from wrapped interferometric phase to generate (c) wrapped differential interferogram. (b) and (d) show a close-up of the Kilauea Caldera. Note how fringes follow topographic features such as the caldera walls. Raster image of unwrapped phase generated with GAMMA command `rasmg`, wrapped phase rasters created with `rasmph_pwr`.

- Generate raster image of correlation coefficient and intensity image (Figure 14: [Interferometric correlation](#) PDF 2.5 MB)

```
rascc diff0_2d/20030212_20030702.cc rml1_1_5/rml1_1_5.ave 5174
```

ISP/DIFF-3 Adjust Itab

Determine for which pairs the processing should be continued (minimum/maximum Bperp) and accordingly adjust the interferometric table.

- View all differential interferograms

```
xv diff0_2d/*.diff.ras&
```

- Based on the quality of the differential interferograms, a new itab is created with perpendicular baselines ranging from -750 m to 750 m.

```
average bperp magnitude (m): 359.564
```

```
number of SLC pairs that meet the bperp_max and delta_T_max criteria: 166
```

```
base_calc SLC_tab slc/20041208.slc.par 20041208.bperp.gr 20041208.bperp itab 1 0 750 -
```

ISP/DIFF-4 Apply ADF Filter to the Stack of Differential Interferograms ([log](#))

```
mk_adf_2d RSLC_tab itab rml1_1_5/rml1_1_5.ave diff0_2d 5 .3 32
```

The script runs the following command for every pair specified in the itab: **adf**

- Adaptive spectral filtering for complex interferograms

```
adf diff0_2d/20030212_20030702.diff diff0_2d/20030212_20030702.adf.diff diff0_2d/20030212_20030702.adf.cc 5174 .3 32 5 8 0 0 .2
```

- Generate raster images of filtered differential interferograms (Figure 15: [Comparison filtered and unfiltered phase](#) PDF 12.8 MB)

```
rasmph_pwr diff0_2d/20030212_20030702.adf.diff rml1_1_5/rml1_1_5.ave 5174 1 1 0 1 1 1. .35
```

ISP/DIFF-5 Unwrap the Differential Interferograms

In the previous step (ISP/DIFF-2) complex interferograms are generated from the co-registered SLC. The interferometric phase is only known modulo 2π . The correct multiple of 2π must be found and added using one of the two common unwrapping algorithms. The method used in this example is the minimum cost flow (MCF) unwrapping algorithm. It is a global optimization technique that utilizes masking, adaptive thinning and processing in patches to allow unwrapping of very large scenes. Even critical areas (low coherence, gaps in the data) can be addressed. It is therefore a widely accepted and preferred method. An alternative method is a branch-cut region growing technique [Rosen, et al., 2000]. Areas of low coherence and layover are masked out, as well as locations around which a closed integral of phase differences would lead to a non-zero result (a.k.a. residue). Branch-cuts are constructed and the phase unwrapped without crossing any branch-cuts. This process is continued by region growing for the entire area connected. GAMMA processing strategies for phase unwrapping are described in Wegmüller et al. [2002].

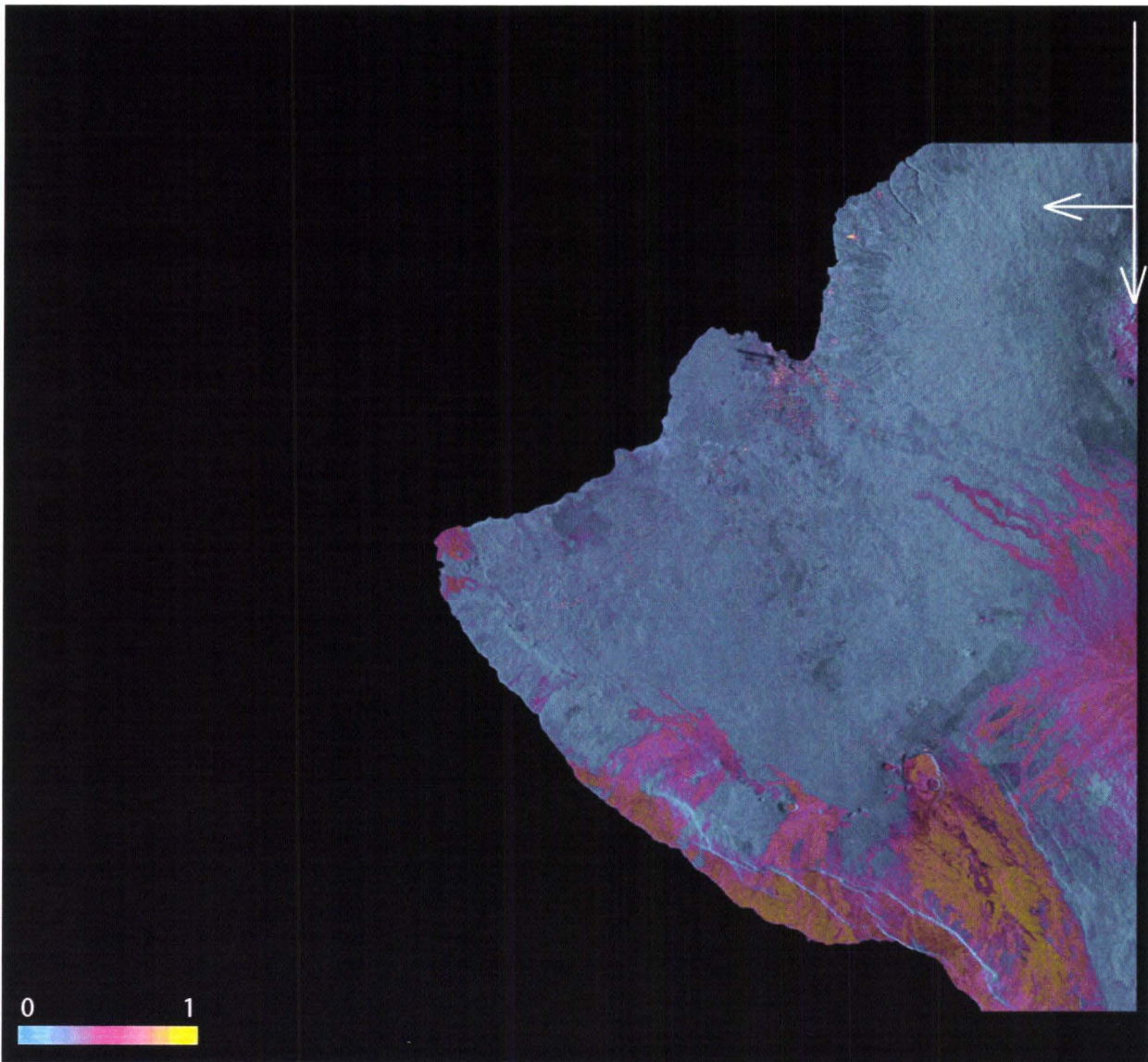


Figure 14: Degree of coherence (interferometric correlation) for pair 20030212_20030702. Color corresponds to interferometric correlation, i.e., low correlation coefficient represented by blue colors and high coefficient by yellow colors, i.e. very good correlation. Brightness relates to backscatter intensity. Arrows indicate satellite's flight- and the SAR's look-direction.

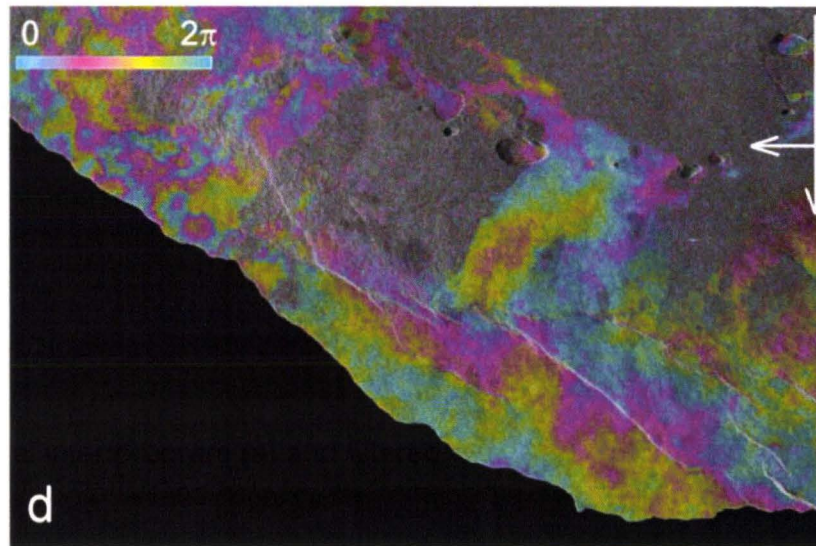
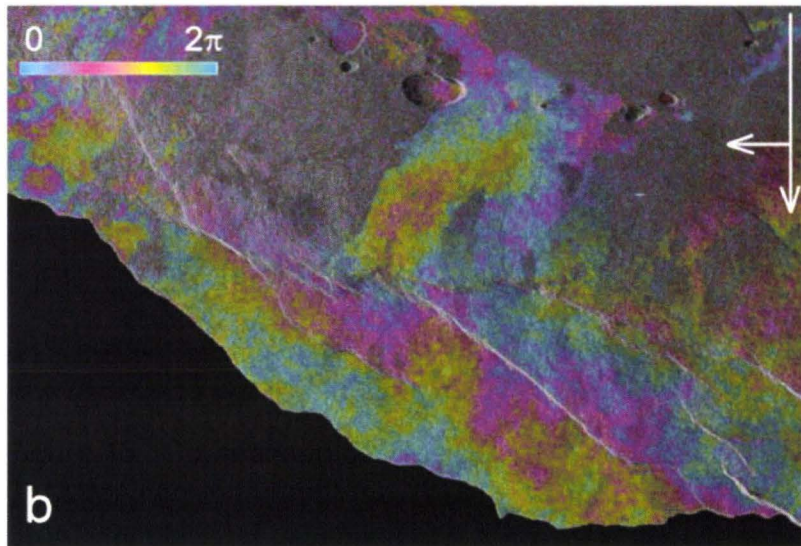
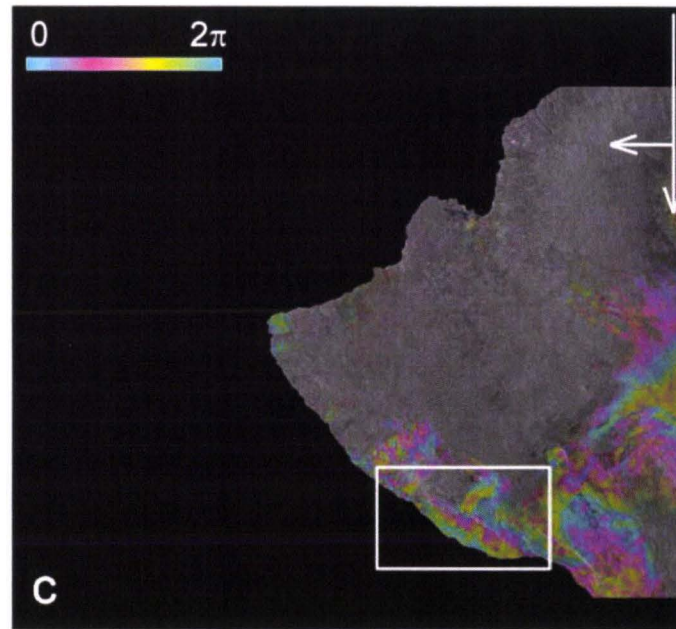
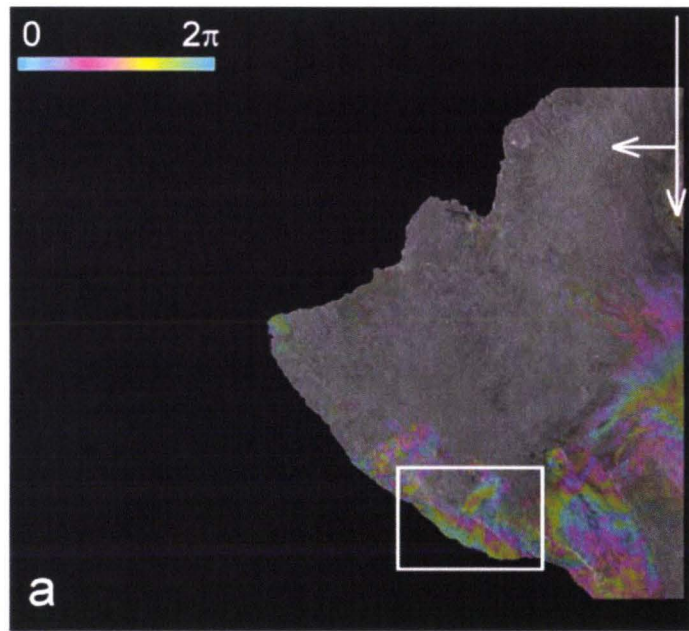


Figure 15: Comparison between unfiltered wrapped differential interferogram (a) and filtered wrapped interferogram (GAMMA command adf) (c). (b) and (d) show close-ups of Hilina Pali. One color-cycle corresponds to 2π radians. Arrows indicate satellite's flight- and the SAR's look-direction.

If the entire scene is processed rather than a small subset, it is recommended to “multi-look down” the wrapped differential interferogram before attempting to unwrap. The differential interferogram as well as the coherence estimate will be scaled down, in this example by a factor of two (-2), cutting the range and azimuth sample size in half. This will minimize the memory allocation necessary for the Minimum Cost Flow (MCF) method ([log](#)).

```
mk_unw_2d RSLC_tab itab rml_i_1_5/rml_i_1_5.ave diff0_2d .4 2 1 1 1
```

The script runs the following commands for every interferogram specified in the itab:

rascc_mask, multi_cpx, multi_real, mcf, multi_real, unw_model, rasrmg, rasrmg

- Generate phase unwrapping validity mask ([Figure 16a](#) PDF 2.8 MB)

Pixels with 0 values are not considered in further processing.

```
rascc_mask diff0_2d_1_5/20030212_20030702.adf.cc rml_i_1_5/rml_i_1_5.ave 5174 1 1 0 2 2 .4 .3 1.0
1. .35 1 diff0_2d_1_5/20030212_20030702.adf.cc_mask.ras
```

- Multi-look the differential interferogram 2 x 2

Transformation between complex image ↔ multi-look complex image (averaging and oversampling)

```
multi_cpx diff0_2d_1_5/20030212_20030702.adf.diff diff0_2d_1_5/20030212_20030702.off diff0_2d_1_5/diff2
diff0_2d_1_5/off2 2 2
```

- Multi-look the coherence estimate 2 x 2

Conversion between real image ↔ multi-look real image (averaging and oversampling)

```
multi_real diff0_2d_1_5/20030212_20030702.adf.cc diff0_2d_1_5/20030212_20030702.off diff0_2d_1_5/cc2
diff0_2d_1_5/off2 2 2
```

- Phase unwrapping using MCF and triangulation

```
mcf diff0_2d_1_5/diff2 diff0_2d_1_5/cc2 diff0_2d_1_5/20030212_20030702.adf.cc_mask.ras diff0_2d_1_5/unw2
2587 1 0 0 2587 2381.5 1 1 128
```

- Over-sample the unwrapped multi-looked differential interferogram -2 x -2

```
multi_real diff0_2d_1_5/unw2 diff0_2d_1_5/off2 diff0_2d_1_5/unw3 diff0_2d_1_5/off3 -2 -2
```

- Phase unwrapping of the not multi-looked differential interferogram using a model of the unwrapped phase (over-sampled multi-looked unwrapped differential phase)

```
unw_model diff0_2d_1_5/20030212_20030702.adf.diff diff0_2d_1_5/unw3 diff0_2d_1_5/20030212_20030702.adf.unw
5174 0 0
```

- Create raster images of unwrapped differential phase and intensity data ([Figure 17a, b](#) PDF 13 MB)

Notice the difference in the amount of fringes by choosing different phase scale factors.

```
rasrmg diff0_2d_1_5/20030212_20030702.adf.unw rml_i_1_5/rml_i_1_5.ave 5174 1 1 0 1 1 1 .35 .0
-1 diff0_2d_1_5/20030212_20030702.adf.unw1.ras diff0_2d_1_5/20030212_20030702.adf.cc 1 .2
rasrmg diff0_2d_1_5/20030212_20030702.adf.unw rml_i_1_5/rml_i_1_5.ave 5174 1 1 0 1 1 .3333 1. .35
.0 -1 diff0_2d_1_5/20030212_20030702.adf.unw_3.ras diff0_2d_1_5/20030212_20030702.adf.cc 1 .2
```

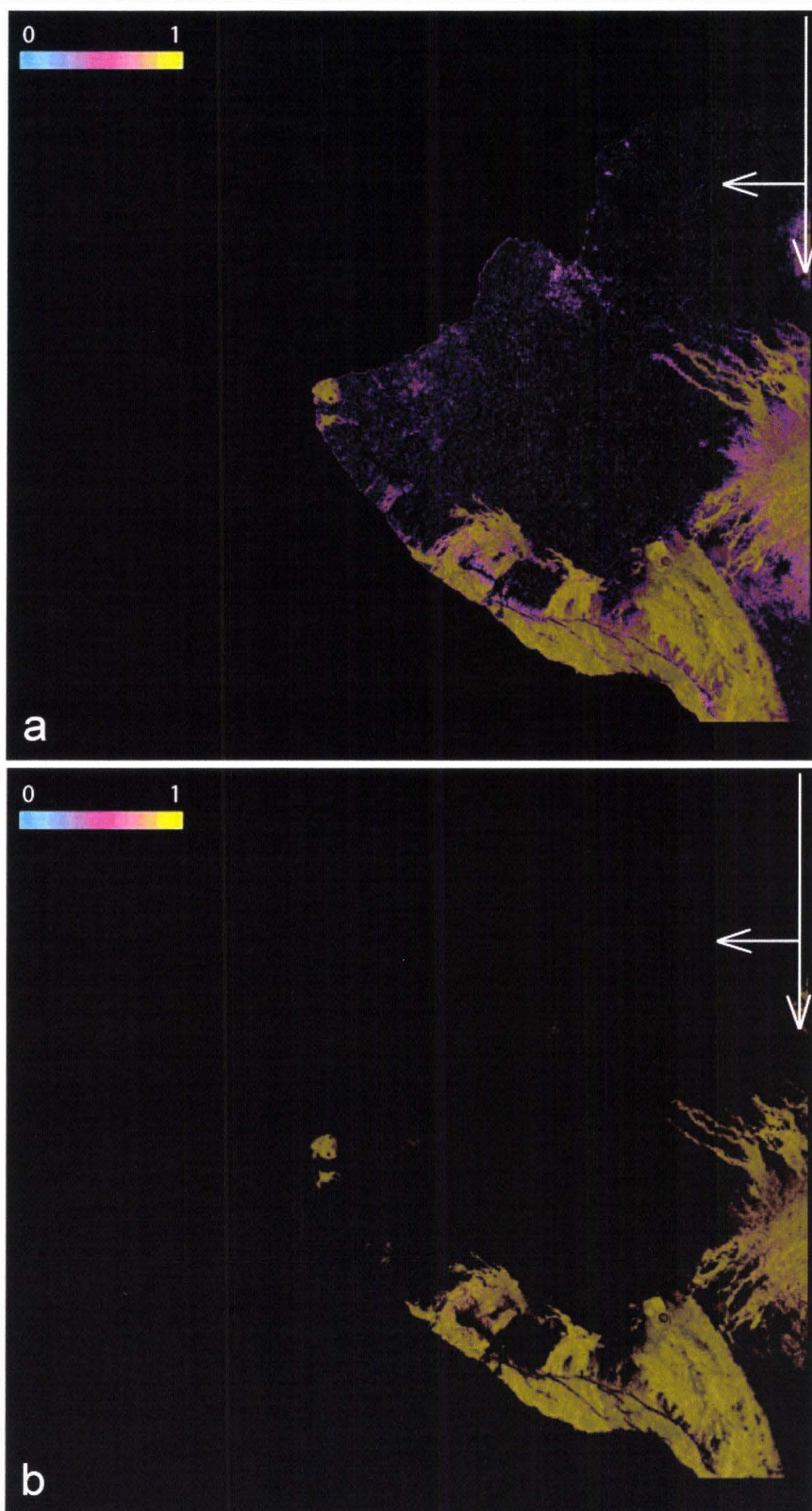


Figure 16: Masks generated using GAMMA command `rascc_mask`. Thresholds are (a) 0.4 and (b) 0.8, pixels with coherence below these values are set to (0,0,0). Color corresponds to interferometric correlation, i.e., low correlation coefficient represented by blue colors and high coefficient by yellow colors, i.e. very good correlation. Brightness relates to backscatter intensity. Arrows indicate satellite's flight- and the SAR's look-direction.

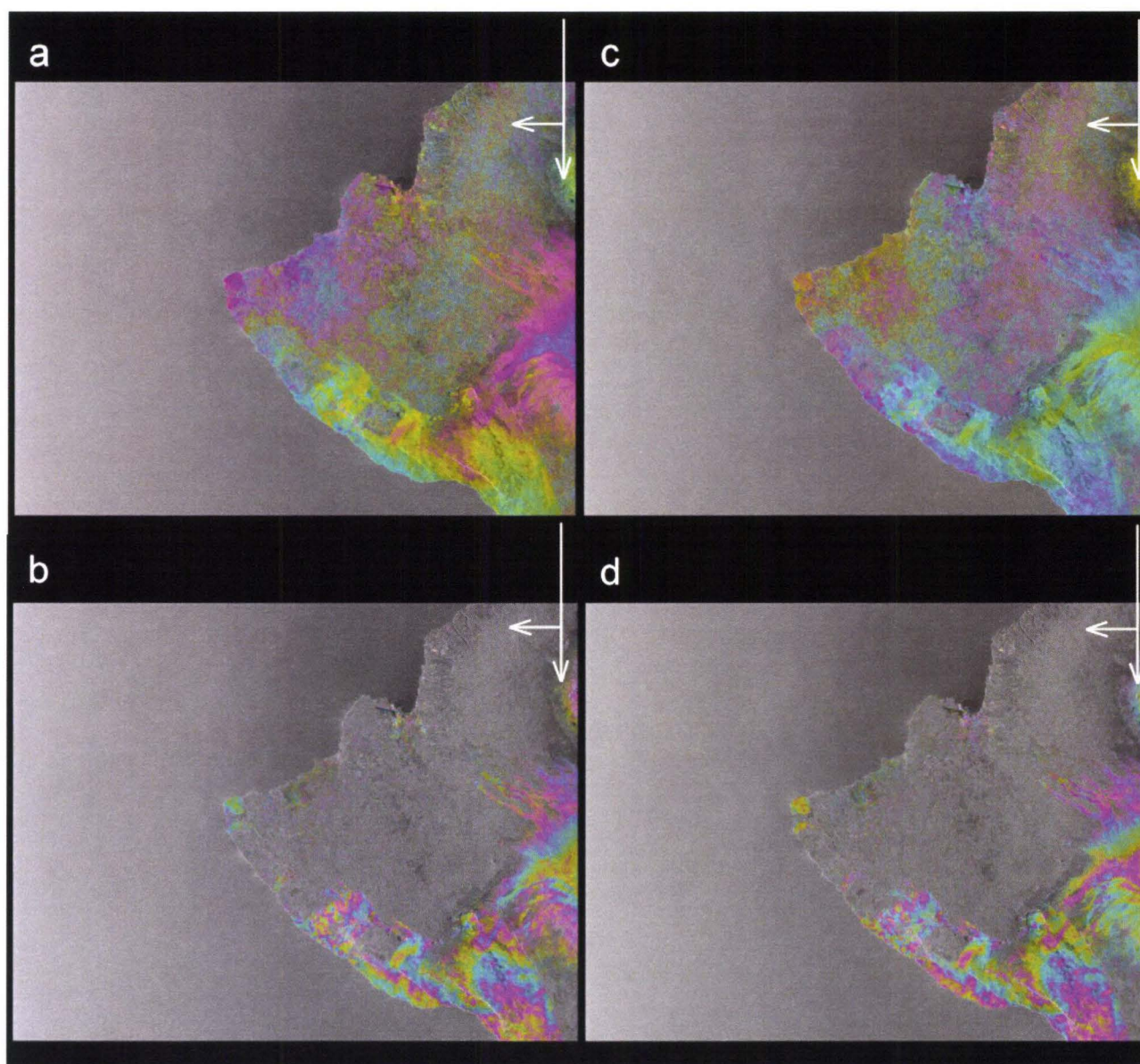


Figure 17: Comparison between unwrapped differential interferograms generated using initial orbit derived baselines (a, b) and refined baselines (c, d). Phase display scale factor 0.3333 (a, c) and 1 (b, d). Arrows indicate satellite's flight- and the SAR's look-direction.

ISP/DIFF-6 Baseline Refinement

The two complex SAR images used to create an interferogram can either be obtained from two temporally separated passes of the same satellite or from two satellites during a tandem mission. In either case, the position of the two acquisition times is separated by a certain distance, also referred to as the baseline. (Figure 18: [Baseline geometry](#) PDF 200 KB). The baseline can be divided into its components, parallel and perpendicular baseline. Knowing this distance very accurately is most important in interferometric processing. The initial baselines are determined from the state orbit vectors. Precision estimation is crucial for interpretation of the satellite phase as elevation. Once the baseline is refined, the topographic phase will be re-simulated and the differential interferograms are re-calculated.

The baseline refinement is based on selecting valid ground control points (GCP). A mask has to be generated that masks out all incoherent areas like water and decorrelation affected areas like vegetation. The unwrapped phase values at the GCP are then extracted to be used in a least-squares fit to calculate more precise baselines. The GAMMA script [mk_base_2d](#) is modified to save the unwrapped interferograms. The topographic phase is re-simulated using the successfully refined baselines ([mk_diff_2d_cs](#)), to be subtracted from the now unwrapped interferograms to create unwrapped differential interferograms. No unwrapping is necessary that way.

ISP/DIFF-6.1 Create Mask

- Create a mask based upon a very high correlation threshold ([Figure 16b](#))

```
rascc_mask diff0_2d/20030212_20030702.adf.cc rml1_1_5/rml1_1_5.ave 5174 1 1 0 1 1 .8 .3 1.0 1.
.35 1 diff0_2d/20030212_20030702.adf.cc_mask.8.ras
```

ISP/DIFF-6.2 Estimate Improved Baselines from Unwrapped Phase and DEM in RDC ([log](#))

```
mk_base_2d_cs RSLC_tab itab geo/BI_dem.rdc diff0_2d - diff0_2d/20030212_
20030702.adf.cc_mask.8.ras 128 128 5 1
```

The script runs the following commands for the pairs specified in the interferometric table itab:

sub_phase, extract_gcp, gcp_phase, base_ls

- Add previously estimated topographic phase back to differential unwrapped interferogram

```
sub_phase diff0_2d/20030212_20030702.adf.unw diff0_2d/20030212_20030702.sim_unw diff0_
2d/20030212_20030702.diff_par diff0_2d/20030212_20030702.unw_int 0 1
```

- Automatically extract [GCP](#) from a DEM in RDC

```
extract_gcp geo/BI_dem.rdc diff0_2d/20030212_20030702.off diff0_2d/gcp 128 128 diff0_
2d/20030212_20030702.adf.cc_mask.8.ras
```

- Extract unwrapped phase ([gcp_ph](#)) at GCP locations

```
gcp_phase diff0_2d/20030212_20030702.unw_int diff0_2d/20030212_20030702.off diff0_2d/gcp diff0_
2d/20030212_20030702.gcp_ph 3
```

- Precisely estimate the interferometric baseline using GCP, corresponding height and unwrapped interferometric phase in a least squares approach

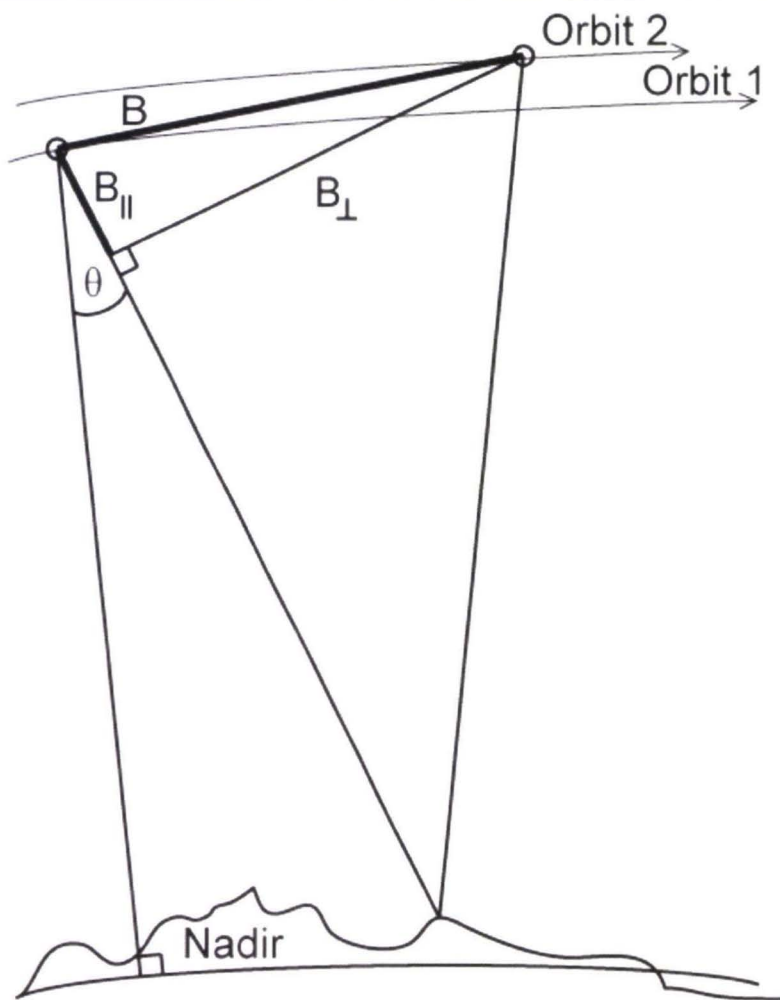


Figure 18: Baseline geometry, after R. Hanssen

```
base_ls rslc/20030212.rslc.par diff0_2d/20030212_20030702.off diff0_2d/20030212_20030702.gcp_ph
diff0_2d/20030212_20030702.base 0 1 1 1 0 1.
```

Check an [example](#) by displaying a baseline file:

```
cat diff0_2d/20030212_20030702.base
```

- Copy necessary files to the new directory *diff1_2d* where 1 indicates that refined baselines are used

```
mkdir diff1_2d
cp diff0_2d/*.base diff1_2d/
mv diff0_2d/*.unw_int diff1_2d/
cp diff0_2d/*.off diff1_2d/
```

The **adf.cc* are moved for displaying purposes.

```
mv /mnt/scsi/PGFSAR/IPTA/2429_3213/diff0_2d/*adf.cc diff1_2d/
```


ISP/DIFF-7 Re-simulate Topography and Create Refined Differential Interferograms

The modified version `mk_diff_2d_cs` uses the unwrapped interferograms and the refined baselines to calculate the updated differential interferograms. Notice the `*.p` naming convention, where `p` indicates precise.

```
mk_diff_2d_cs RSLC_tab itab 1 geo/BI_dem.rdc - rml1_1_5/rml1_1_5.ave rml1_1_5 diff1_2d 1 5 3 1
0
```

- Create new differential parameter file

```
create_diff_par diff1_2d/20030212_20030702.off - diff1_2d/20030212_20030702.diff_par <diff1_2d/diff_par.in
```

- Re-simulate topographic phase using refined baseline

```
phase_sim rslc/20030212.rslc.par diff1_2d/20030212_20030702.off diff1_2d/20030212_20030702.base
geo/BI_dem.rdc diff1_2d/20030212_20030702.sim_unw.p 0 1 - 140
```

- Subtract simulated topographic phase from unwrapped interferograms to create refined differential unwrapped interferograms

```
sub_phase diff1_2d/20030212_20030702.unw_int diff1_2d/20030212_20030702.sim_unw.p diff1_2d/20030212_20030702.diff_par
diff1_2d/20030212_20030702.diff_unw 0
```

- Create images of unwrapped differential interferograms ([Figure 17c, d](#))

```
rasrmg diff1_2d/20030212_20030702.diff_unw rml1_1_5/rml1_1_5.ave 5174 1 1 0 1 1 1 1. .35 .0 -1
diff1_2d/20030212_20030702.adf.diff_unw1.ras diff1_2d/20030212_20030702.adf.cc 1 .2
rasrmg diff1_2d/20030212_20030702.diff_unw rml1_1_5/rml1_1_5.ave 5174 1 1 0 1 1 .3333 1. .35 .0
-1 diff1_2d/20030212_20030702.adf.diff_unw_3.ras diff1_2d/20030212_20030702.adf.cc 1 .2
```

- Compare the refined baseline interferograms with the initial baseline ones ([Figure 17: Comparison initial and refined baselines](#))

```
dis2ras diff0_2d/20030212_20030702.adf.unw1.ras diff1_2d/20030212_20030702.adf.diff_unw1.ras
```

ISP/DIFF-8 Stack the Data

Stacking requires unwrapped phase and gives the average deformation rate over the time period of the different interferograms summed. A list of unwrapped differential interferograms is created only considering pairs with appropriate baseline and time spans. Depending on the project, only short time intervals, short perpendicular baselines or certain times might be considered. The `diff_tab` will consist of two columns, where the first column is a list of the differential unwrapped interferograms (refined) and the second column lists the corresponding `delta_t`.

- Create `diff_tab` for all pairs of interest, use `base_calc` output for `delta_t`
- Stack the data

Estimate the phase rate (linear velocity) by stacking multiple unwrapped differential interferograms. A phase reference region has to be determined by viewing all differential interferograms and finding a stable region that is coherent throughout the entire stack i.e., has values at that point. `Xv` can be used for displaying. All motion is with respect to the reference region. The expected nominal error for an interferogram with relatively good coherence is about 1/20 of the wavelength accuracy in the path length. The path length errors originate through atmosphere, ionospheric delay, residual topographic phase due to baseline uncertainties and DEM errors. The

goal is to determine the deformation due to the path length difference. The approach of stacking averages out the atmosphere and baseline related errors. One complete cycle (2π) is equivalent to $\frac{1}{2}$ the wavelength difference in path length (the wavelength for Envisat is 5.6 cm). The convention in the GAMMA software is that positive phase in a differential interferogram means negative deformation (i.e., subsidence). This assumes that the interferogram is generated using the earlier scene as the reference. *Sig_ph_rate* and *sig_ph* are standard deviation of (1) the estimated phase rate (in radians/year) and of (2) the residual phases (in radians). For a more detailed discussion see the GAMMA help on [stacking](#).

```
stacking diff_tab 5174 ph_rate sig_ph_rate sig_ph 2041 2413 - - -
```

ISP/DIFF-9 Create Displacement Map in RDC

- Conversion of unwrapped differential phase to displacement map (m)

```
dispmap ph_rate - slc/20041208.slc.par diff1_2d/20051228_20060308.off ph_rate_disp
```

- Display displacement map in GAMMA to view actual values (Figure 19: [Displacement close-ups](#) PDF 9.5 MB)

```
dishgt ph_rate_disp rml1_1_5/rml1_1_5.ave 5174 1 1 0 0.01 1 .35
```

- Create a raster of the displacement (Figure 20a: [Displacement in RDC](#) PDF 1.7 MB)

```
rashgt ph_rate_disp rml1_1_5/rml1_1_5.ave 5174 1 1 0 1 1 .01 1. .35 -1 ph_rate_disp.ras
```

ISP/DIFF-10 Geocode Back Displacements to Transform to UTM

- Transform the displacement into UTM coordinates

```
geocode_back ph_rate_disp 5174 geo/BI_1.map_to_rdc ph_rate_disp.utm 8334 0 0 0
```

- Transform coherence estimate to UTM

```
geocode_back diff1_2d/20031119_20040303.adf.cc 5174 geo/BI_1.map_to_rdc disp.cc 8334 7825 1 0
```

- Display the UTM displacement in GAMMA (Figure 21: [Displacement close-ups in UTM](#) PDF 8.2 MB)

```
dishgt ph_rate_disp.utm geo/BI_map.mli 8334 1 1 0 0.01 1 .35
```

- Create a raster of the UTM displacement (Figure 20b: [Displacement in UTM](#))

```
rashgt ph_rate_disp.utm geo/BI_map.mli 8334 1 1 0 1 1 .01 1. .35 1 ph_rate_disp.utm.ras
```

- Create a raster with the pwr image in the background (Figure 22: [Displacement map with pwr background](#) PDF 1.3 MB)

Here, you must manually determine out the phase display scale factor to match the m/color-cycle chosen in rashgt.

In this case 0.01 m per color-cycle in rashgt match a phase display scale factor of 630 ($200*\pi$).

```
rasrmg ph_rate_disp.utm geo/BI_map.mli 8334 1 1 0 1 1 630 1. .35 0 1 ph_rate_disp.utm1.ras  
disp.cc 1 .2|
```

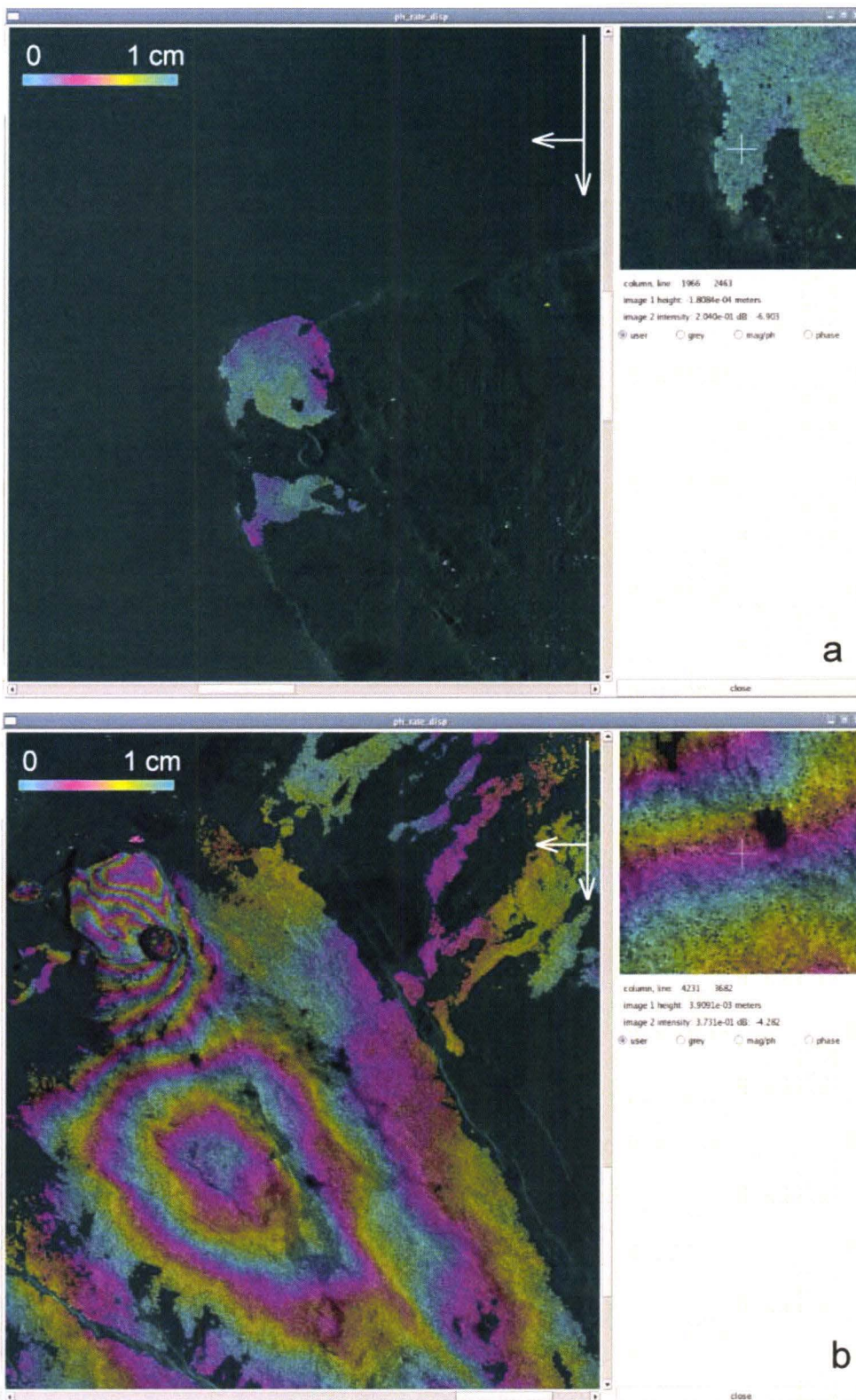


Figure 19: Displacement map in RDC using GAMMA command dishgt at 0.01 m/color-cycle. The displacement is along the look-vector of the SAR and showing the average motion from 20030212 and 20060308. (a) Kapoho region, magenta values range between -2.7 and -3.5 mm, cyan values: -0.14 and -0.25 mm. (b) Kilauea, center of the ellipsoidal shape south of Kilauea caldera measures ~ -10 mm, whereas left of Halemaumau shows values of ~ 50mm. Arrows indicate satellite's flight- and the SAR's look-direction.

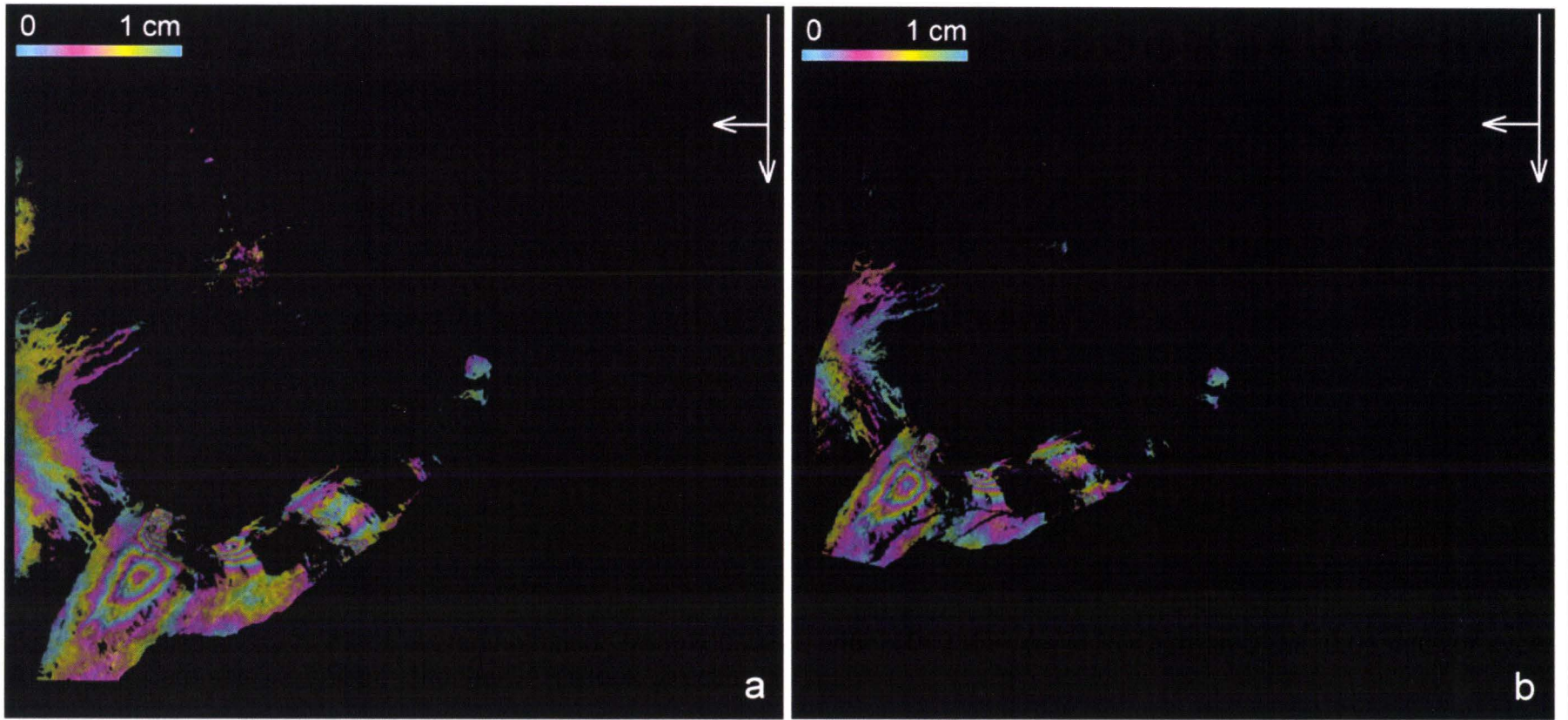


Figure 20: Displacement maps averaging motion from 20030212 until 20060308. (a) in RDC, (b) in UTM. 0.01 m/color-cycle. Arrows indicate satellite's flight- and the SAR's look-direction.

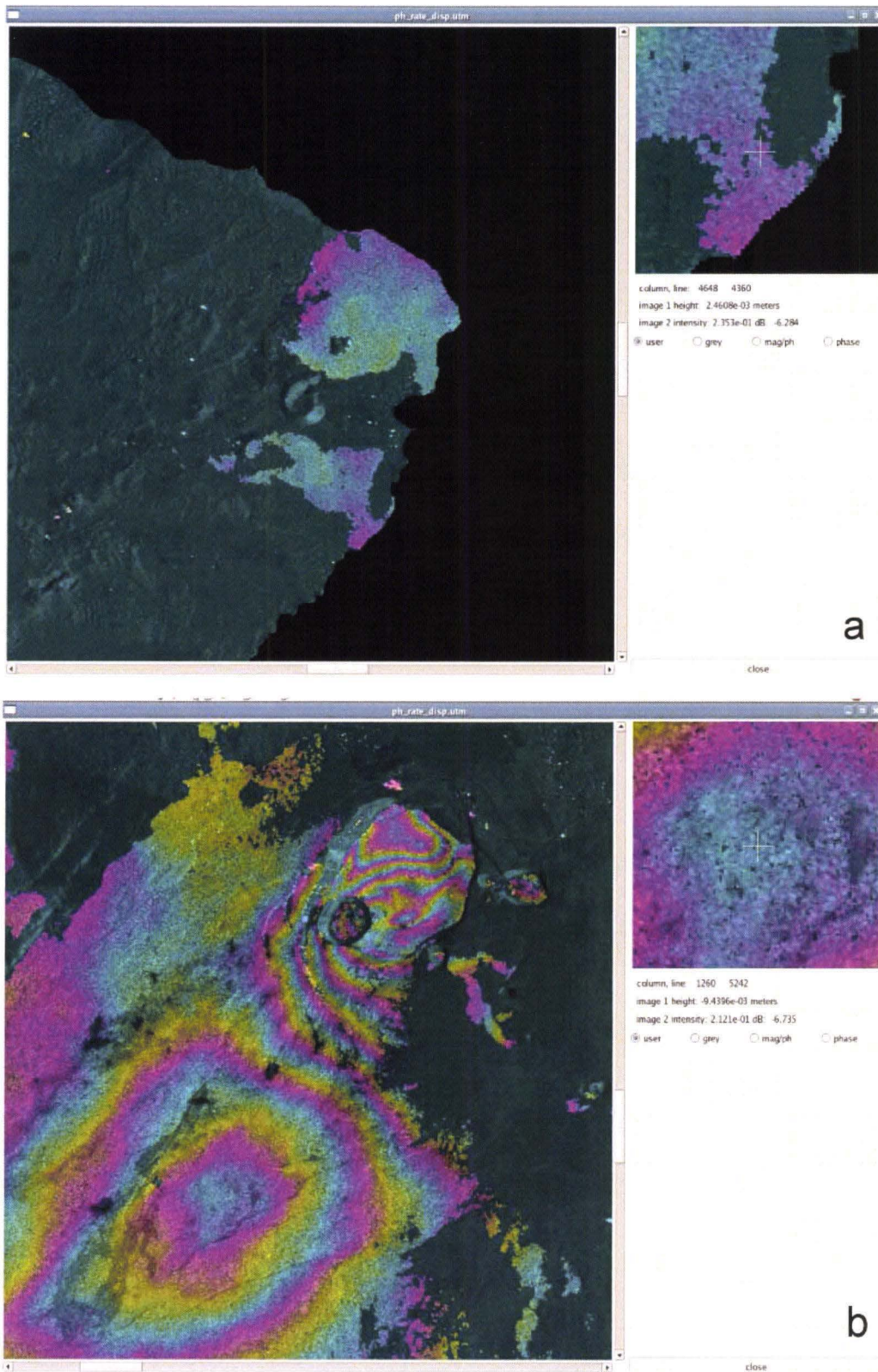


Figure 21: Displacement maps in UTM (0.01 m/color-cycle) showing (a) Kapoho region and (b) Kilauea. North is up.

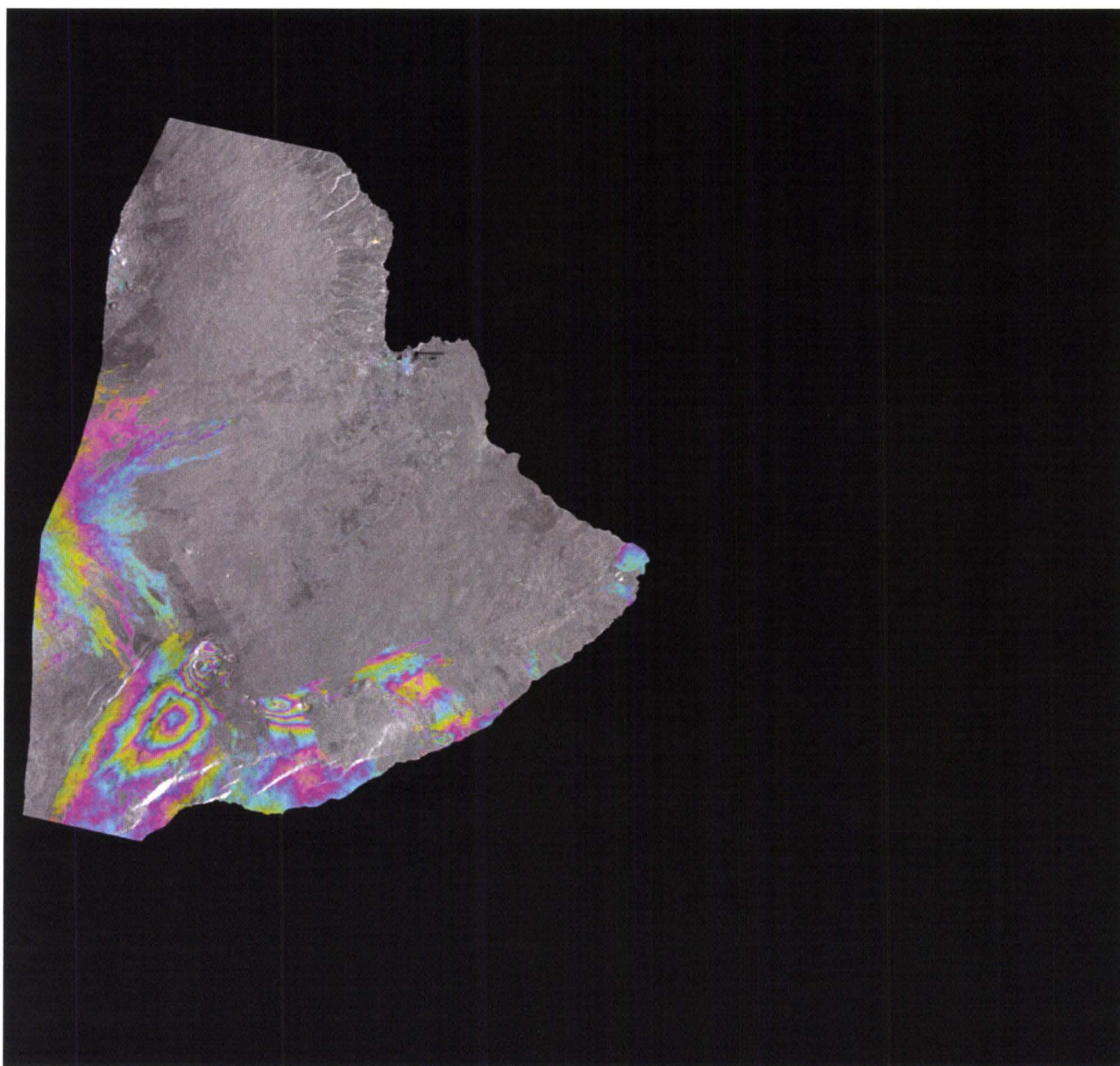


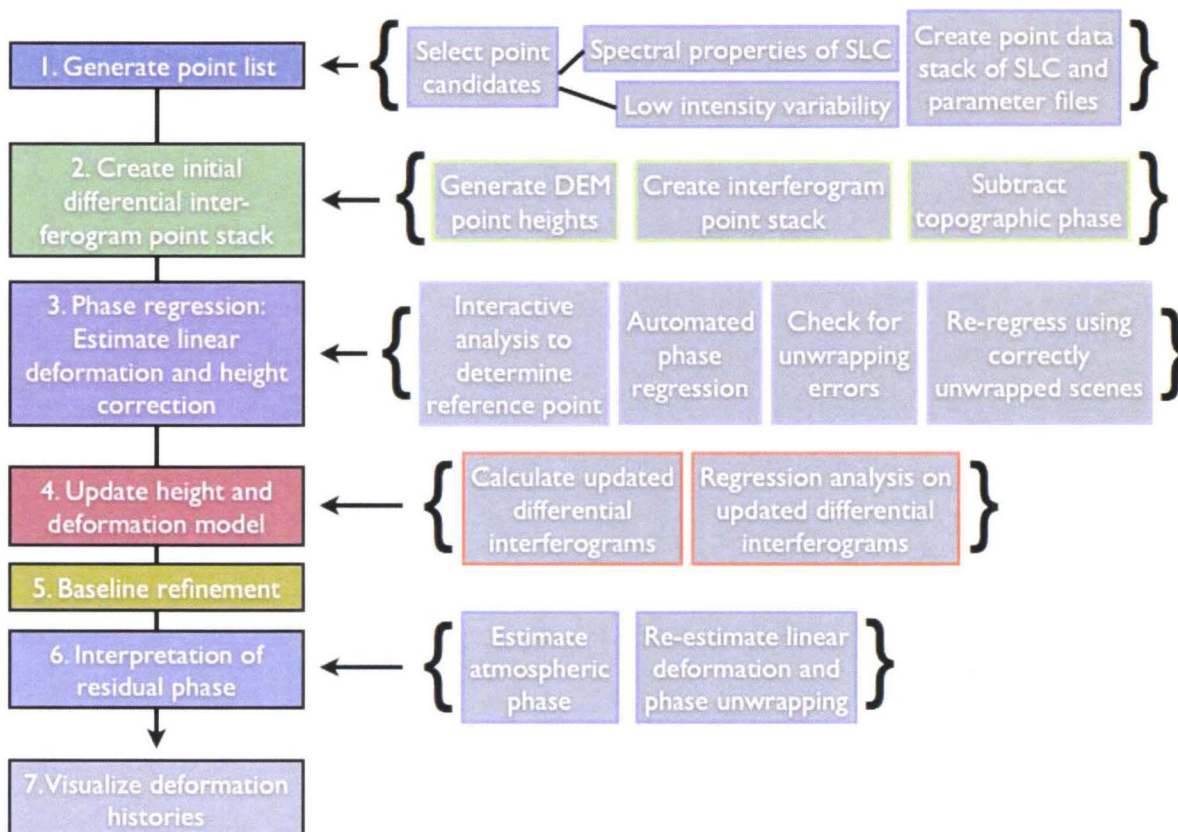
Figure 22: Displacement in UTM averaging motion between 20030212 and 20060308. One color-cycle corresponds to 0.01 m. Raster generated using GAMMA command rasrmg with a phase display scale factor of 200π . North is up.

IPTA – [Interferometric Point Target Analysis](#)

Please also refer to the [GAMMA IPTA Processing Example Luxembourg](#)

The IPTA is an alternative approach to the traditional interferometric spatial processing. The phases of points rather than from the entire interferogram are going to be interpreted. The chapter will go through point candidate selection algorithms, interferogram generation, simulation of topographic phase and regression analyses to model unwrapped phase.

The following flow chart describes the basic processing flow where each step is going to be illuminated in greater detail throughout this chapter. Processing scripts consist of a series of different commands and sometimes can be run in various modes, utilizing a different sequence of commands. Text displayed in the font `courier new` refers to either the content of a file or the script/command that is run and their standard output. The input arguments are separated by spaces. For an explanation of the input/output parameters please refer to **Appendix E** (Index of commands/scripts and their syntax). A command history can be found in the [README](#) file and in the example of the MSP application in **Appendix H**.



One of the biggest limitations in utilizing SAR interferometry for surface deformation mapping is decorrelation in the temporal and spatial domain. Temporal decorrelation is due to changes in the characteristics of reflectors on the ground, for example caused by changing vegetation and/or low temporal sampling frequency.

Spatial or geometric decorrelation is usually related to long baselines. Interferometric point target analysis (IPTA) is used as an alternative approach to monitor crustal deformation by exploiting the interferometric signatures from pixels on the ground that behave somewhat like point scatterers [Werner, et al., 2003], also referred to as point targets (GAMMA), permanent scatterers [Ferretti, et al., 2001] or persistent scatterers [Hooper, et al., 2004]. Scatterers with point-like characteristics are stable natural reflectors which are coherent over long time intervals. By identifying point targets in low-coherence areas, spatial gaps in deformation maps may be filled and the temporal sampling is improved by inclusion of large baselines pairs [Wiesmann, 2004].

IPTA-1 Focus on Smaller Area of Interest (optional)

Depending on the nature of the research project, the processing region may be cut down to a smaller area (Figure 23: [New area](#) PDF 2.6 MB). This has the advantage of quicker processing and smaller file sizes. The new area dimensions can be obtained by viewing one of the differential interferograms in Image processing software Gimp. It is important to determine the range and azimuth offset to the first sample of the new region as well as the desired width and length. Caution: the azimuth values need to be corrected for the MLR (1:5). That is, if the new area starts at azimuth line 2300 as determined in the multi-looked differential interferogram image, the corresponding RSLC azimuth line is 11500. The cropped files will be saved in a different folder (in this case *_kapoho).

IPTA-1.1 Crop the RSLC to Desired Region

Copy RSLC with options for data format conversion, segment extraction, and swapping

Note the change in width and height: [SLC par in](#) → [SLC par out](#)

```
SLC_copy rslc/20030212.rslc rslc/20030212.rslc.par rslc_kapoho/20030212.rslc
rslc_kapoho/20030212.rslc.par 1 - 1912 500 11500 2600
```

IPTA-1.2 Create new [RSLC tab](#)

Change the [existing RSLC tab](#) in order to include the new path: [RSLC tab kapoho](#)

IPTA-1.3 Generate new RMLI, Average RMLI and [Parameter File](#) (Figure 24: [Cropped MLI](#) PDF 930 KB)

```
mk_mli_all RSLC_tab_kapoho rml_i_5_kapoho 1 5 1
cp rml_i_5_kapoho/20041208.rml_i.par rml_i_5_kapoho/rml_i_5.ave.par
```

IPTA-1.4 Crop DEM (MLI format) to Desired Region ([MLI out par](#))

```
MLI_copy geo/BI_dem.rdc rml_i_5/rml_i_5.ave.par geo/BI_dem_kapoho.rdc
geo/BI_dem_kapoho.rdc.par 1912 500 11500 2600
```

- Generate raster image of DEM as shaded relief

```
rasshd geo/BI_dem_kapoho.rdc 500 15 15 1 0 1 1
```

- Display of Sun raster images (Figure 25: [Cropped DEM in RDC](#) PDF 1 MB)

```
disras geo/BI_dem_kapoho.rdc.ras&
```

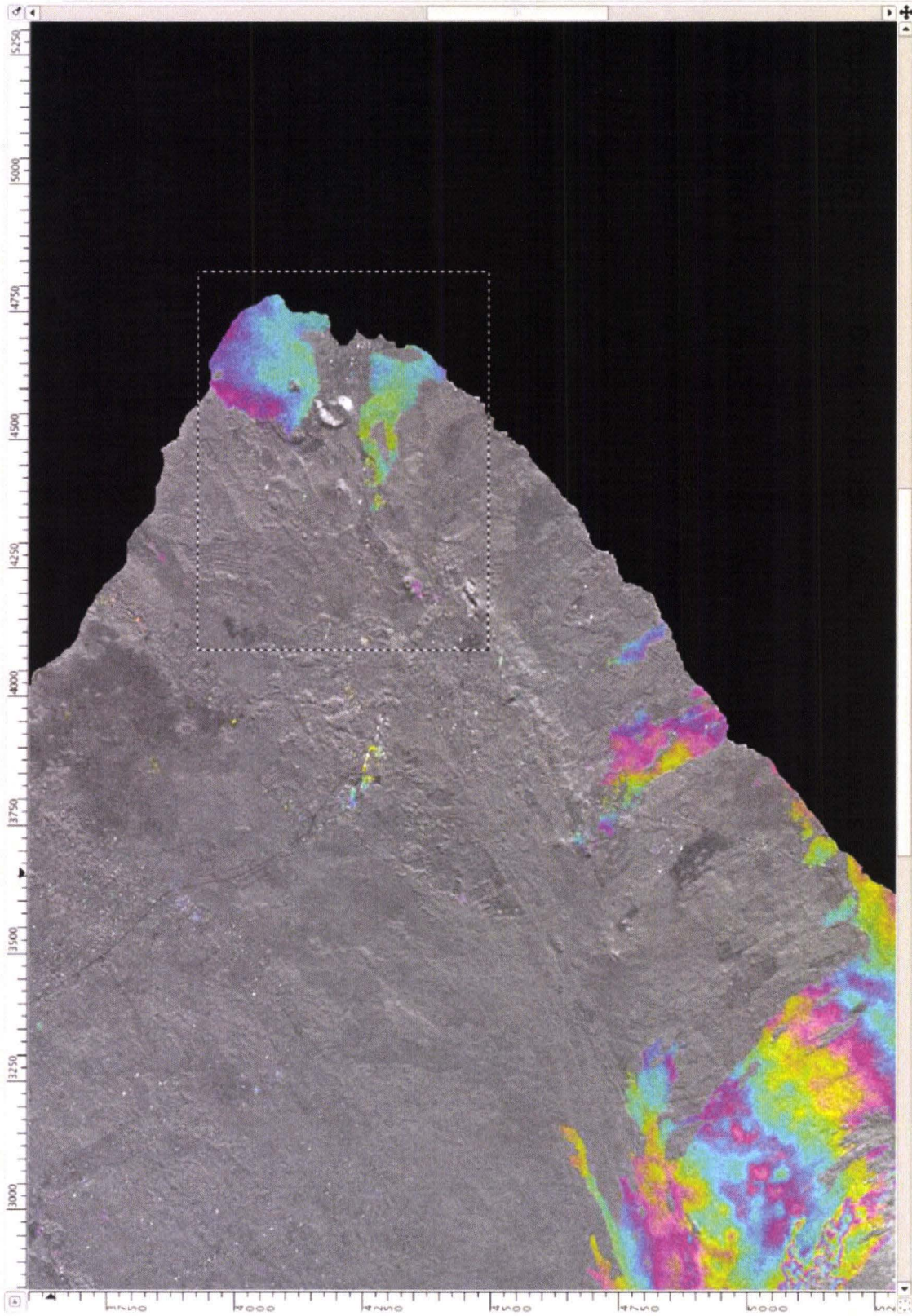



Figure 23: Defining the dimension of a smaller area using the image processing software Gimp. North is up.

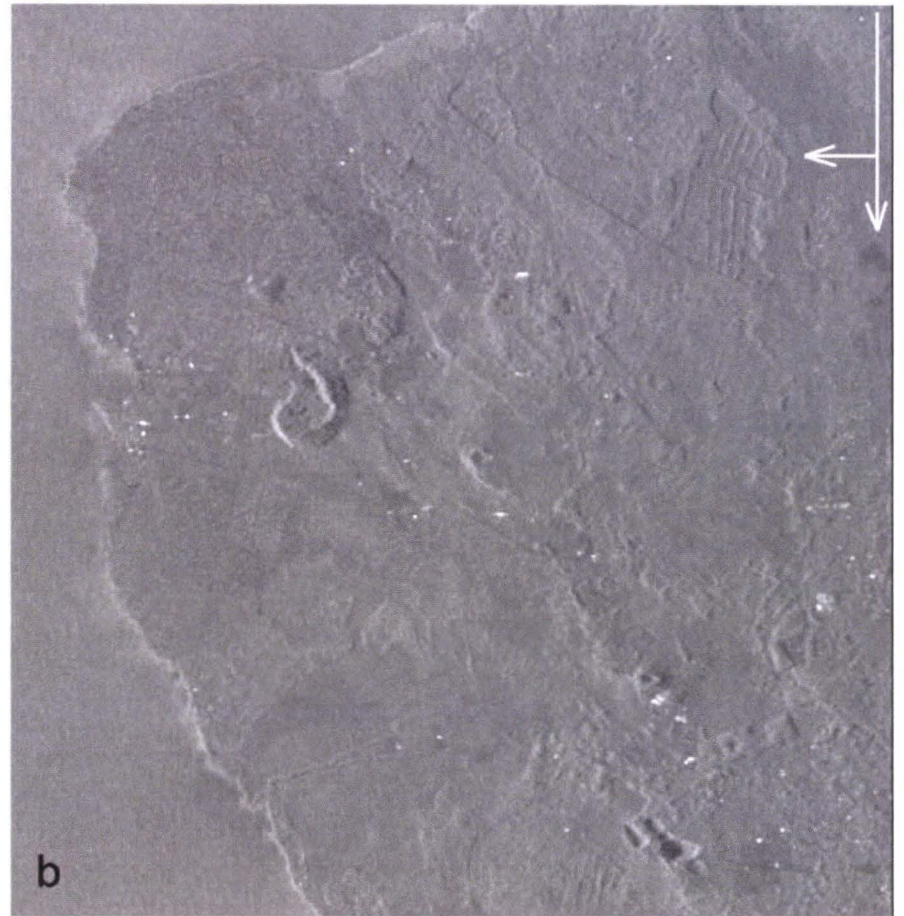
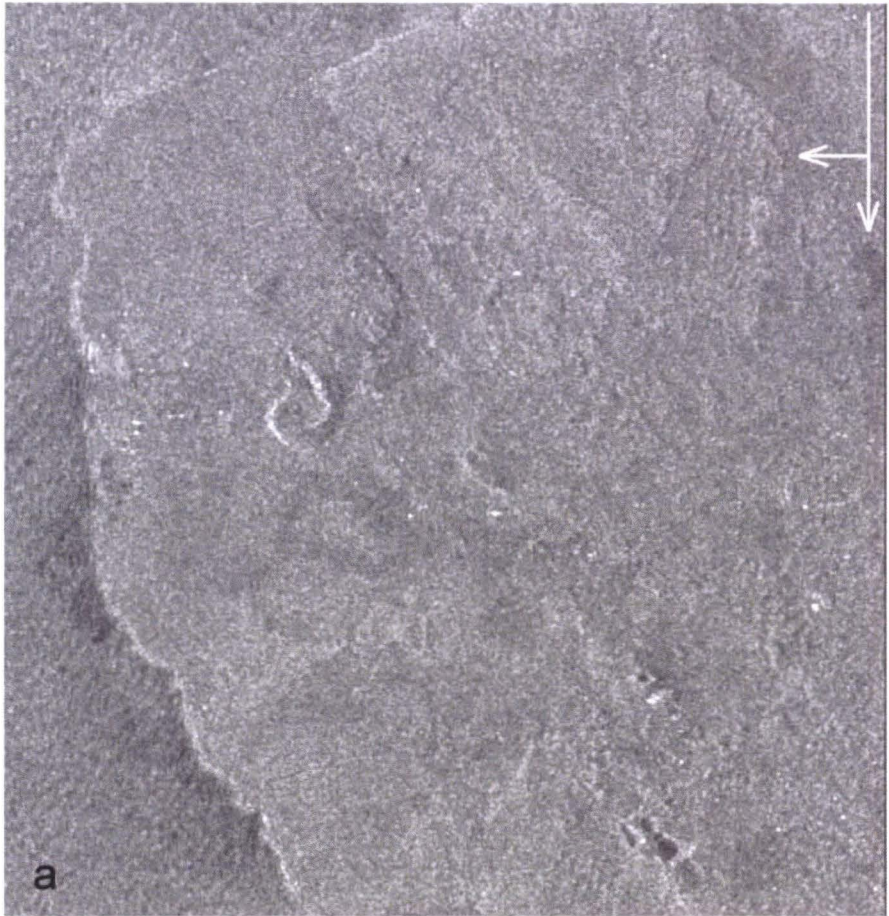


Figure 24: Multi-Look Intensity (MLI) images of the cropped (a) co-registered Single-Look Complex (RSLC) Image and (b) the average image. Arrows indicate satellite's flight- and the SAR's look-direction.

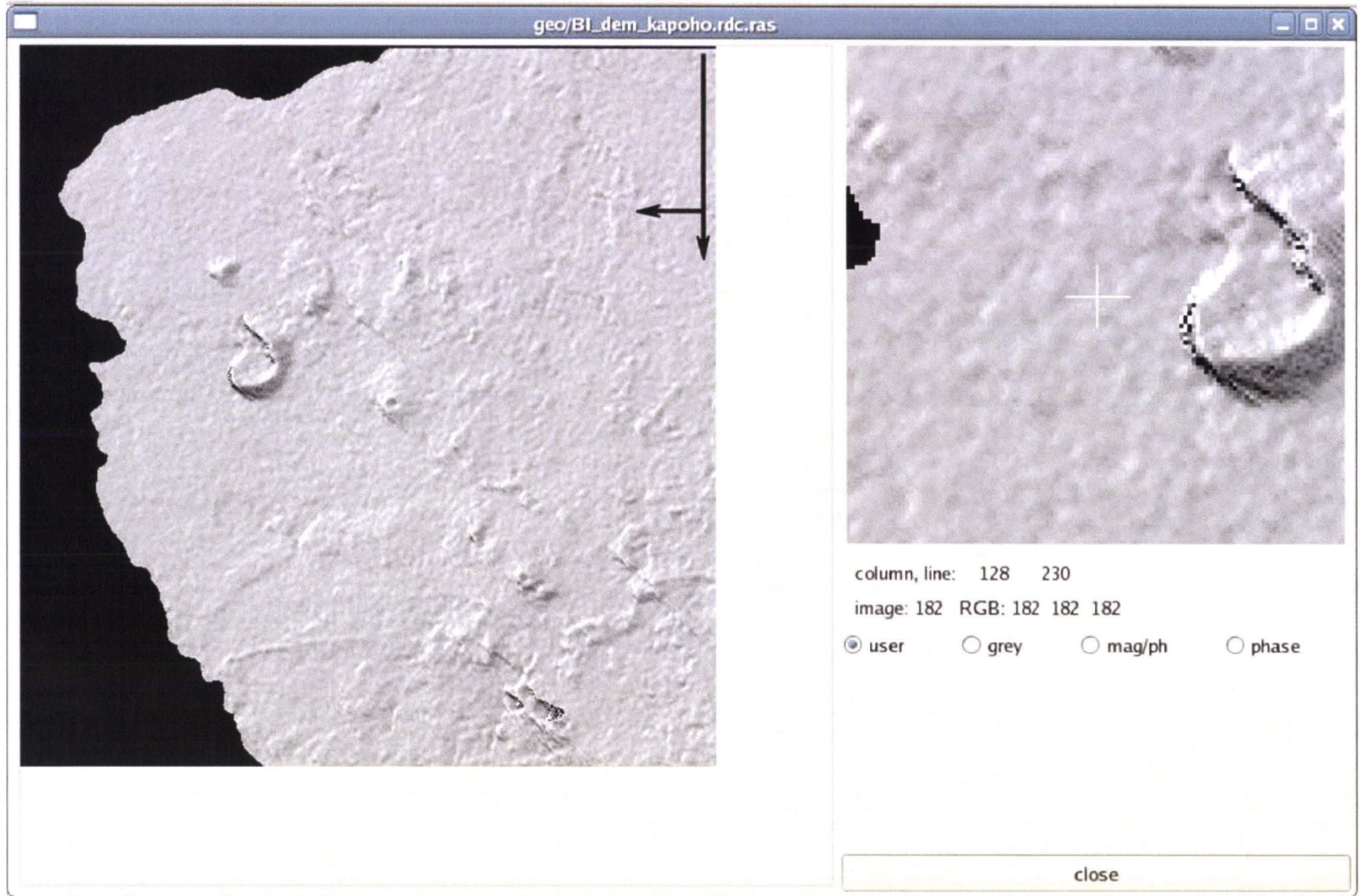


Figure 25: Cropped Digital Elevation Model (DEM) in Range-Doppler Coordinates (RDC) as displayed using GAMMA command disras. Arrows indicate satellite's flight- and the SAR's look-direction.

IPTA-2 Create [IPTA Interferogram Table](#)

- Use command **base_calc** to create a single-reference table [itab_single](#) (click [here](#) to see the full std out)

A single-reference itab will only create pairs of the reference scene with every available scene, including the reference scene itself. In contrast, an itab with all possible combinations will include all possible pairs.

```
base_calc RSLC_tab_kapoho rslc_kapoho/20041208.rslc.par 20041208.bperp.gr 20041208.bperp
itab_single 0
```

```
number of baselines in the ITAB with length > 0: 20
average bperp magnitude (m): 391.634
average bperp (m): 141.368
```

IPTA-3 Generate [Point List](#)

IPTA conducts interferometric analysis on selected points, defined in a point list by range and azimuth coordinates. Candidates are selected as point target based on their stable and point-like scattering characteristics. Point targets have no temporal and low to no geometric decorrelation, which enables inclusion of pairs with very large baselines. Two methods are utilized to select candidate points based on (1) spectral diversity and (2) temporal variability. The first approach considers the spectral characteristics of the co-registered SLC. Low spectral diversity, i.e., high spectral correlation is the selection criteria for this method. The second approach is based on point targets not showing speckle behavior and therefore having low temporal intensity variability (high Mean to Sigma Ratio MSR). The point candidate lists generated by these two methods are then merged to a single point list. Areas of no interest or over water are masked out.

IPTA-3.1 Generation of Point Target Candidate List Based on Spectral Properties of Individual SLC

```
mk_sp_all RSLC_tab_kapoho sp 4 4 0.5 0.4 1.2
```

The following commands are run for the entire list of RSLC specified in the RSLC_tab:

sp_stat, ave_image, ras_linear, ave_image, ras_linear

- Generate point list using low spectral diversity to identify point targets

```
sp_stat rslc_kapoho/20030212.rslc - sp/20030212.sp_cc sp/20030212.sp_msr sp/20030212.pt_ext
500 0.5 0.4 1.2 4 4 - - - - - 0
```

- Calculate average of spectral correlation (using the [MLI tab](#) as input list)

```
ave_image sp/cc_list 500 sp/ave.sp_cc
```

- Generate raster image of spectral correlation using a linear scale

```
ras_linear sp/ave.sp_cc 500 1 0 1 1 0. 1.
```

- Calculate average of mean/sigma ratio

```
ave_image sp/msr_list 500 sp/ave.sp_msr
```

- Generate raster image of mean/sigma ratio using a linear scale

```
ras_linear sp/ave.sp_msr 500 1 0 1 1 0. 5.
```

IPTA-3.2 Use Spectral Correlation Average and Different Thresholds to Create Point List

- Generate point list containing coordinates of image points that satisfy constraints regarding the selected thresholds

```
thres_im_pt sp/ave.sp_cc 500 sp/pt_cc_3 0.3 - 1 1
number of points that satisfy threshold constraints: 49209
```

```
thres_im_pt sp/ave.sp_cc 500 sp/pt_cc_318 0.318 - 1 1
number of points that satisfy threshold constraints: 21340
```

```
thres_im_pt sp/ave.sp_cc 500 sp/pt_cc_35 0.35 - 1 1
number of points that satisfy threshold constraints: 5828
```

- Draw point list locations on a SUN/BMP raster image (Figure 27: [Spectral correlation points](#) PDF 3 MB)

```
ras_pt sp/pt_cc_3 - rml1_1_5_kapoho/rml1_1_5.ave.ras sp/pt_cc_3.ras 1 5
```

```
ras_pt sp/pt_cc_318 - rml1_1_5_kapoho/rml1_1_5.ave.ras sp/pt_cc_318.ras 1 5
```

```
ras_pt sp/pt_cc_35 - rml1_1_5_kapoho/rml1_1_5.ave.ras sp/pt_cc_35.ras 1 5
```

IPTA-3.3 Generation of Point Target Candidate List Based on Low Intensity Variability

- Generate a set of point target lists with different thresholds (determined based on predetermined values by GAMMA)

```
mk_msr_pt RSLC_tab_kapoho rslc_kapoho/20040303.rslc.par rml1_1_5_kapoho/rml1_1_5.ave.par
rml1_1_5_kapoho/rml1_1_5.ave.ras msr .5 2 1.5 1 1.2 10 .1
```

The script [mk_msr_pt](#) runs the following commands: **pwr_stat**, **ras_pt**, **npt**

- Generate point list based on temporal variability of SLC intensity

```
pwr_stat RSLC_tab rslc/20040303.rslc.par msr/msr_1.2 msr/pt_1.2 1.2 .5 0 0 - - 0 2 1.5 1
```

- Draw point list locations on a raster image (Figure 28: [Low intensity variability points](#) PDF 3.8 MB)

```
ras_pt msr/pt_1.2 - rml1_1_5_kapoho/rml1_1_5.ave.ras msr/pt_1.2.ras 1 5 255 255 0
```

```
ras_pt msr/pt_1.4 - rml1_1_5_kapoho/rml1_1_5.ave.ras msr/pt_1.4.ras 1 5 255 255 0
```

```
ras_pt msr/pt_1.7 - rml1_1_5_kapoho/rml1_1_5.ave.ras msr/pt_1.7.ras 1 5 255 255 0
```

```
ras_pt msr/pt_2.1 - rml1_1_5_kapoho/rml1_1_5.ave.ras msr/pt_2.1.ras 1 5 255 255 0
```

- Determine number of points in a point list

```
npt msr/pt_1.2
```

IPTA-3.4 Merging of Point Target Candidate Lists to Create the Point List Pt

- Create a [list](#) that contains the chosen cc and msr list

```
ls -l sp/pt_cc_318 msr/pt_1.7 > plist_tab
```

- Merge the two point lists

```
merge_pt plist_tab pt 1 0 0
```

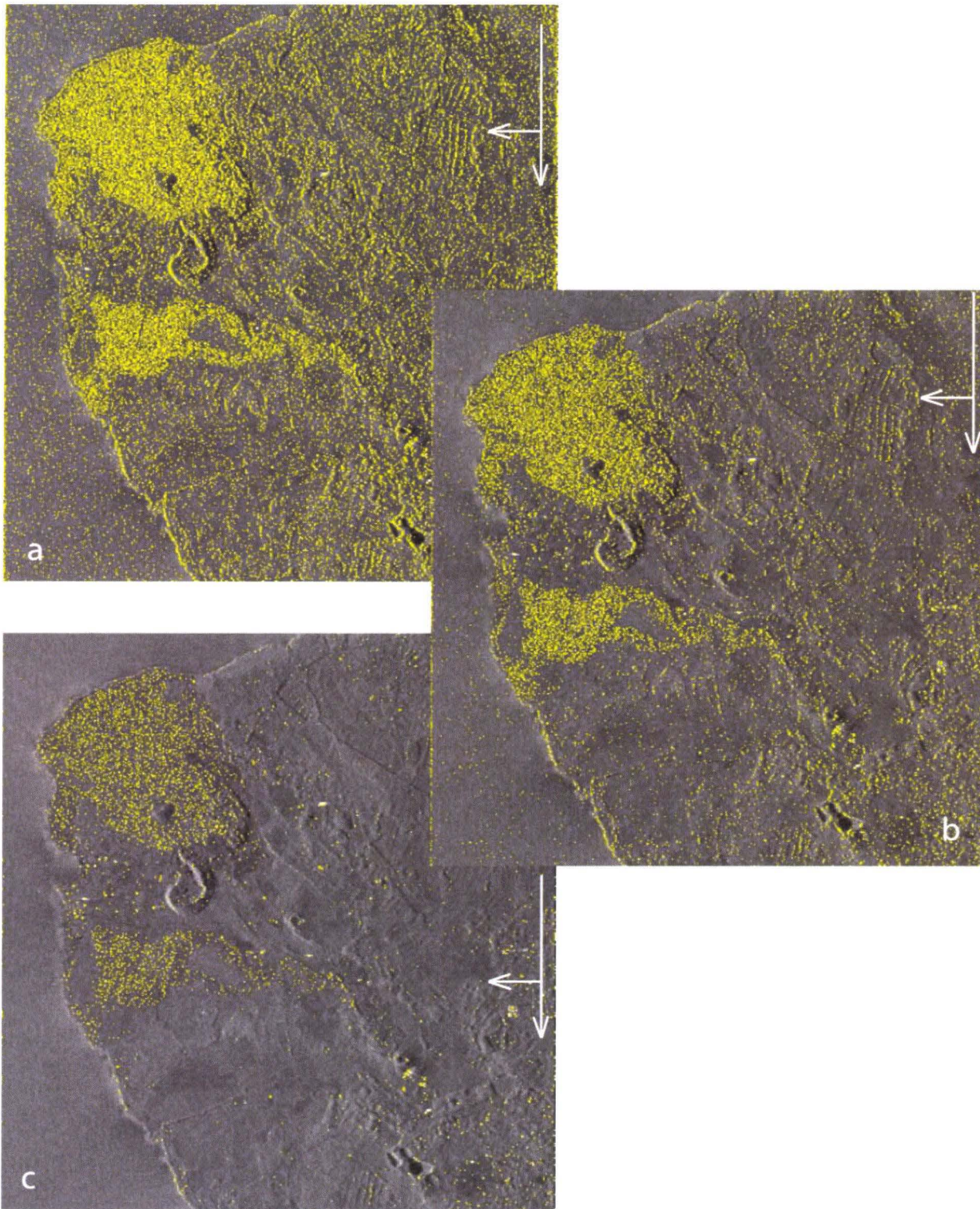


Figure 27: Point locations (shown in yellow dots) of three point lists created using spectral correlation thresholds of (a) 0.3, (b) 0.318, and (c) 0.35. The backscatter intensity image is in the background. Arrows indicate satellite's flight- and the SAR's look-direction.

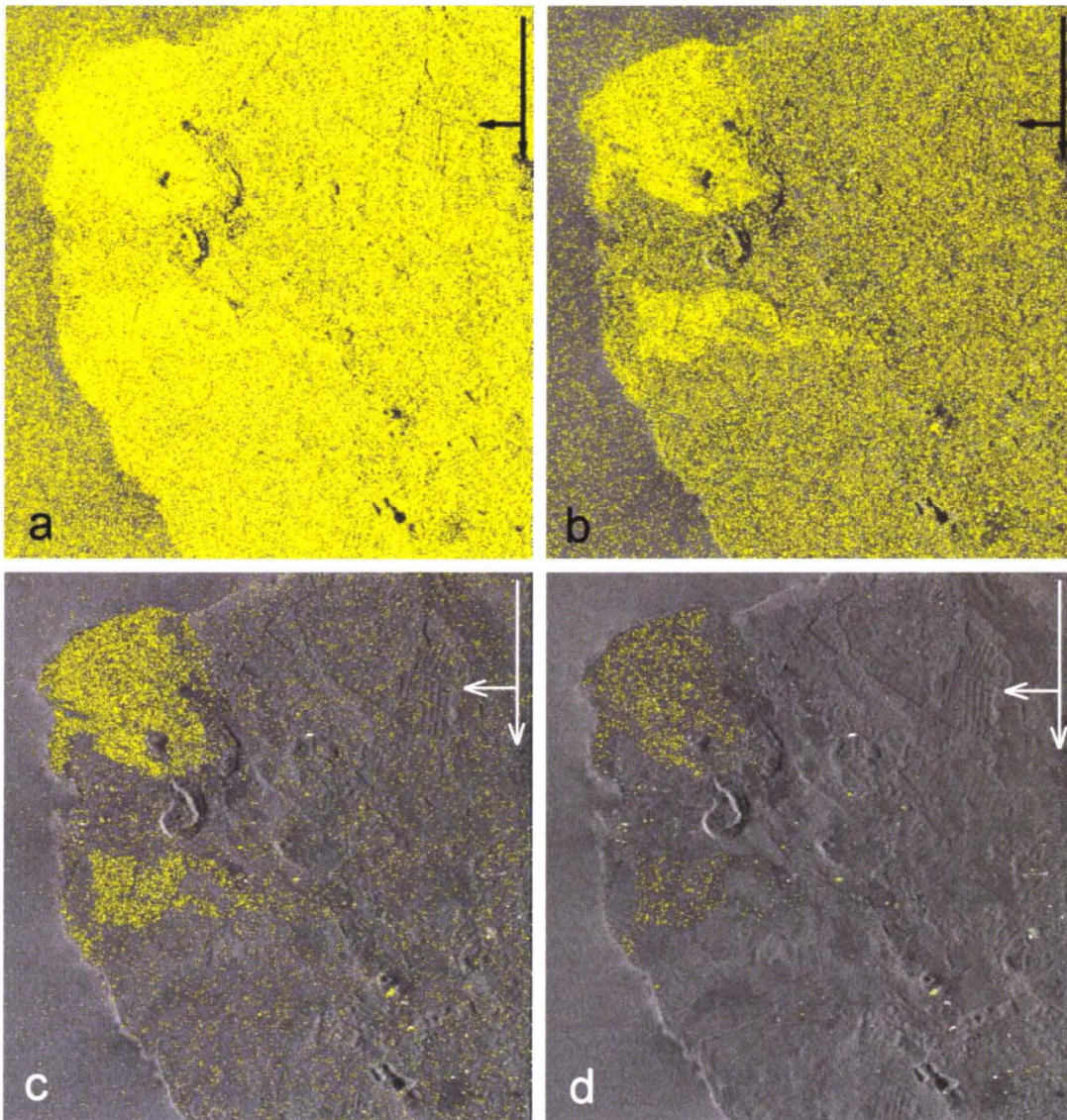


Figure 28: Point locations (yellow dots) of four point lists created temporal variability of SLC intensity thresholds of (a) 1.2, (b) 1.4, (c) 1.7 and (d) 2.1. Backscatter intensity image is in the baackground. Arrows indicate satellite's flight- and the SAR's look-direction.

- Create a raster image of the merged points (Figure 29a: [Merged points](#) PDF 2.2 MB)

```
ras_pt pt - rml1_1_5_kapoho/rml1_1_5.ave.ras pt.ras 1 5 255 0 0 3
```

IPTA-3.5 Create a Mask Raster File

Use Gimp or equivalent software such as Photoshop, Xv etc. to color all the areas that are not of interest black.

Zero values are not considered in the processing.

- Generate a point list and mask vector using a raster image mask

```
msk_pt pt - rml1_1_5_kapoho/mask.ras pt_masked pmask 1 5
```

- Save original point list without mask and rename masked point list to *pt*

```
cp pt pt_nomask
mv pt_masked pt
```

- Create raster of masked points (Figure 29b: [Masked points](#))

```
ras_pt pt - rml1_1_5_kapoho/rml1_1_5.ave.ras pt.ras 1 5
```

IPTA-4 Generate pSLC_par and SLC Point Data Stacks pSLC ([log](#))

```
SLC2pt RSLC_tab_kapoho pt - pSLC_par pSLC -
```

The script runs the following commands for the entire list of RSLC specified in the RSLC_tab:

SLC_par_pt, data2pt

- Read/write record of SLC/MLI parameter stack

```
SLC_par_pt rslc_kapoho/20030212.rslc.par pSLC_par 1 1
```

- Raster to vector format conversion

Extract values from a 2-D data file at specified point locations

```
data2pt rslc_kapoho/20030212.rslc rslc_kapoho/20030212.rslc.par pt
rslc_kapoho/20030212.rslc.par pSLC 1 0
```

IPTA-5 Generate Initial Differential Interferograms pdiff0

The initial differential interferograms do not include a height corrected DEM or a linear deformation model.

IPTA-5.1 Generate DEM Point Data File pdem

- Convert raster DEM to vector DEM

```
data2pt geo/BI_dem_kapoho.rdc geo/BI_dem_kapoho.rdc.par pt rslc_kapoho/20041208.rslc.par pdem2
1 2
```

- Display point height data

```
pdisdt_pwr24 pt - rslc_kapoho/20041208.rslc.par pdem2 1 rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 256. 1
```


The script runs the following commands: `pt2data, disdt_pwr24`

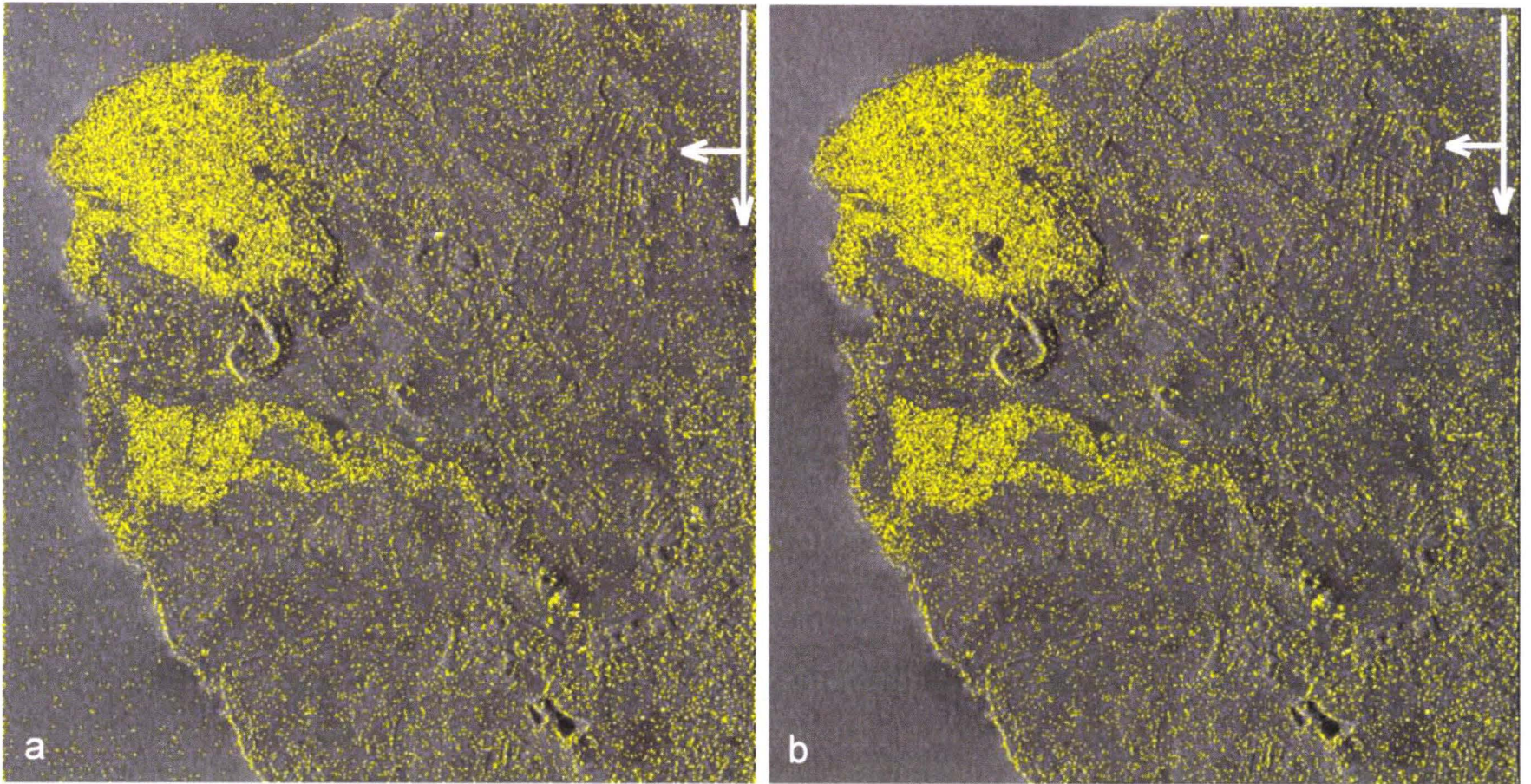


Figure 29: Raster images of (a) merged point list and (b) masked point list. The mask has zero values (black) in areas of water, for this example; other low correlation areas or regions of no interest can also be masked out. Points are shown as yellow dots. Backscatter intensity image is in the background. Arrows indicate satellite's flight- and the SAR's look-direction.

- Vector to raster format conversion

Interpolate point data values into a 2-D data file. Uses SLC/MLI parameter file as input.

```
pt2data pt - pSLC_par pdem2 1 tmp.pdisdt_pwr24 rml_i_1_5_kapoho/rml_i_1_5.ave.par 2 3 4 1
```

- Display float parameter (e.g., height) and intensity image (Figure 30a: [Point heights](#) PDF 2.6 MB)

```
disdt_pwr24 tmp.pdisdt_pwr24 rml_i_1_5_kapoho/rml_i_1_5.ave 500 1 1 0 256. 1. .4
```

IPTA-5.2 Generate Interferogram Point Data Stack pint0

Calculate the initial interferograms and display the individual layers

IPTA-5.2.1 Create Interferogram and Baseline Data Stacks ([log](#))

```
mk_int_all pt - pSLC_par itab_single pSLC pbase pint mk_int_all.log
```

The script [mk_int_all](#) runs the following commands: **SLC_par_pt**, **base_orbit_pt**, **intf_pt**

- Generate stack of parameter files

```
SLC_par_pt /tmp/slc.par pSLC_par 1 0
```

- Estimate baseline from orbit state vectors (single record or all records)

```
base_orbit_pt pSLC_par itab_single - pbase
```

- Generate interferogram from SLC point data stack (single record or all records)

```
intf_pt pt - itab_single - pSLC pint 0
```

IPTA-5.2.2 Display the Layers of pint0

Generate 2-D images of point data using **pt2data**, **rasdt_pwr24** and **rasmph_pwr24** ([log](#))

```
mk_2d_im pt - itab_single rslc_kapoho/20041208.rslc.par pint 1 - rml_i_1_5_kapoho/rml_i_1_5.ave.par 0 3 1.2 1 int0 rml_i_1_5_kapoho/rml_i_1_5.ave - 0
```

The script runs the following commands for each record in the stack:

pt2data, **rasmph_pwr24** (fcomplex) or **rasdt_pwr24** (float)

- Convert vector data to raster format for each record

```
pt2data pt - rslc_kapoho/20041208.rslc.par pint 20 int0/pint_020 rml_i_1_5_kapoho/rml_i_1_5.ave.par 0 3 1.2 1
```

Generate 24-bit SUN/BMP raster file of interferogram phase and intensity image (Figure 31: [Pint0](#) PDF 5 MB)

```
rasmph_pwr24 int0/pint_020 rml_i_1_5_kapoho/rml_i_1_5.ave 500 1 1 0 1 1 .9
```

IPTA-5.3 Calculate and Subtract Topographic Phase Model psim_unw

Linear deformation estimates, height corrections, refined baselines and atmospheric distortions are **not** available yet.

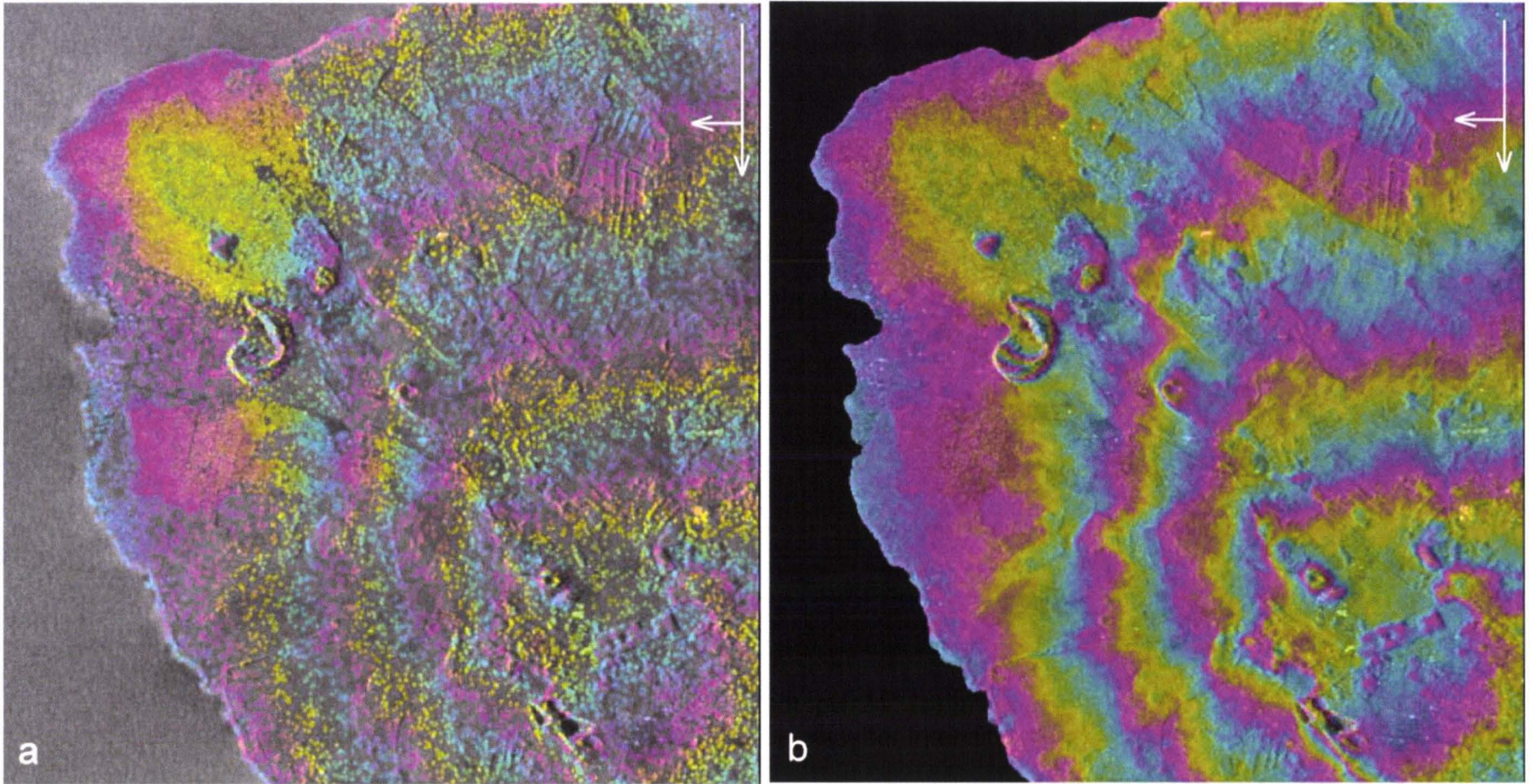


Figure 30: (a) Display of point heights using GAMMA script `pdisdt_pwr24`. (b) In comparison the DEM heights in RDC generated using `rashgt`. One color-cycle corresponds to 50 m. Backscatter intensity image is in the background. Arrows indicate satellite's flight- and the SAR's look-direction.

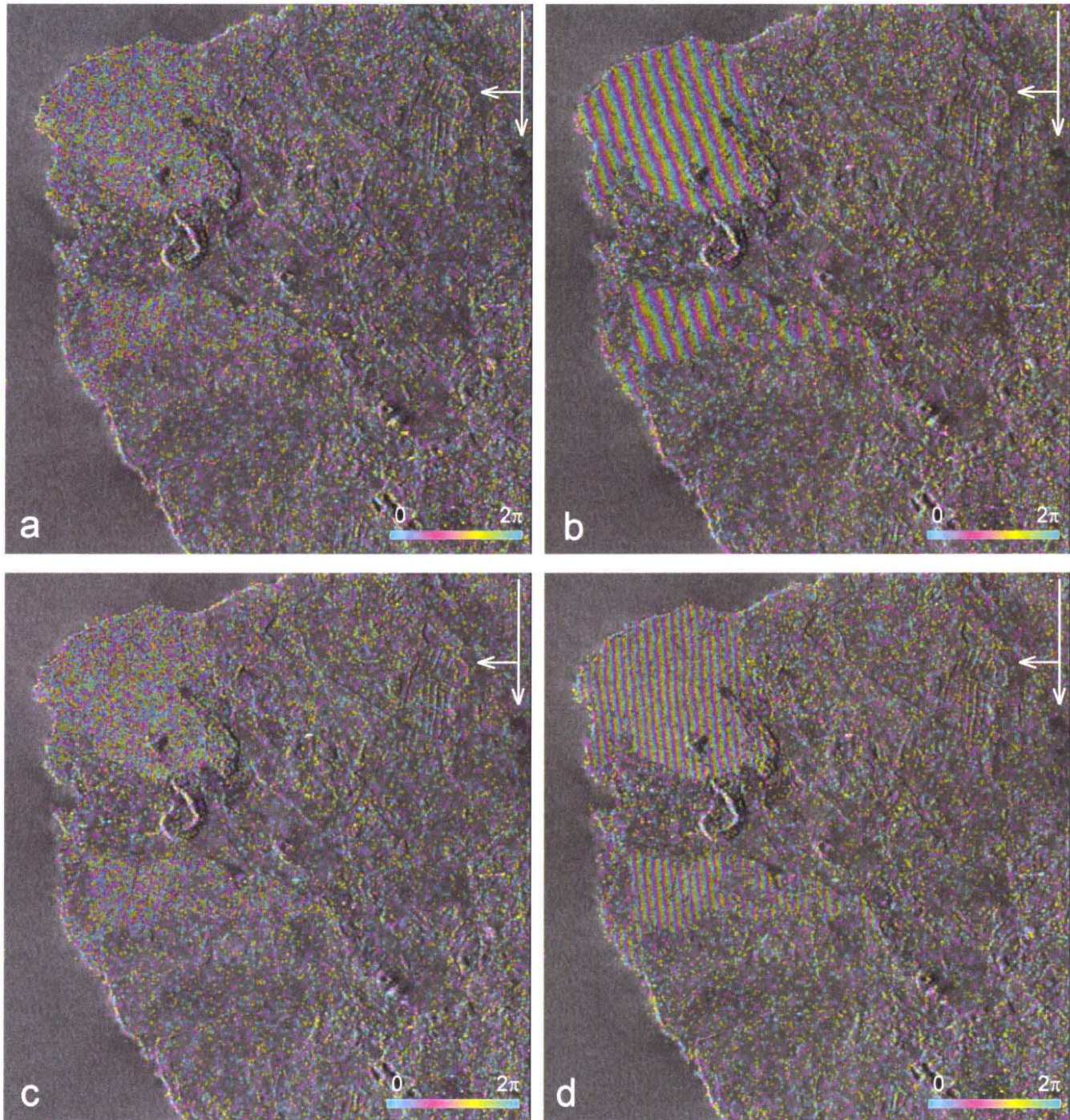


Figure 31: Four layers of the initial wrapped interferometric point data stack pint0: (a) pint0_001, (b) pint0_002, (c) pint0_004, (d) pint0_008. Generated using GAMMA script mk_2d_im. One color-cycle corresponds to 2π radians. Arrows indicate satellite's flight- and the SAR's look-direction.

IPTA-5.3.1 Simulate Topographic Phase and Subtract it from Interferogram ([log](#))

```
mk_diff_all pt - pSLC_par itab_single pbase 0 pint 1 pdem2 - - psim_unw0 pdiff0
mk_diff_all0.log
```

The script [mk_diff_all](#) first deletes the simulated phase if it already exists from previous processing.

```
rm -rf psim_unw0
```

Then it runs the following commands: **phase_sim_pt**, **sub_phase_pt**

- Simulate unwrapped interferometric phase

```
phase_sim_pt pt - pSLC_par - itab_single - pbase pdem2 psim_unw0 - 0 0
```

The script [mk_diff_all](#) then deletes already existing differential interferogram stacks (remnants from earlier processing).

```
rm -rf pdiff0
```

- Subtract phase from interferograms

```
sub_phase_pt pt - pint - psim_unw0 pdiff0 1
```

It is recommended to save the initial differential interferograms, because in future processing the *pdiffs* are going to be overwritten.

```
cp pdiff0 pdiff0_no_phgt
cp -r diff0 diff0_no_phgt
```

IPTA-5.3.2 Create Raster Images of Differential Interferograms (Figure 32: [Diff interferograms](#) PDF 5 MB)

```
mk_2d_im pt - itab_single rslc_kapoho/20041208.rslc.par pdiff0 1 - rml1_1_5_kapoho/rml1_1_5.ave.par 0 3 1.2 1 diff0 rml1_1_5_kapoho/rml1_1_5.ave - 0
```

IPTA-6 Phase Regression Analysis for Pairs of Points

The general phase model for IPTA is identical to conventional interferometry. The unwrapped interferometric phase may be expressed as the sum of topographic phase, deformation phase, atmospheric phase (also referred to as differential path delay phase) and phase noise (or decorrelation) terms [*Werner, et al.*, 2003]. IPTA is a method to utilize the spatial and temporal characteristics of these phase terms (see table below) to precisely calculate height maps, deformation histories and the atmospheric path delay. The simulated (modeled) unwrapped topographic phase and its model parameters are iteratively improved throughout the process, which also includes a baseline refinement.

The phase model point data stack (simulated topography) may be calculated with respect to a linear deformation rate estimate and/or height corrections for the available DEM. At first, though, no deformation and no height corrections are available. The initial phase model is subtracted from the complex valued wrapped point interferograms to obtain complex valued wrapped point differential interferograms.

The next processing step consists of least-squares regression analyses performed on the stack of differential interferograms. The interferometric phase model suggests linear phase dependence for the perpendicular baseline as well as for time.

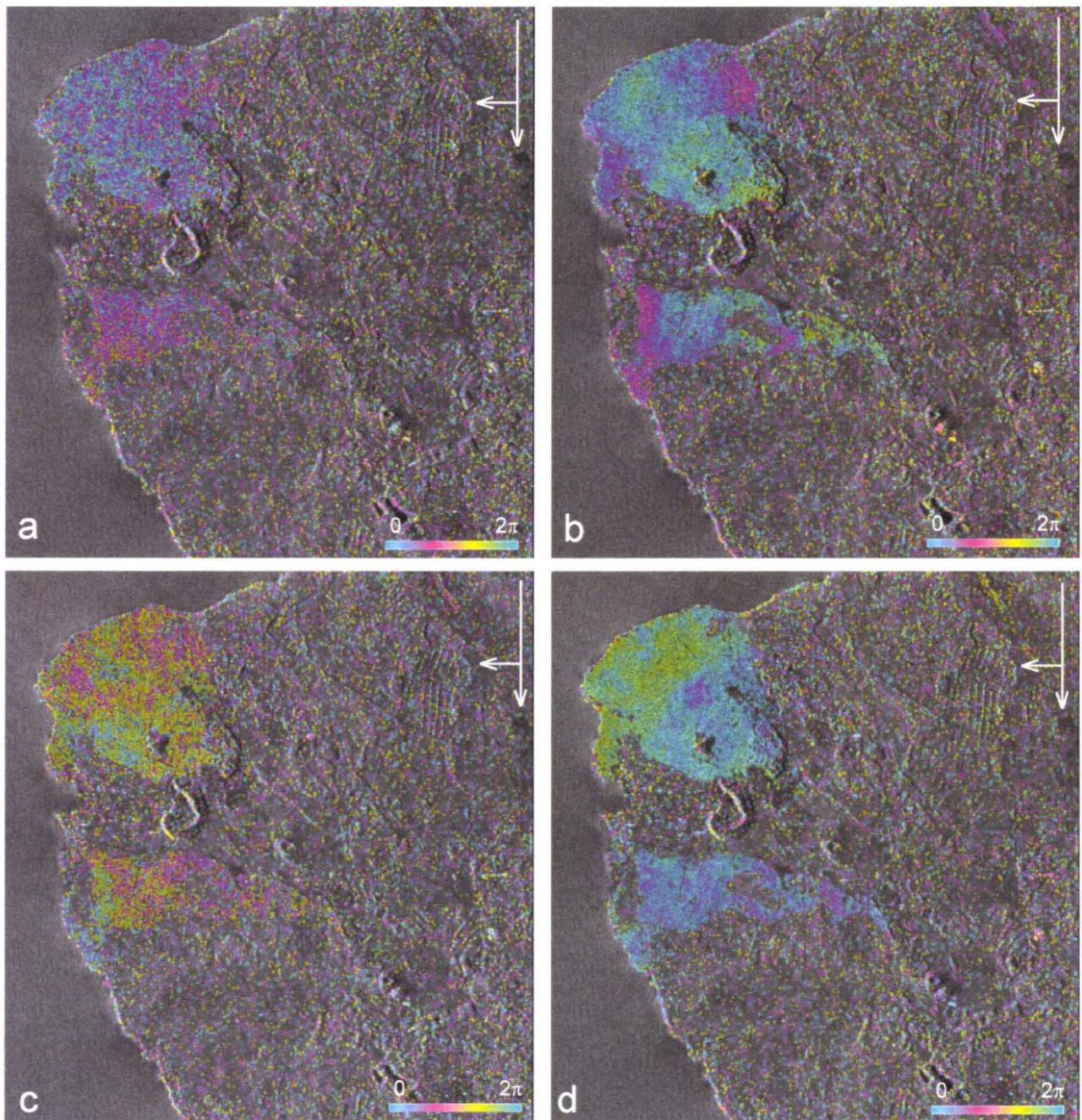


Figure 32: Four layers of the initial wrapped differential interferometric point data stack pdiff0: (a) pdiff0_001, (b) pdiff0_002, (c) pdiff0_004, (d) pdiff0_008. Raster images created using GAMMA script mk_2d_im. One color-cycle corresponds to 2π radians. Arrows indicate satellite's flight- and the SAR's look-direction.

The slopes of the regression represent height corrections and linear deformation rate, respectively. Those terms may then be used to refine the model of the simulated unwrapped topographic phase. The regression analysis is primarily done across the data stack, i.e., in the temporal domain. Additional output parameters are the unwrapped phase, a quality measure and residual phases. The table below shows the spatial and temporal characteristics of the components of the residual phase which have to be utilized in order to be able to distinguish between atmospheric and non-linear deformation phase.

Residual phase components	Spatial domain	Temporal domain
Atmosphere	Correlated/Low pass	Random/High pass
Non-linear deformation	Correlated/Low pass*	Correlated/Low pass*
Baseline errors	Correlated/Low pass	Random/High pass
Phase noise	Random/High pass	Random/High pass

*This is generally the case, non-linear deformation can also be highly uncorrelated (earthquakes, landslides etc.)

There are six models that can be applied for phase regression analysis

- 1: $a_0 + a_1 \cdot b_{\text{perp}} [i]$
- 2: $a_0 + a_1 \cdot b_{\text{perp}} [i] + a_2 \cdot \Delta_t [i]$ (default)
- 3: $a_1 \cdot b_{\text{perp}} [i]$
- 4: $a_1 \cdot b_{\text{perp}} [i] + a_2 \cdot \Delta_t [i]$
- 5: $a_0 + a_2 \cdot \Delta_t [i]$
- 6: $a_2 \cdot \Delta_t [i]$

The parameter a_0 is an optional phase constant that needs to be calculated if a single SLC is used as the reference for all interferograms. If different SLC are used as references for the interferogram pairs in the stack (**base_calc**: all combinations) the phase constant is 0.0 and does not need to be estimated. Use of the one-dimensional regression models constrains the regression and gives better results for small stacks (25+) or early on in the analysis when large height corrections (50+ m) are expected [Wegmüller, 2005].

For the Kapoho project, a two-dimensional phase model is chosen to calculate height corrections and linear deformation simultaneously. Calculating the height corrections in a one-dimensional regression analysis first would limit the number of usable scenes, considering the fact that only short time intervals can be used to prevent hidden deformation signals.

IPTA-6.1 Interactive Point-wise Phase Regression Analysis

- Display regressions of phase with respect to time interval and baseline

In the case of wrapped phase values (**dis_ipta_unw** can be used for unwrapped phases) the regression is done in two steps by the script: 1) the best regression function among a predefined set of functions is determined and 2) a least-squares analysis is conducted. The predefined regression functions for 1) are derived based on the indicated height correction and maximum deformation rate. The higher those values, the more regression models are tested, increasing the running time. Allowing a too high maximum height correction further reduces the efficiency and might result in more phase unwrapping errors [Wegmüller, 2005]. Choosing values too small

might result in failure to resolve the regression for a point with a large height correction or linear deformation rate value. Note that height corrections and linear deformation rate estimates are relative to the reference point, they are not absolute values. The regression is therefore more robust for points near the reference point.

The reference point is determined using `dis_ipta`. It should be close to the region of interest in a relatively stable area. By double clicking the right mouse button the reference point can be locked in and individual regressions can be visually evaluated by using the right mouse button once. The left mouse button changes the zoom window.

Experimenting with the maximum height correction and maximum linear deformation rate resulted in parameter values of 30 and 0.01, respectively for the example of Kapoho. The reference point (23558) for the regression analysis is chosen in the southern flow, close to the town of Kapoho (Figure 33: [Dis_ipta output PDF 3.3 MB](#)).

```
dis_ipta pt - pSLC_par - itab_single pbase 0 pdiff0 1 diff0/pdiff0_008.ras 30 0.01 2 - - 3 128
ref.:23558 x:104 y:1502 point:13003 x:149 y: 714
dphase/dBperp: 1.41411e+03 delta2 - delta1 (m): 0.01192 alpha: 7.06403e-04

look angle (deg.): 19.54137 interior angle (deg.): 2.52874 range (m): 840996.409
number of phase steps: 8 step size: 0.78540
steps nrpm: 27 radians/meter step: 1.52009e-03 steps nrpd: 6 radians/day step: 2.03933e-03
init. slope (rad./m):-7.600e-04 def rate:-2.039e-03 phase offset: -1.571 stddev: 1.23941
model 2 slope(rad/day):-1.07310e-03 slope (rad/m):-4.89132e-04 offset: -1.43635 sigma:
1.24922

delta height (m): 0.6924 estimated uncertainty (m): 0.9815
deformation rate m/year: 1.754e-03 estimated uncertainty (m/year): 1.508e-03
phase constant (rad): -1.4363 estimated uncertainty (rad): 0.3303
```

IPTA-6.2 Automated Phase Regression Analysis

For the automated regression analysis it is recommended to save the results of every run in a different folder. That makes it easier to keep track of the outputs and prevent accidentally overwriting files.

IPTA-6.2.1 Multi-patch Estimation of Linear Deformation, Height Corrections and Residual Phase

```
mkdir multi_def_pt1

multi_def_pt pt - pSLC_par - itab_single pbase 0 pdiff0 1 23558 multi_def_pt1/pres1
multi_def_pt1/pdh1 multi_def_pt1/pddef1 multi_def_pt1/punw1 multi_def_pt1/psigma1
multi_def_pt1/pmask1 30. 0.01 70 1.2 1.0 2 0 -1 -1

number of points with sigma < sigma_max = 1.200 : 19791
```

`Multi_def_pt` uses a multi-patch approach to conduct a regression analysis. Each patch has a local reference point, making the regression more robust. The unwrapping might result in errors related to patch boundaries. It might help to choose smaller/larger patches if the sample size permits it.

- Create images of output ([Figure 34 PDF 4.7 MB](#))

```
prasd_t_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pdh1 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 30.0 1
```

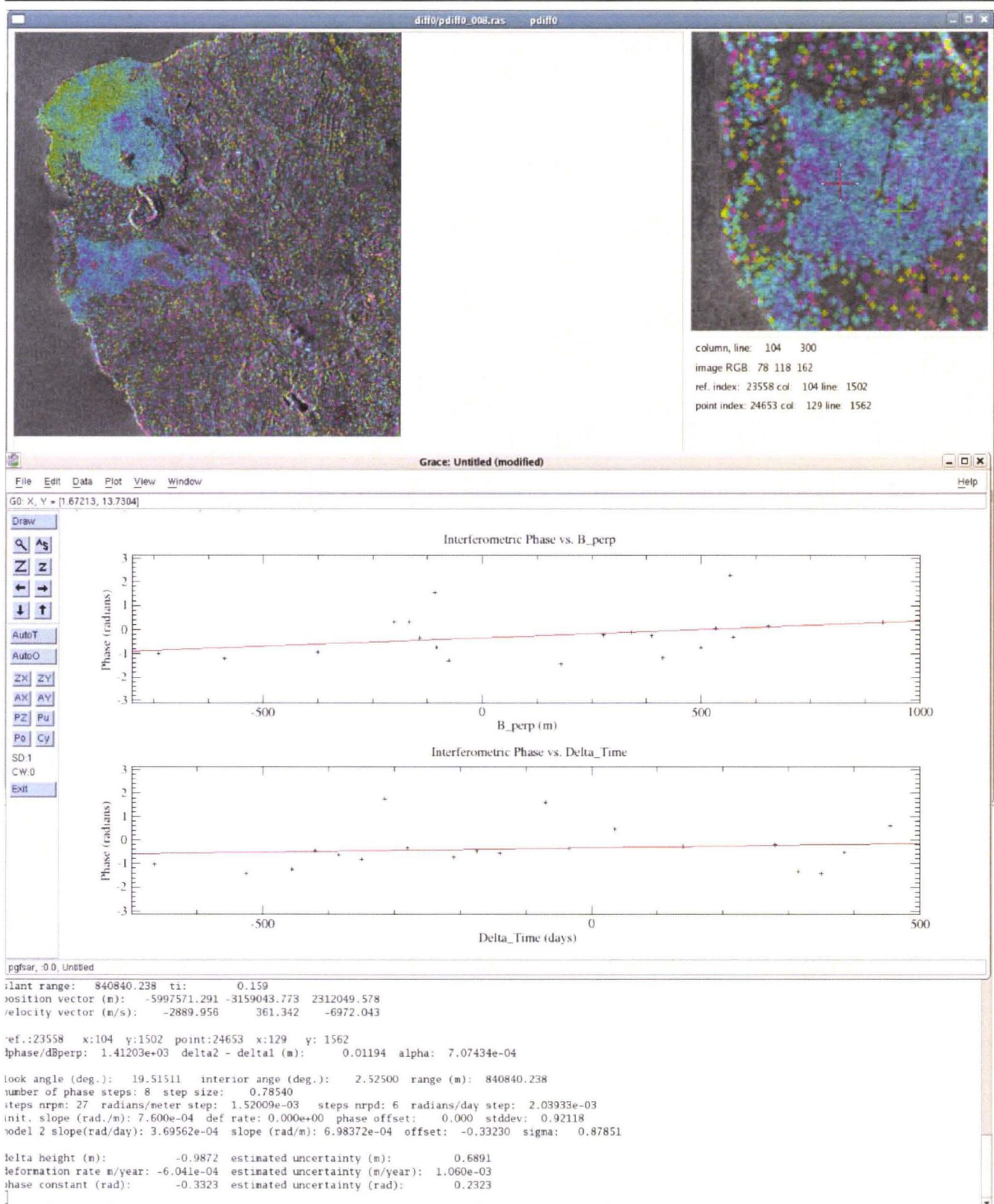



Figure 33: Conducting a manual regression to determine a reference point using GAMMA command `dis_ipta`. Location of the reference point is indicated by the red cross, the 2nd point for the regression by the green cross. The Grace plot shows the regressions, top: height corrections (Bperp vs. phase), bottom: linear deformation rate (Delta_t vs. phase).

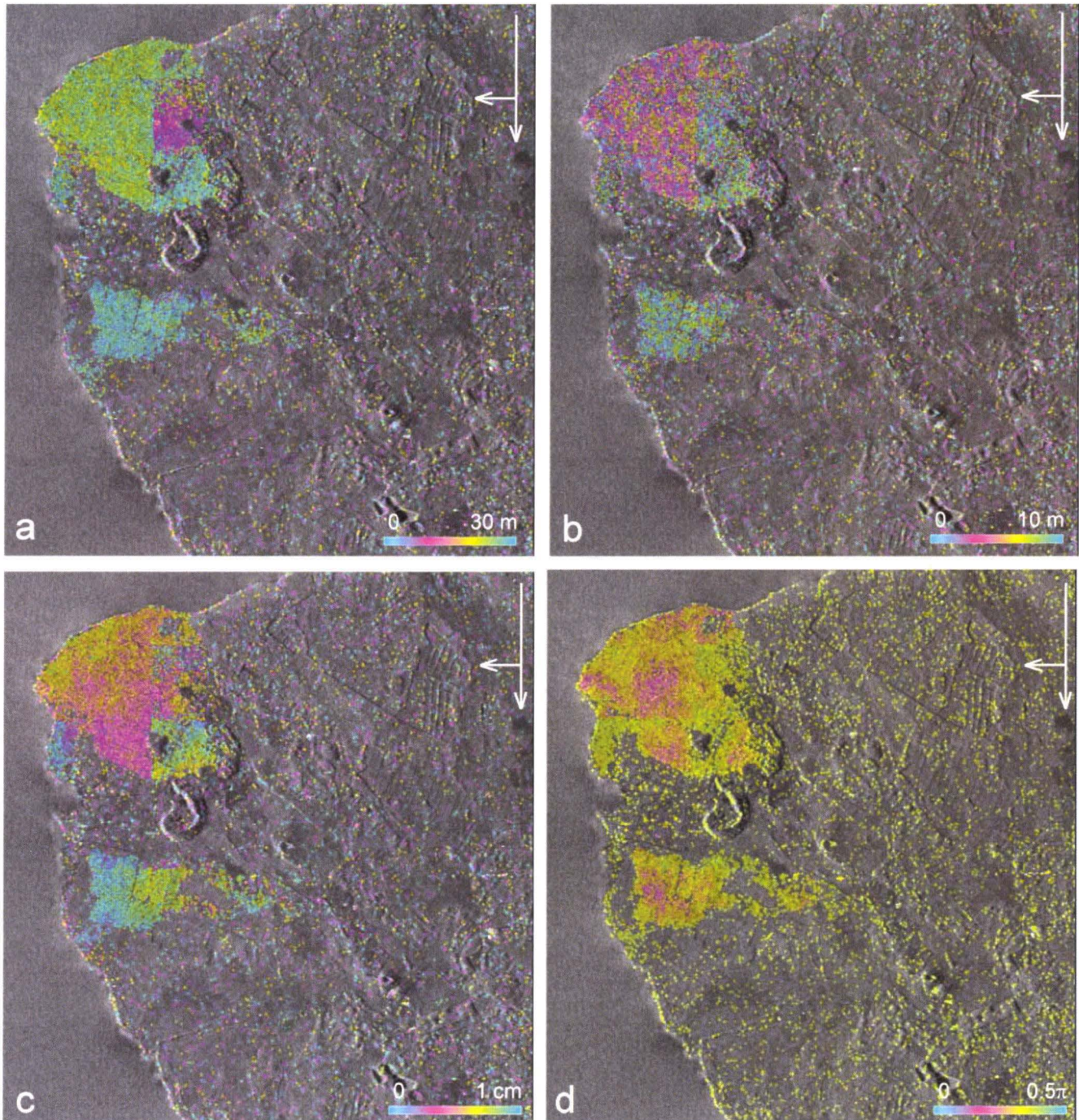


Figure 34: Results from 1st multi-patch regression (GAMMA command `multi_def_pt`). All initial wrapped differential interferograms used as input point data stack. (a) Height correction pdh1 (30 m/color-cycle, values in the southern flow range from -10 to 10 m), (b) pdh1 (10 m/color-cycle), (c) linear deformation rate (0.01 m/color-cycle, cyan: ~ -1.5 mm, yellow: ~ -2.5 mm), (d) phase standard deviation from regression-fit ($1/2 \pi$ rad/color-cycle, red: ~ 0.7 rad). Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

The areas highlighted by the selected points on the east tip of the Big Island can be divided in two main areas, a northern and a southern lava flow created in the 1960 eruption. The northern flow shows an obvious color discontinuity in the initial height corrections ([Figure 34a, b](#)), related to unwrapping errors along the patch boundaries. The project focuses on the southern flow which does not show any errors. It is therefore acceptable to disregard the patch boundary discontinuities in the northern flow. Height corrections range between -10 and 10 m and will be used to correct the available Digital Elevation Model (DEM). Height corrections are determined through the linear dependence of the interferometric phase on the perpendicular baseline. Note: `prasdt_pwr24` creates raster images, whereas `pdisd_t_pwr24` displays the data in GAMMA so the actual values can be observed.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pddef1 1 rml_i_1_5
_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 0.01 1
```

The linear deformation rate estimate is based on the linear dependence of the interferometric phase on temporal baseline. Again, the discontinuities in the northern flow ([Figure 34c](#)) can be ignored, since the project focuses on the southern flow.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/psigma1 1 rml_i_1_5
_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 1.5 1
```

The standard deviation from the regression fit is calculated for each point. It can be used as a quality measure ([Figure 34d](#)) for the regression. Sigma increases in value with distance from reference point.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pres1 - rml_i_1_5
_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 2.5 1
```

The residual phases ([Figure 35: Residuals at 2.5 radians](#) PDF 4.6 MB, [Figure 36: Residuals at 12.6 radians](#) PDF 4.7 MB) represent the deviation from the regression fit and are calculated for each record in the stack. Using a phase scale of 12.6 radians results in a smooth southern flow indicating a good phase model [*Wegmüller, 2005*] whereas 2.5 radians is used to show greater detail in the variations.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/punw1 - rml_i_1_5
_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 12.6 1
```

Unwrapped interferometric phase examples ([Figure 37](#) PDF 4.7 MB) are calculated for each record in the stack.

Multi_def_pt unwraps the interferometric phases temporarily throughout the entire stack. Depending on baseline related errors or influence of atmospheric distortions, some of the unwrapping might be incorrect. Looking at each residual record, four scenes with patch related phase unwrapping errors in the southern flow can be identified (2, 3, 7, 12). In the next step **def_mod_pt** is run without those “incorrectly unwrapped” scenes. The IPTA command **def_mod_pt** conducts the regression analysis using a single patch approach, which is possible because the data is already unwrapped.

IPTA-6.2.2 Estimation of Linear Deformation Rate, Height Correction and Residual Phase over the Entire Scene
Re-running the regression analysis for correctly unwrapped layers ([itab selection](#)).

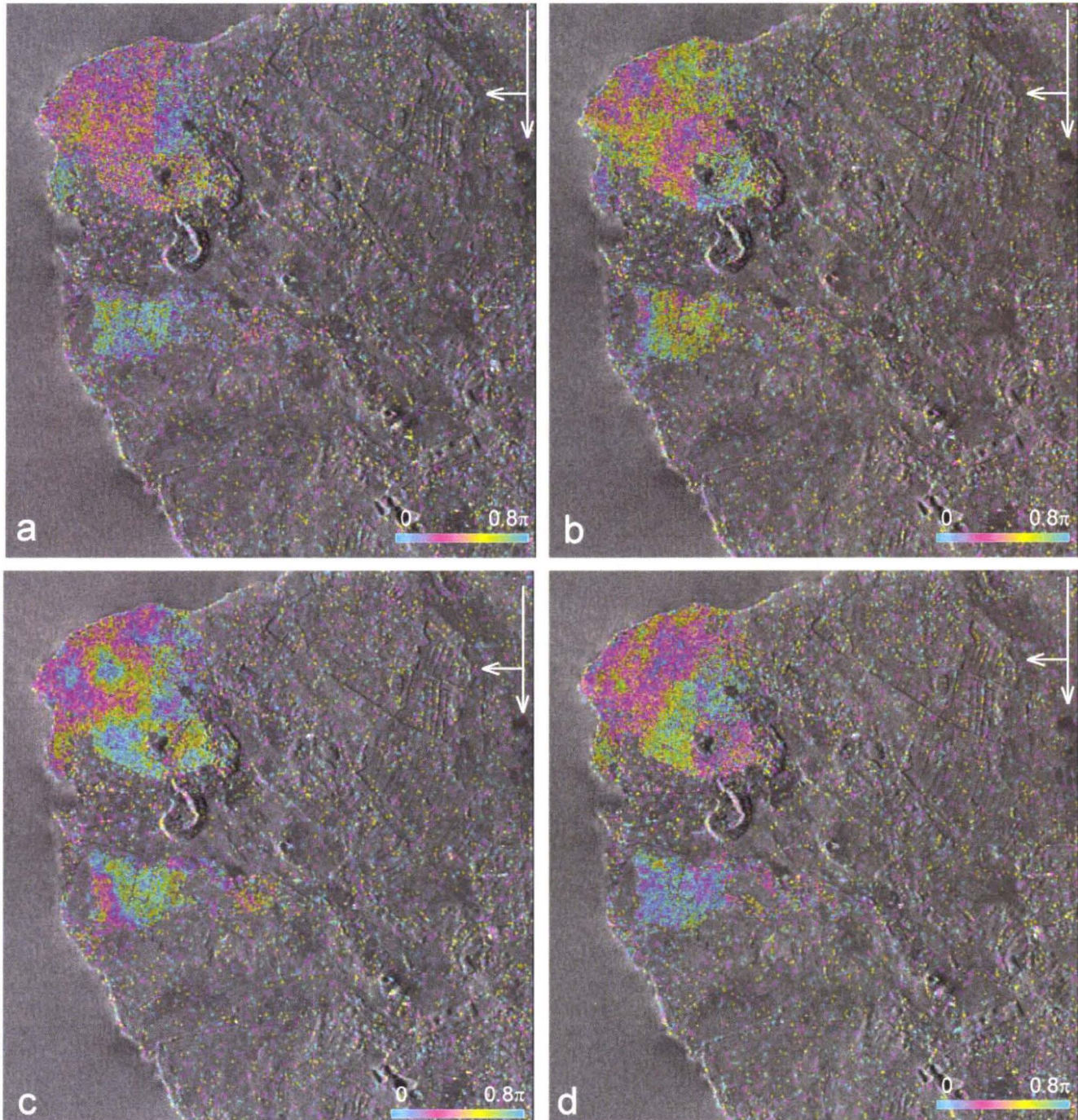


Figure 35: Residual phases from 1st multi-patch regression (GAMMA command `multi_def_pt`). One color-cycle corresponds to 2.5 rad to show minute difference to residuals from later processing. All initial wrapped differential interferograms used as input point data stack. (a) Pres1_001, (b) pres1_002, (c) pres1_004, (d) pres1_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

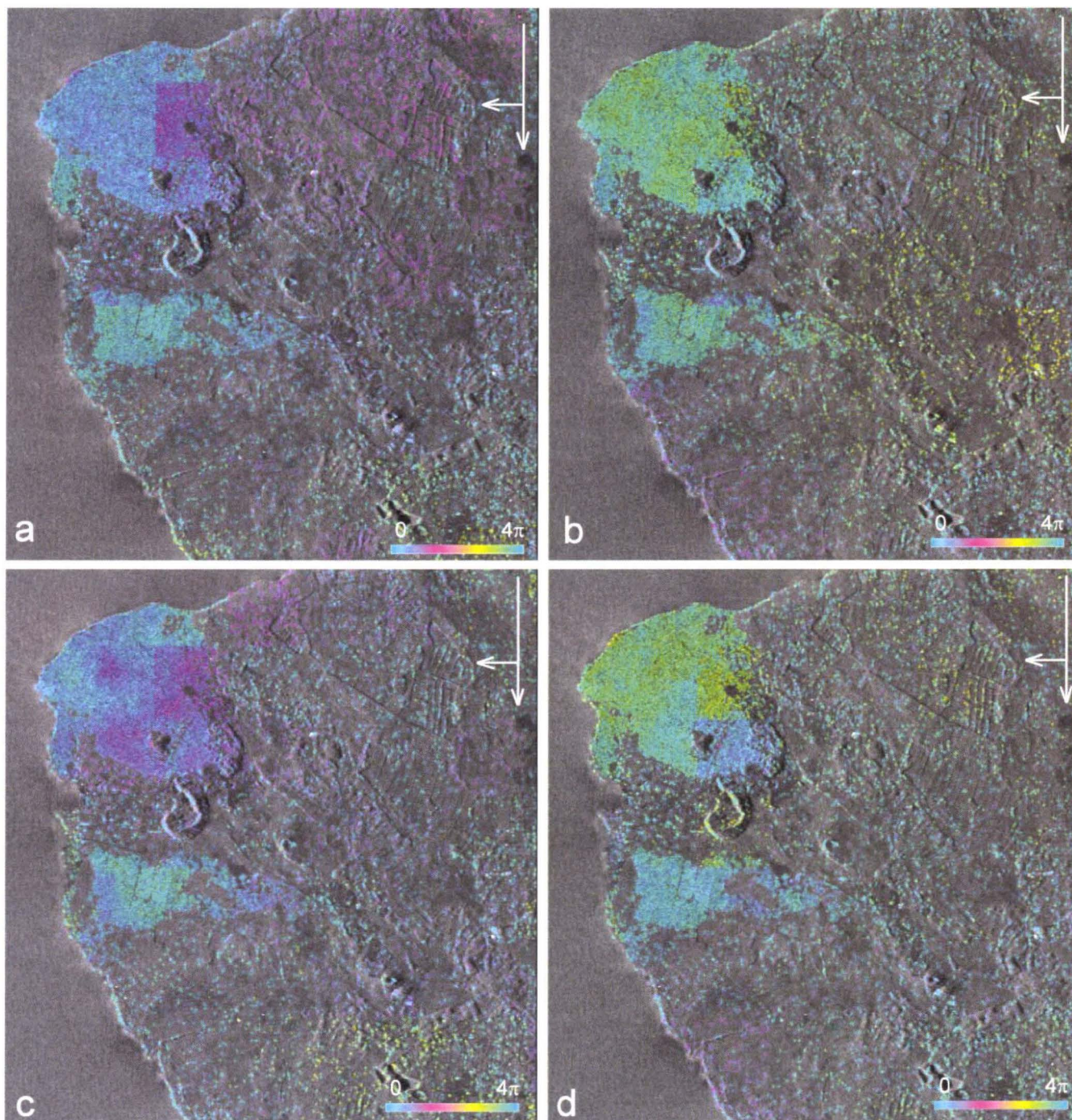


Figure 36: Residual phases from 1st multi-patch regression (GAMMA command `multi_def_pt`). One color-cycle corresponds to 12.6 rad. All initial wrapped differential interferograms used as input point data stack. (a) Pres1_001, (b) pres1_002, (c) pres1_004, (d) pres1_008. Generated using GAMMA script `prastd_pwr24`. Note the smoothness of the southern flow. Arrows indicate satellite's flight- and the SAR's look-direction.

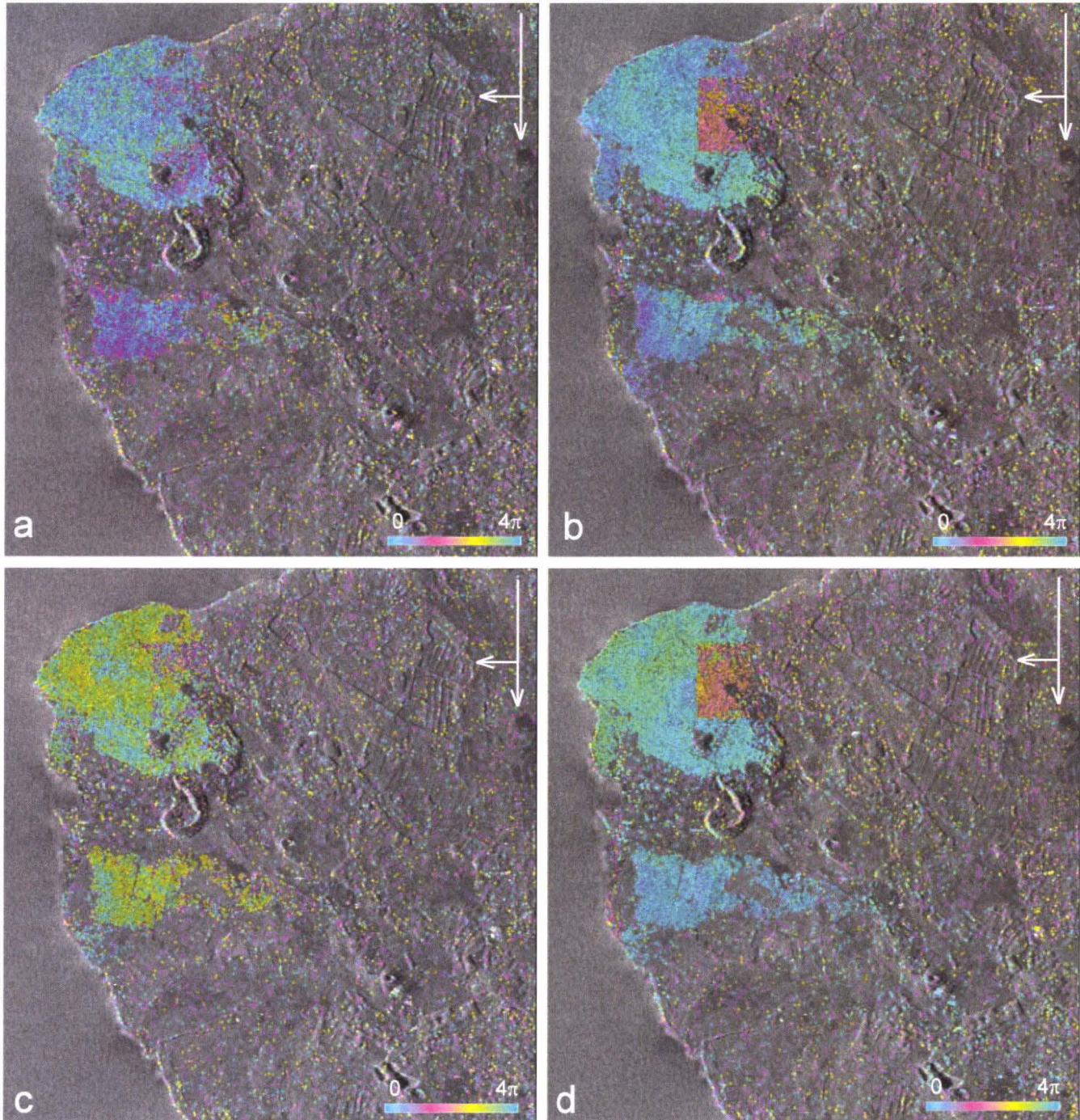


Figure 37: Unwrapped interferometric phases from 1st multi-patch regression (GAMMA command `multi_def_pt`). One color-cycle corresponds to 12.6 rad. All initial wrapped differential interferograms used as input point data stack. (a) Punw1_001, (b) punw1_002, (c) punw1_004, (d) punw1_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

```
mkdir def_mod_pt
```

```
def_mod_pt pt multi_def_pt1/pmask1 pSLC_par - itab_selection pbase 0 multi_def_pt1/punw1 0
23558 def_mod_pt/pres2 def_mod_pt/pdh2 def_mod_pt/pddef2 def_mod_pt/punw2 def_mod_pt/psigma2
def_mod_pt/pmask2 25. 0.03 3.0 2 - - -
```

```
number of points with sigma < sigma_max=3.000 : 19646
```

- Create images of output ([Figure 38](#) PDF 4.6 MB)

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pdh2 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 30.0 1
```

Height corrections ([Figure 38a, b](#)) are a little bit smaller without the influence of “incorrectly unwrapped” data.

The northern flow still shows an obvious color discontinuity, related to unwrapping errors, which still can be neglected since the project only focuses on the southern flow. The differences are more visible using a lower phase scaling of 2.5 radians. The slight improvements in the height corrections can be observed by comparing the actual values using the command `disdt_pwr24`.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pddef2 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 0.01 1
```

The linear deformation estimate ([Figure 38c](#)) without influence of “incorrectly unwrapped” data does not change significantly. Again, the discontinuities in the northern flow can be ignored.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/psigma2 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 1.5 1
```

The point quality measure sigma ([Figure 38d](#)) increases in value with distance from reference point. Notice how the amount of red points increases in the southern flow (especially around the reference point), indicating lower sigma values, hence an improvement in the model.

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pres2 - rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 2.5 1
```

The residual phases ([Figure 39: Residuals at 2.5 radians](#) PDF 4 MB, [Figure 40: Residuals at 12.6 radians](#) PDF 4.1 MB) represent the deviation from the regression fit and are calculated for each record in the stack. The scale factor of 2.5 allows detailed comparison to residuals from `multi_def_pt` run ([Figure 35](#)). Using a phase scale of 12.6 results in a even smoother southern flow indicating a better phase model [*Wegmüller, 2005*].

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/punw2 - rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 12.6 1
```

Unwrapped interferometric phases ([Figure 41](#) PDF 4.1 MB) are calculated for each record in the stack.

IPTA-7 Update the Height and Deformation Model

- Calculate linear combination of records in point data stacks

Apply height correction to `pdem`, `phgt1` becomes the current height model

```
lin_comb_pt pt def_mod_pt/pmask2 pdem2 1 def_mod_pt/pdh2 1 phgt1 1 -0. 1. 1. 2 0
```

- Copy the refined linear deformation rate estimate to the current deformation model `pdef1`

```
cp def_mod_pt/pddef2 pdef1
```

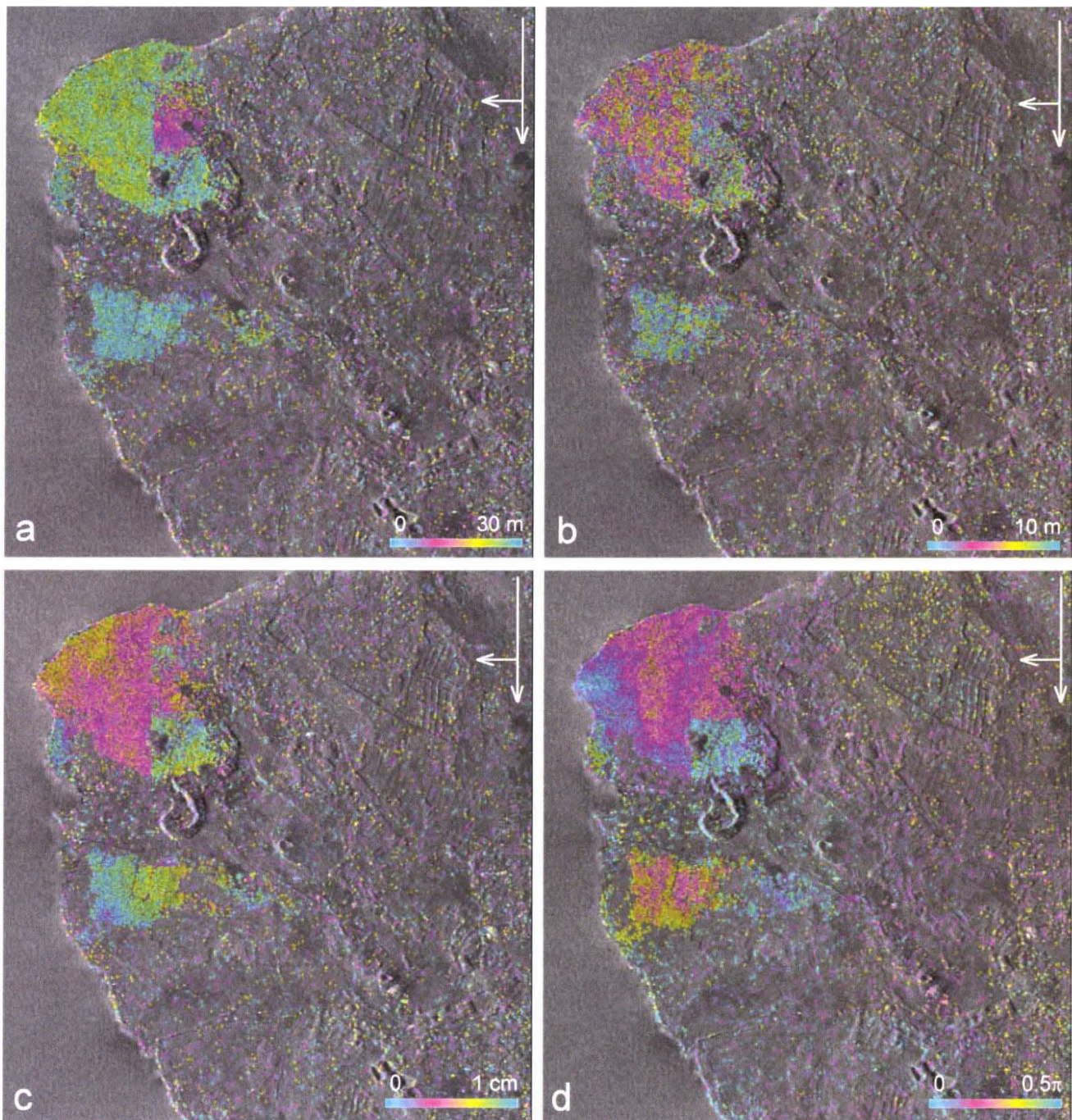


Figure 38: Results from 1st single-patch regression (GAMMA command `def_mod_pt`). Only correctly unwrapped differential interferograms used as input point data stack. Compare to Figure 34. (a) Height correction `pdh2` (30 m/color-cycle, most values in the southern flow are between -8 and 8m), (b) `pdh2` (10 m/color-cycle), (c) linear deformation rate (0.01 m/color-cycle, cyan: ~ -1.5 mm, yellow: ~ -2.5 mm), (d) phase standard deviation from regression-fit ($1/2 \pi$ rad/color-cycle, red: ~ 0.7 rad). Notice the increased number of red sigma values. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

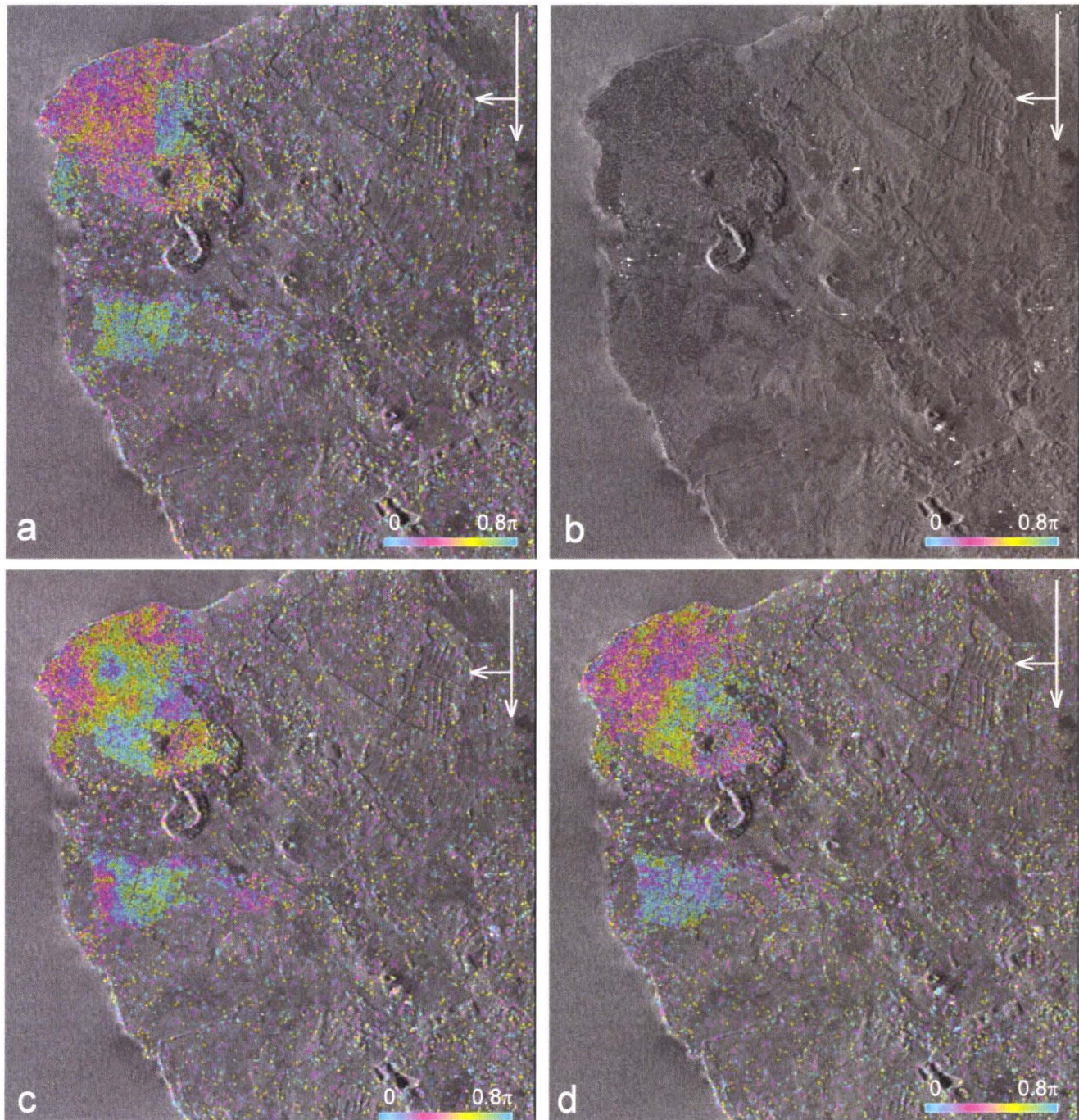


Figure 39: Residual phases from 1st single-patch regression (GAMMA command `def_mod_pt`). One color-cycle corresponds to 2.5 rad to show minute difference to residuals from earlier processing (Figure 35). Only correctly unwrapped differential interferograms used as input point data stack. (a) Pres2_001, (b) pres2_002, no data because pres1_002 showed unwrapping errors and was not included in this regression. (c) pres2_004, (d) pres2_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

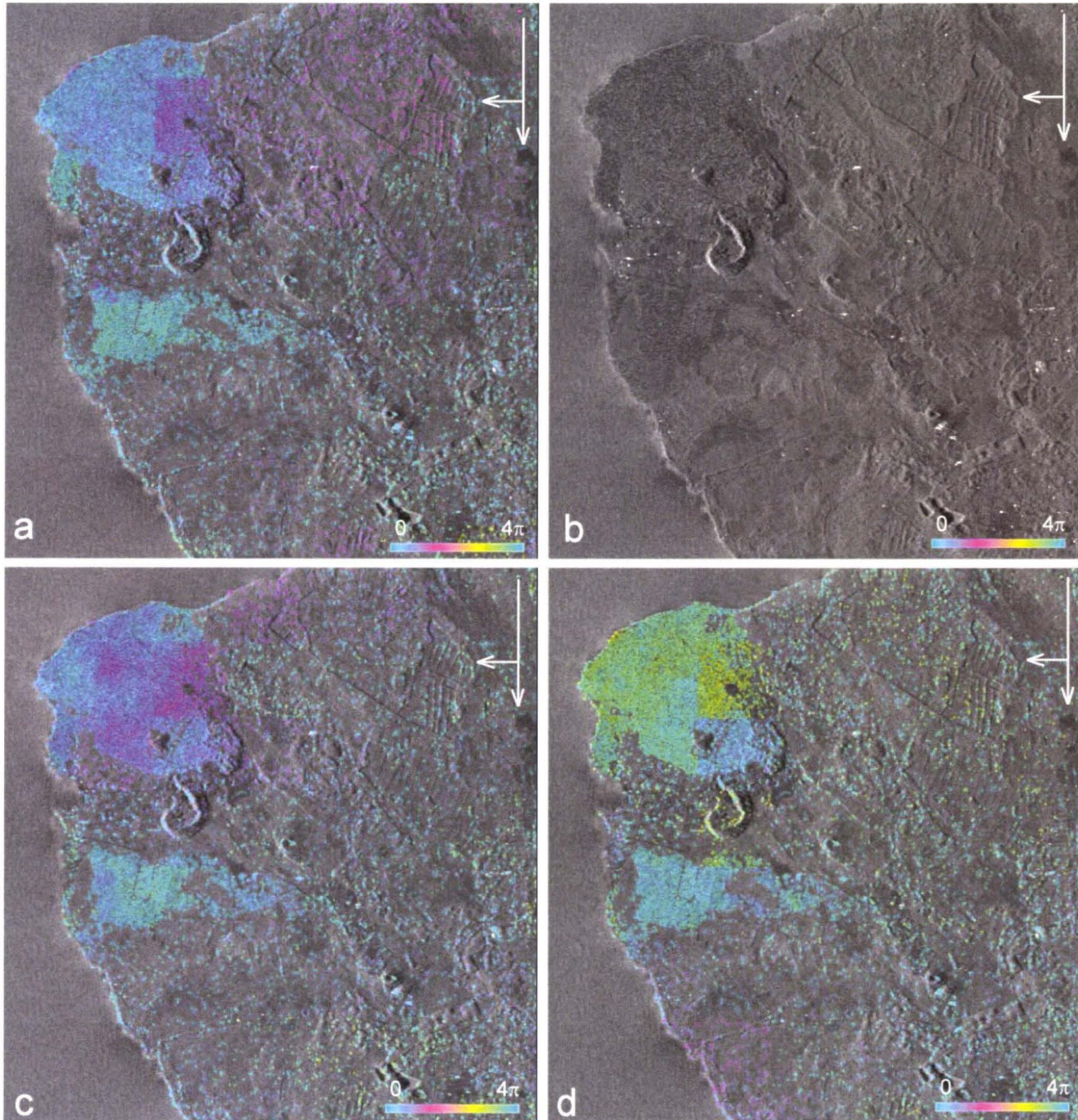


Figure 40: Residual phases from 1st single-patch regression (GAMMA command `def_mod_pt`). One color-cycle corresponds to 12.6 rad. Only correctly unwrapped differential interferograms used as input point data stack. (a) Pres2_001, (b) pres2_002, no data because pres1_002 showed unwrapping errors and was not included in this regression, (c) pres2_004, (d) pres2_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

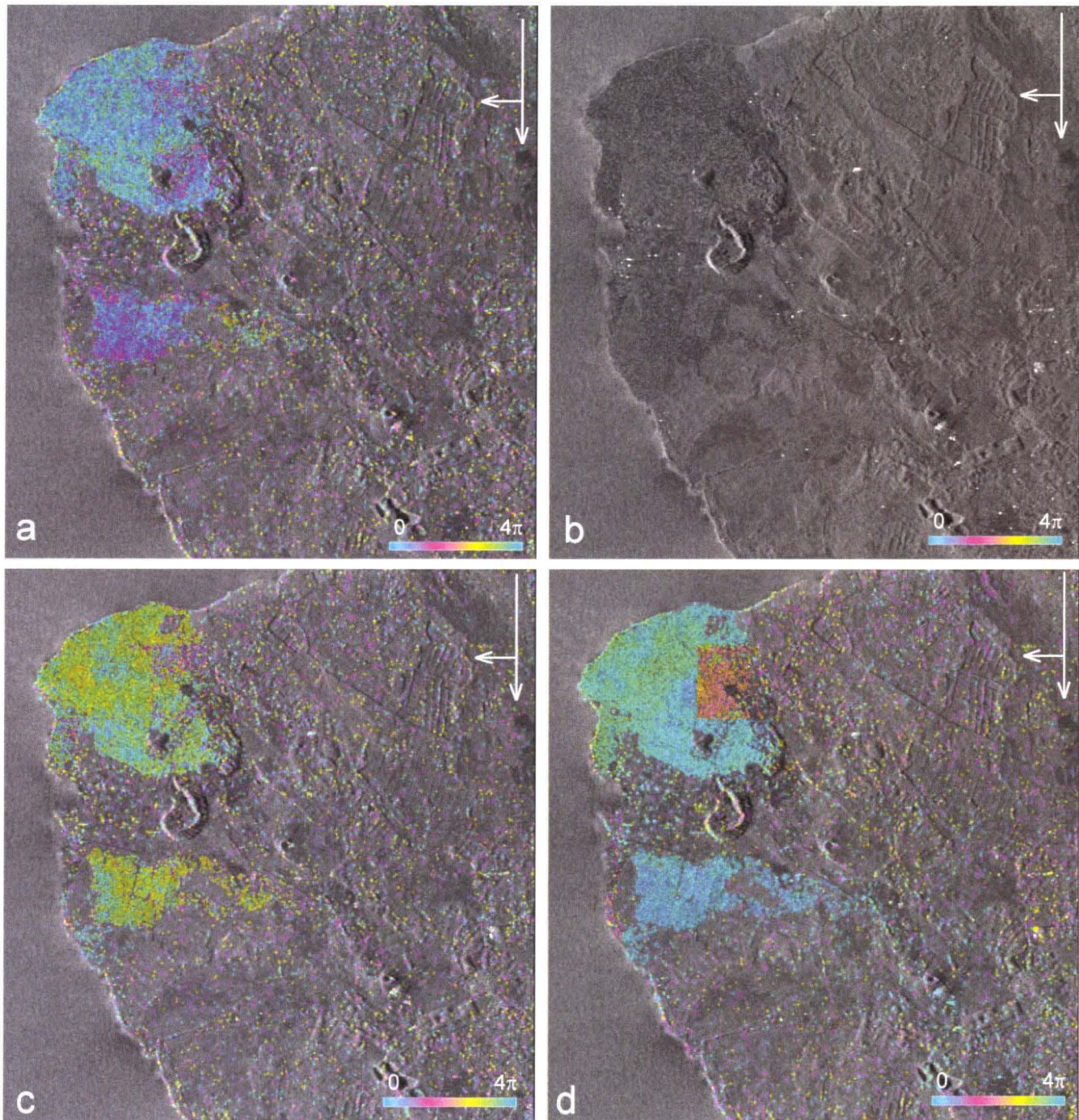


Figure 41: Unwrapped interferometric phases from 1st single-patch regression (GAMMA command `def_mod_pt`). One color-cycle corresponds to 12.6 rad. Only correctly unwrapped differential interferograms used as input point data stack. (a) Punw2_001, (b) punw2_002, no data because pres1_002 showed unwrapping errors and was not included in this regression, (c) punw2_004, (d) punw2_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

IPTA-8 Calculate Updated Differential Interferograms

- Calculate differential interferograms using *phgt1* and *pdef1* ([log](#))

```
mk_diff_all pt def_mod_pt/pmask2 pSLC_par itab_single pbase 0 pint 1 phgt1 pdef1 -
def_mod_pt/psim_unw0 def_mod_pt/pdiff0 def_mod_pt/mk_diff_all0.log
```

- Create images of updated differential interferograms ([log](#)) ([Figure 42](#) PDF 4.5 MB)

```
mk_2d_im pt def_mod_pt/pmask2 itab_single rslc_kapoho/20041208.rslc.par def_mod_pt/pdiff0 1 -
rml1_1_5_kapoho/rml1_1_5.ave.par 0 3 1.2 1 def_mod_pt/diff0 rml1_1_5_kapoho/rml1_1_5.ave - 0
```

IPTA-9 Regression Analysis on Updated Differential Interferograms

- Include additional pairs into solution of model

```
mkdir multi_def_pt2
multi_def_pt pt def_mod_pt/pmask2 pSLC_par - itab_single pbase 0 def_mod_pt/pdiff0 1 23558
multi_def_pt2/pres1 multi_def_pt2/pdh1 multi_def_pt2/pddef1 multi_def_pt2/punw1
multi_def_pt2/psigma1 multi_def_pt2/pmask1 3. 0.003 70 1.3 0.9 2 1 -1 -1
```

```
number of points with sigma < sigma_max = 1.300 : 12728
```

- Create images of output ([Figure 43](#) PDF 3.4 MB)

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pdh1 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 30.0 1
```

The height corrections ([Figure 43a, b](#)) appear smooth in the southern flow, indicating that major corrections have already been done. The earlier height corrections, applied to the DEM in IPTA-7, corrected errors in the DEM sufficiently.

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pddef1 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 0.01 1
```

The linear deformation rate estimate ([Figure 43c](#)) shows a similar behavior, a smooth southern flow, again, indicating that major corrections have already been done.

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/psigma1 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 1.5 1
```

The point quality measure sigma ([Figure 43d](#)) did not change significantly in comparison to [Figure 34d](#).

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pres1 - rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 12.6 1
```

As one can see in [Figure 44](#) (PDF 3.4 MB), fewer points are displayed in comparison to the residual plots from [Figure 34](#), indicating that only “good” points are used. Higher thresholds have selected better points. The residuals themselves have not changed significantly. [Figure 45](#) (PDF 3.4 MB) shows the residual phases at a phase scale of 12.6.

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/punw1 - rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 12.6 1
```

Example of unwrapped phase can be seen in the four examples in [Figure 46](#) (PDF 3.4 MB).

Check residuals carefully for unwrapping errors: none in this case.

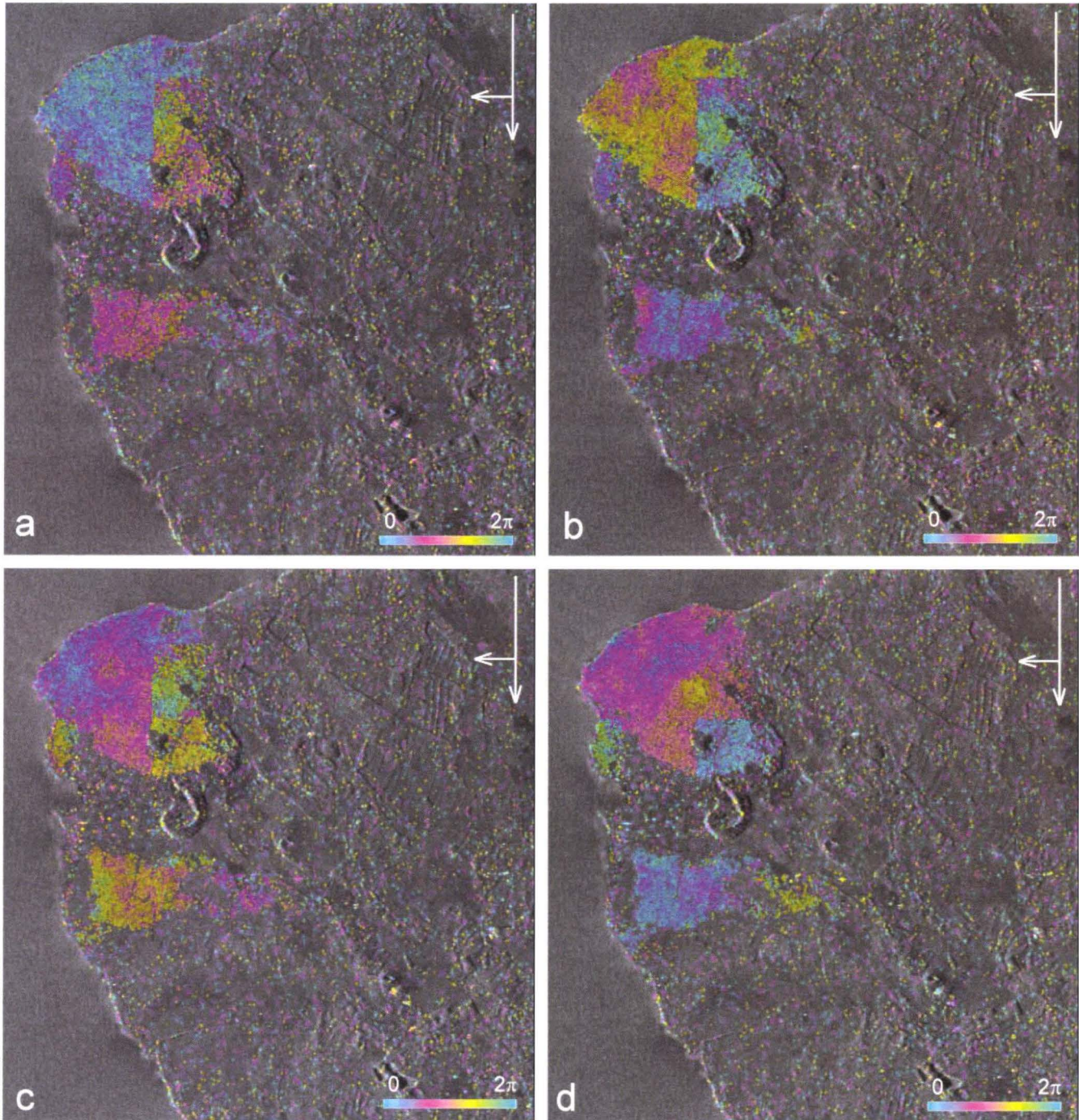


Figure 42: Four layers of the updated wrapped differential interferometric point data stack `pdiff0`: (a) `pdiff0_001`, (b) `pdiff0_002`, (c) `pdiff0_004`, (d) `pdiff0_008`. The topographic phase subtracted from the interferograms now accounts for the height corrections and the linear deformation estimate. Generated using GAMMA script `mk_2d_im`. Arrows indicate satellite's flight- and the SAR's look-direction.

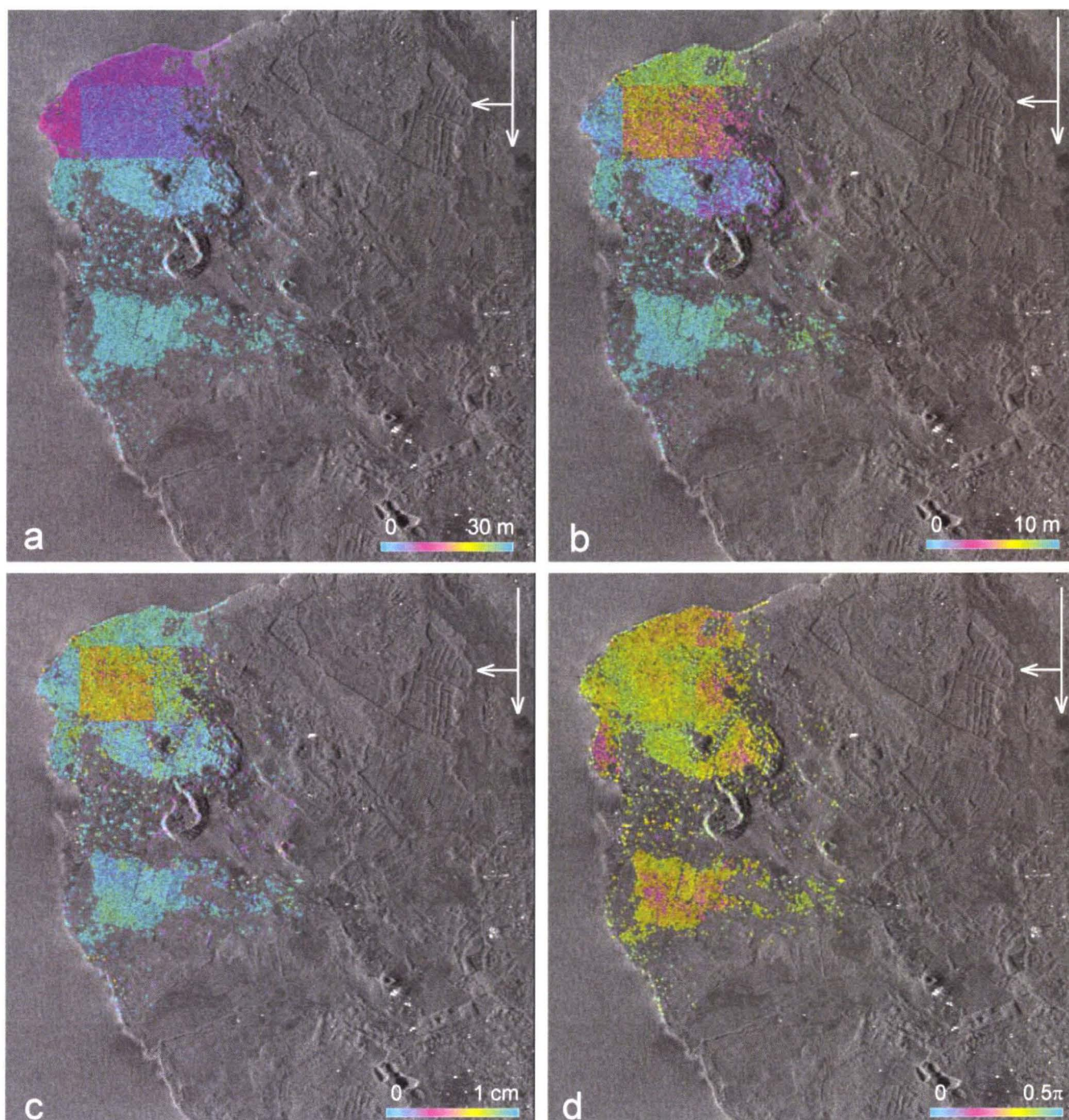


Figure 43: Results from 2nd multi-patch regression (GAMMA command `multi_def_pt`). Inclusion of previously not accepted wrapped differential interferograms into solution. (a) Height correction `pdh1` (30 m/color-cycle, values in the southern flow are mostly below 15 cm), (b) `pdh1` (10 m/color-cycle), (c) linear deformation rate (0.01 m/color-cycle, values in submillimeter range), (d) phase standard deviation from regression-fit ($1/2 \pi$ rad/color-cycle, red: ~ 0.7 rad). Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

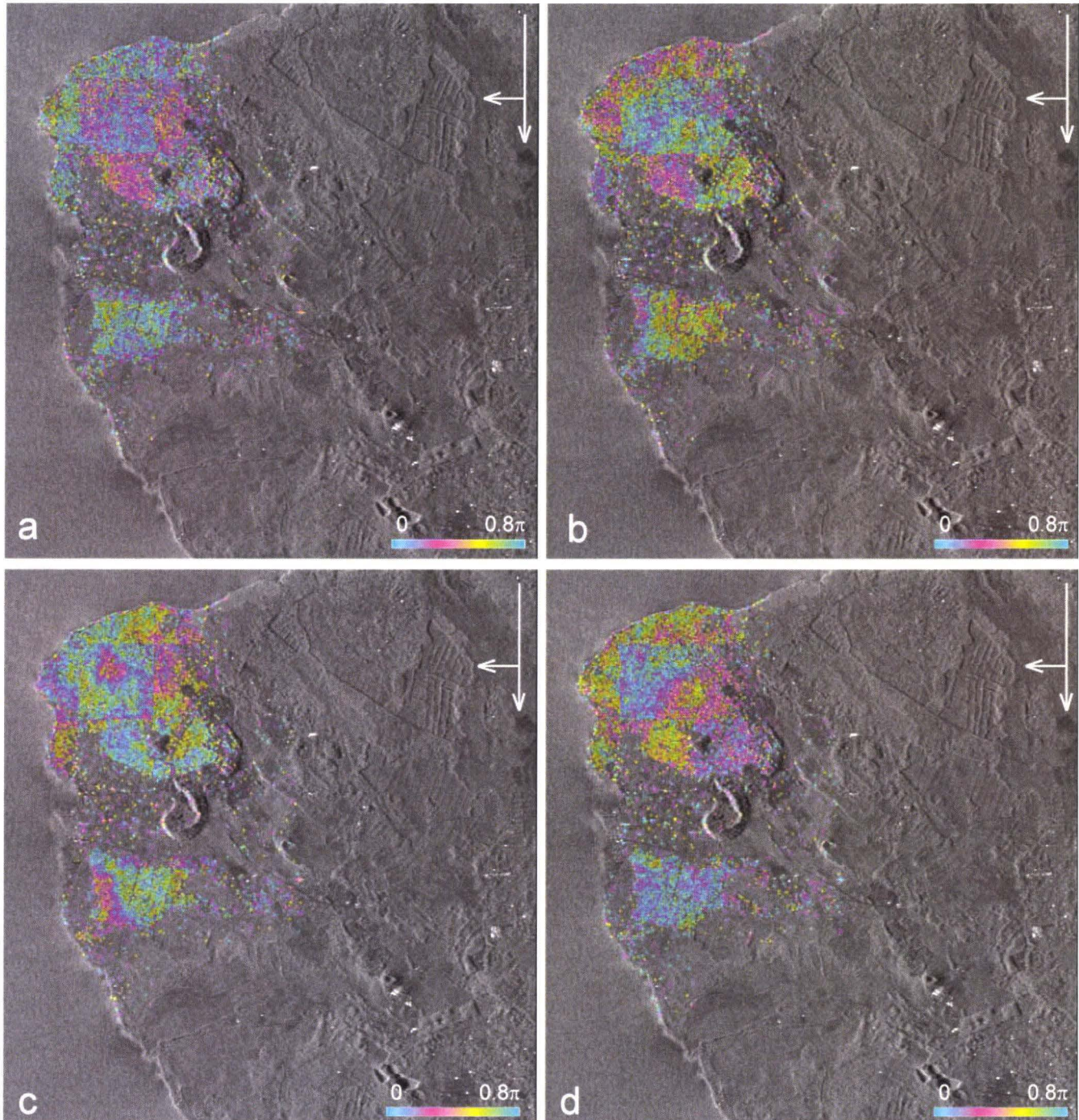


Figure 44: Residual phases from 2nd multi-patch regression (GAMMA command `multi_def_pt`). One color-cycle corresponds to 2.5 rad to show minute difference to residuals from later processing. Inclusion of previously not accepted wrapped differential interferograms into solution. (a) Pres1_001, (b) pres1_002, (c) pres1_004, (d) pres1_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

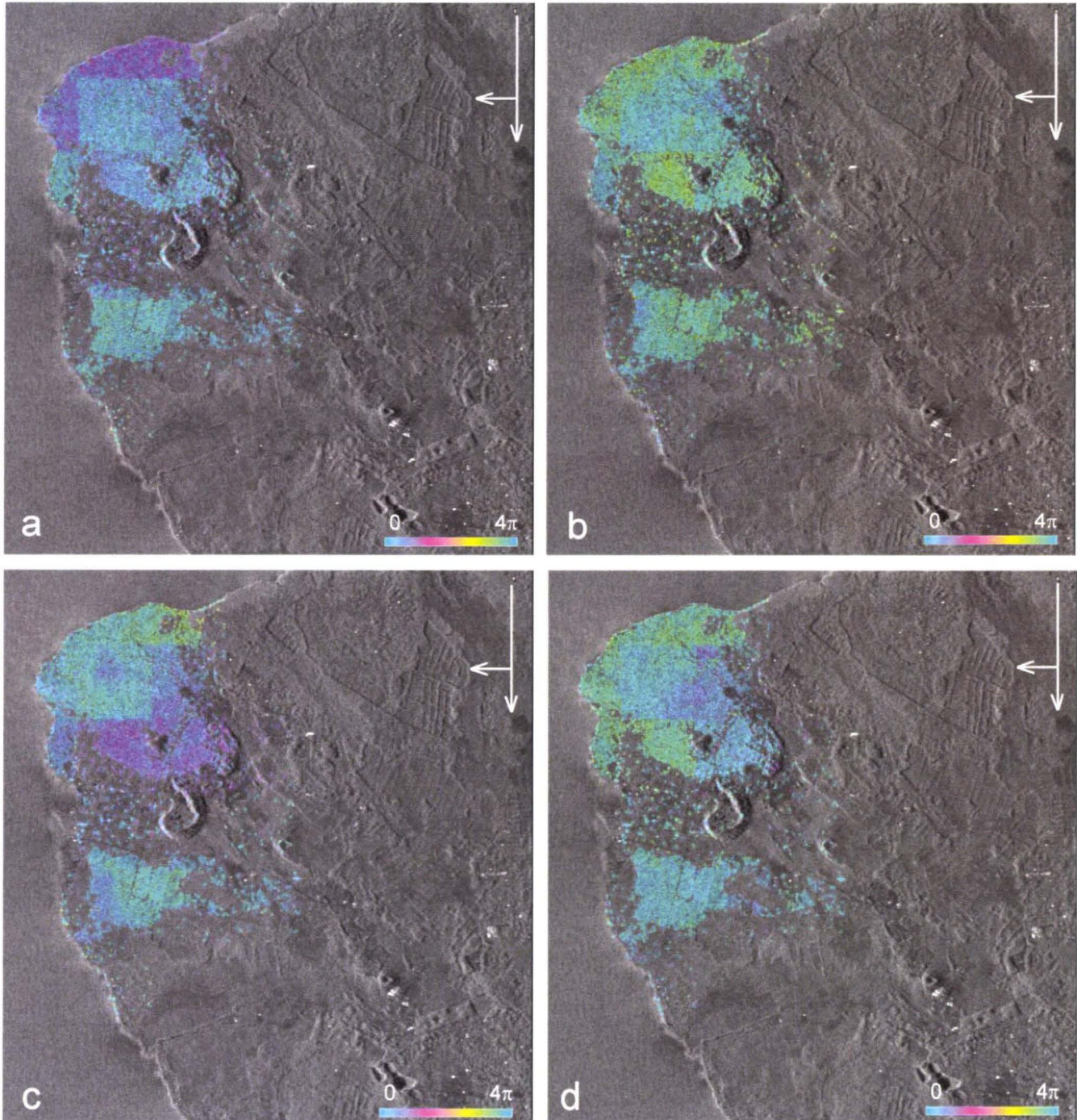


Figure 45: Residual phases from 2nd multi-patch regression (GAMMA command `multi_def_pt`). One color-cycle corresponds to 12.6 rad to show minute difference to residuals from later processing. Inclusion of previously not accepted wrapped differential interferograms into solution. (a) Pres1_001, (b) pres1_002, (c) pres1_004, (d) pres1_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

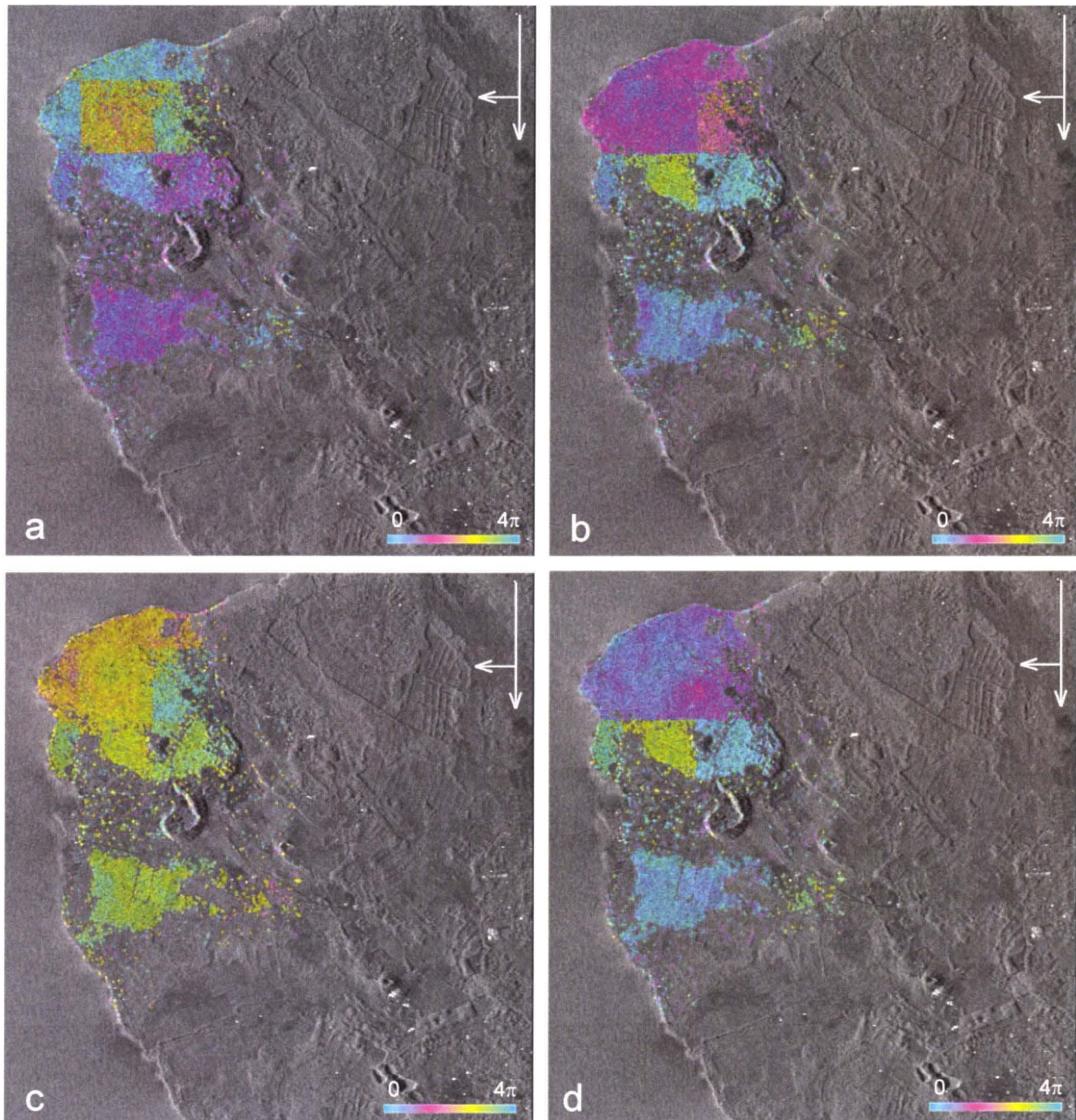


Figure 46: Unwrapped interferometric phases from 2nd multi-patch regression (GAMMA command `multi_def_pt`). One color-cycle corresponds to 12.6 rad. Inclusion of previously not accepted wrapped differential interferograms into solution. (a) Punw1_001, (b) Punw1_002, (c) Punw1_004, (d) Punw1_008. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

IPTA-10 Baseline Refinement

A baseline refinement can only be done if the area is large enough to provide a sufficient number of points to refine. In this case there were not enough points. In future processing a larger area could be selected to allow a bigger number of good points. The principle flow of commands is similar to the ISP refinement. For details refer to IPTA example by Wegmüller [2005].

IPTA-11 Interpretation of Residual Phase

The residual phases are real valued unwrapped phases which can have values outside the interval $(-\pi, \pi)$. The residual phase is considered to be the sum of non-linear deformation phase, atmospheric phase and phase noise (see above). These three components have different spatial and temporal characteristics which can be used to discriminate among them. The non-linear deformation phase is the difference between the total and the linear deformation phase. It can be depicted as non-uniform movements of any scale. Atmospheric path delay is caused by heterogeneities in the tropospheric water vapor content and in the ionosphere. It is largely random in the temporal domain. Atmospheric signatures are usually at a larger spatial scale; however distortions can occur at any scale and are mostly below one phase cycle. Phase noise is temporally and spatially random. The level of phase noise is directly related to the point quality [Wegmüller, 2005].

The Kapoho project investigates the small-scale deformation features identified in the differential interferograms. Taking into account the absence of any large-scale non-linear deformation signals, all large-scale deviations from the regression fit are remnants of atmospheric path delay, which is determined to be mostly uncorrelated between different pairs. It also is important to minimize phase noise in order to ensure high quality point deformation.

IPTA-11.1 Estimate and Apply Atmospheric Phase

- Spatially filter the residual phases to estimate large-scale components and remove reference point bias


```
spf_pt pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/pres1 atm/pres1_spf_25_1 -
2 25 1 -

spf_pt pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/pres1 atm/pres1_spf_10_1 -
2 10 1 -
```
- Rename filtered residual phase to “atmospheric phase” ([Figure 47](#) PDF 3.4 MB)


```
cp atm/pres1_spf_25_1 atm/patm1
```
- Calculate atmospheric corrections for the interferometric pairs (auto-interferogram influence)


```
lin_comb_pt pt multi_def_pt2/pmask1 atm/patm1 - atm/patm1 14 atm/patm1x - 0. 1. -1. 2 1
```
- Remove atmospheric phases from unwrapped differential interferograms


```
sub_phase_pt pt multi_def_pt2/pmask1 multi_def_pt2/punw1 - atm/patm1x multi_def_pt2/punw1_noatm
0 0
```
- Display the unwrapped phases (without auto-interferogram influence) ([Figure 48](#) PDF 3.4 MB)


```
prasd_t_pwr24 pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/punw1_noatm - rml_i_
1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 12.6 1
```

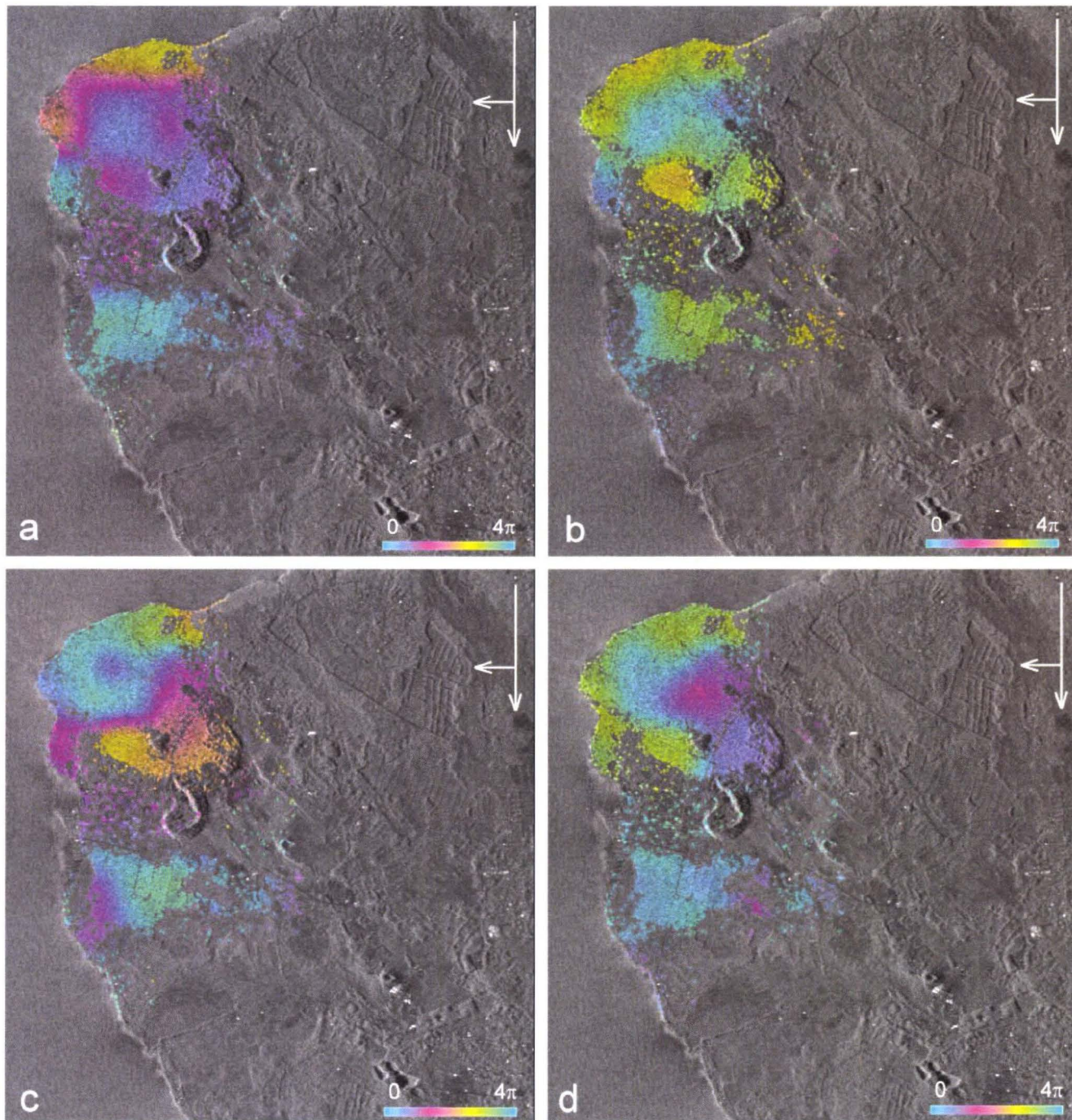


Figure 47: Spatially filtered residual phases from 2nd multi-patch regression (GAMMA command `multi_def_pt`). Triangular weighted average is the spatial filter used with a 25 sample wide radius. One color-cycle corresponds to 4 rad. (a) `Pres1_spf_25_001`, (b) `pres1_spf_25_002`, (c) `pres1_spf_25_004`, (d) `pres1_spf_25_008`. Generated using GAMMA script `prasdt_pwr24`. Arrows indicate satellite's flight- and the SAR's look-direction.

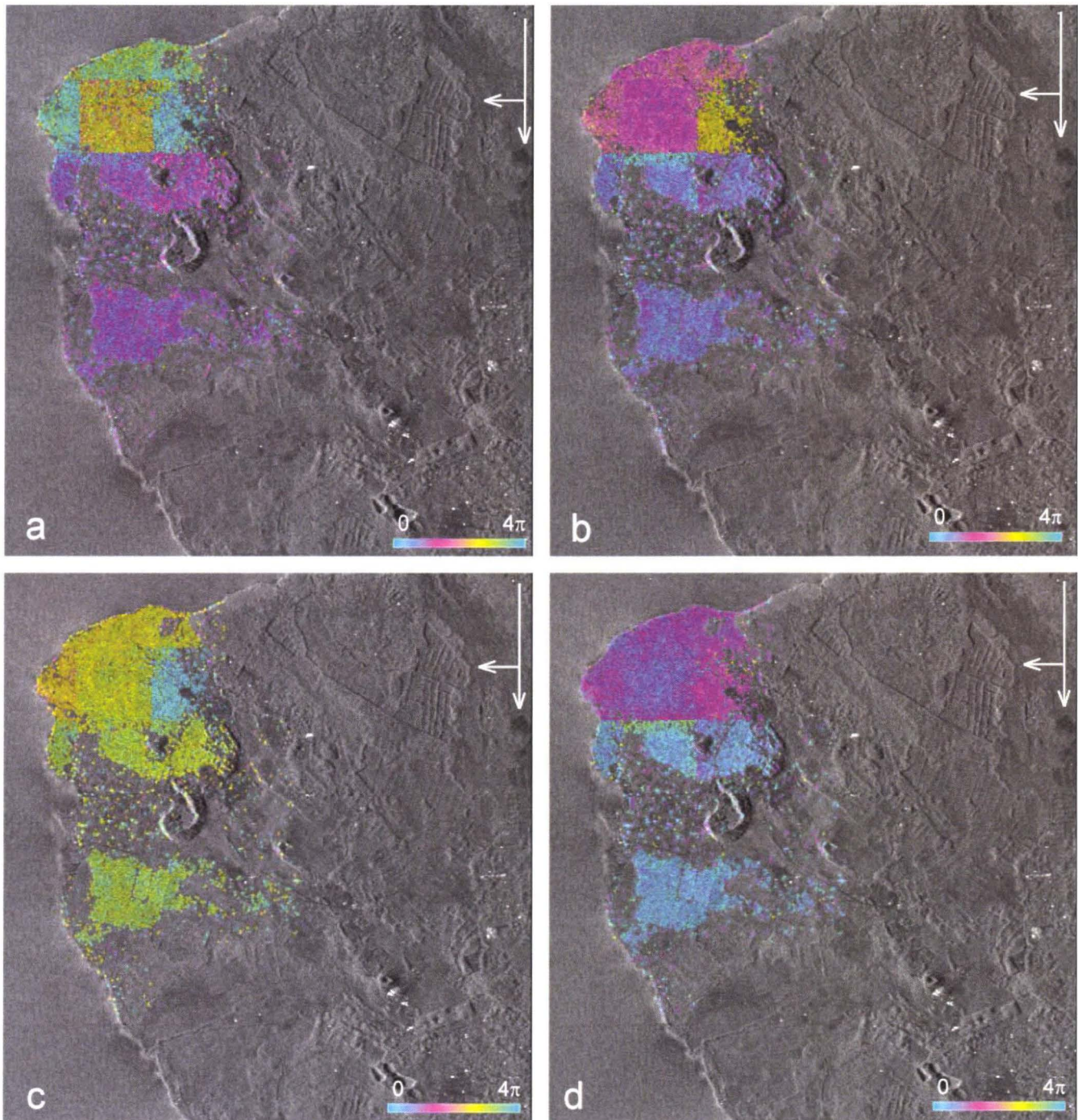


Figure 48: Unwrapped interferometric phases without the atmospheric path delay. One color-cycle corresponds to 12.6 rad. (a) Punw1_noatm_001, (b) punw1_noatm_002, (c) punw1_noatm_004, (d) punw1_noatm_008. Generated using GAMMA script prasdt_pwr24. Arrows indicate satellite's flight- and the SAR's look-direction.

- Reduce phase noise of interferometric phase phases relative to reference point

```
spf_pt pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/punw1_noatm
multi_def_pt2/punw1_noatm_spf - 2 25 0 - 23558 0
```

- Display the noise-free interferometric phases ([Figure 49](#) PDF 3.4 MB)

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/punw1_noatm_spf -
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 12.6 1
```

IPTA-11.2 Re-estimate Linear Deformation Rates and Phase Unwrapping

- Including atmospheric correction and using a filtered reference (*pdiff1_unwa*)

```
mkdir def_mod_pt2
def_mod_pt pt multi_def_pt2/pmask1 pSLC_par - itab_single pbase 0 multi_def_pt2/punw1
_noatm_spf 0 23558 def_mod_pt2/pres2 def_mod_pt2/pdh2 def_mod_pt2/pddef2 def_mod_pt2/punw2
def_mod_pt2/psigma2 def_mod_pt2/pmask2 25. 0.03 2.0 2 def_mod_pt2/pdh_err2a
def_mod_pt2/pdef_err2a def_mod_pt2/ppc_err2a
```

```
number of points with sigma < sigma_max=2.000 : 12721
```

IPTA-11.3 Visualization of Deformation Histories

- Update point heights and linear deformation rates

```
lin_comb_pt pt def_mod_pt2/pmask2 phgt1 1 def_mod_pt2/pdh2 1 phgt2 1 -0. 1. 1. 2 0
lin_comb_pt pt def_mod_pt2/pmask2 pdef1 1 def_mod_pt2/pddef2 1 pdef2 1 -0. 1. 1. 2 0
```

- Calculate deformation phase corresponding to estimated linear deformation rates

```
phase_sim_pt pt def_mod_pt2/pmask2 pSLC_par - itab_single - pbase - ptmp1 pdef2 1 0
```

Add last residual phase (incl. non-linear deformation phase and phase noise)

```
lin_comb_pt pt def_mod_pt2/pmask2 ptmp1 - def_mod_pt2/pres2 - pdef_phase1 - 0.0 1. 1. 2 1
```

Note: Atmospheric phase is not included.

- Convert combined unwrapped differential phase to line-of-sight displacement (m)

```
dispmap_pt pt def_mod_pt2/pmask2 pSLC_par itab_single pdef_phase1 phgt2 pdisp1 0
```

- Create optical reference to select points (displacement at the earliest date (layer 1))

```
prasdt_pwr24 pt def_mod_pt2/pmask2 rslc_kapoho/20041208.rslc.par pdisp1 1 rml1_1_5
_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 0.05 0
```

- Display IPTA deformation time series ([Figure 50](#) PDF 3 MB)

```
vu_disp pt def_mod_pt2/pmask2 pSLC_par itab_single pdisp1 pdef2 phgt2 def_mod_pt2/psigma2
def_mod_pt2/pdh_err2a def_mod_pt2/pdef_err2a - pdisp1.ras
```

Using the left-mouse button the deformation histories for each point will display in Grace. The deformation history for the reference point shows zero displacement. Clicking close to the area of interest, subsidence in the order of about 1-2 cm for the time between the first and last acquisition dates (20030212 – 20060308) can be found, corresponding to the blue values. The displacements are relative to the reference point.

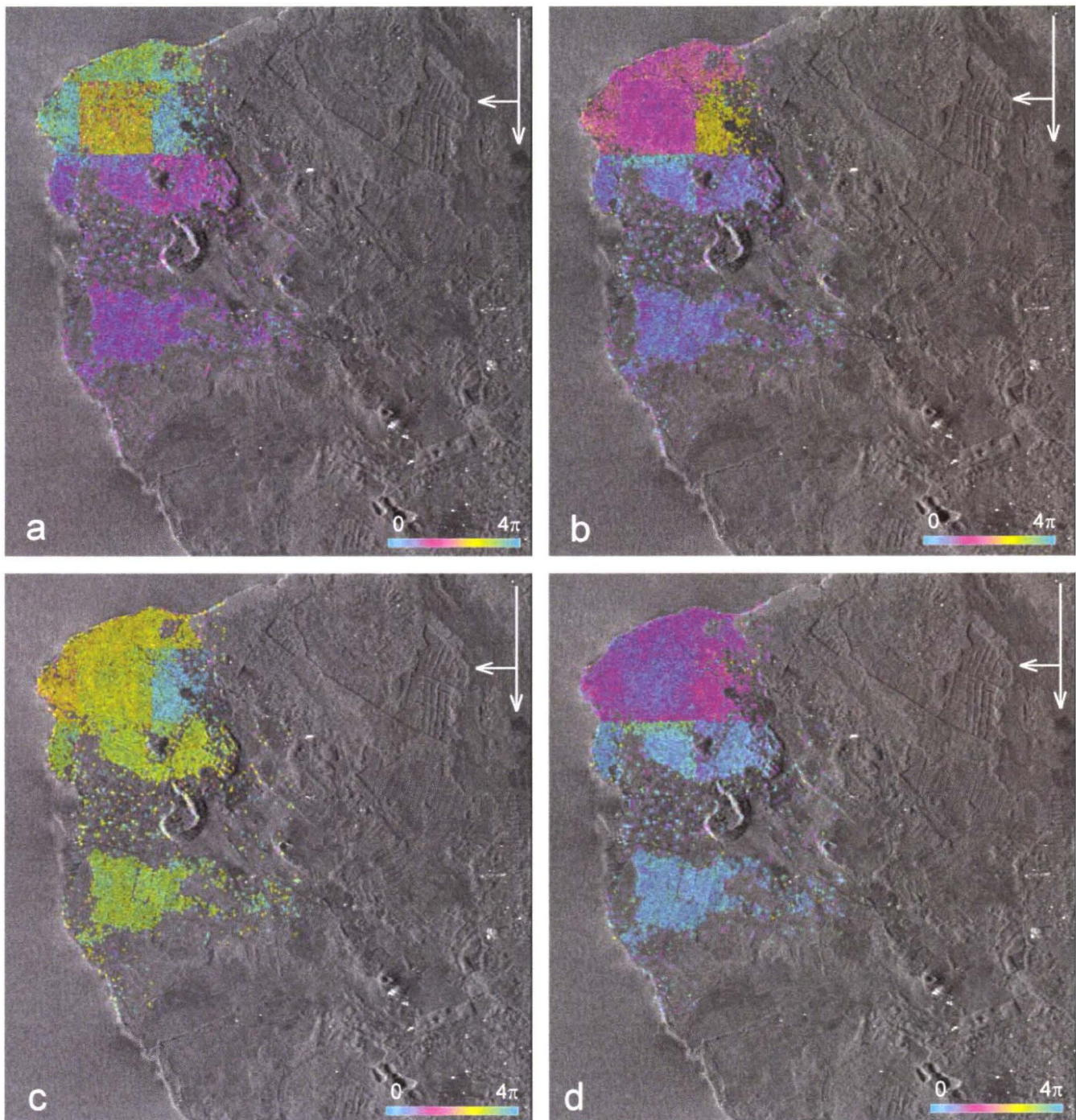


Figure 49: Filtered unwrapped interferometric phases without the atmospheric path delay. The interferometric phase is now noise free. One color-cycle corresponds to 12.6 rad. (a) Punw1_noatm_spf_001, (b) punw1_noatm_spf_002, (c) punw1_noatm_spf_004, (d) punw1_noatm_spf_008. Generated using GAMMA script prasdt_pwr24. Arrows indicate satellite's flight- and the SAR's look-direction.

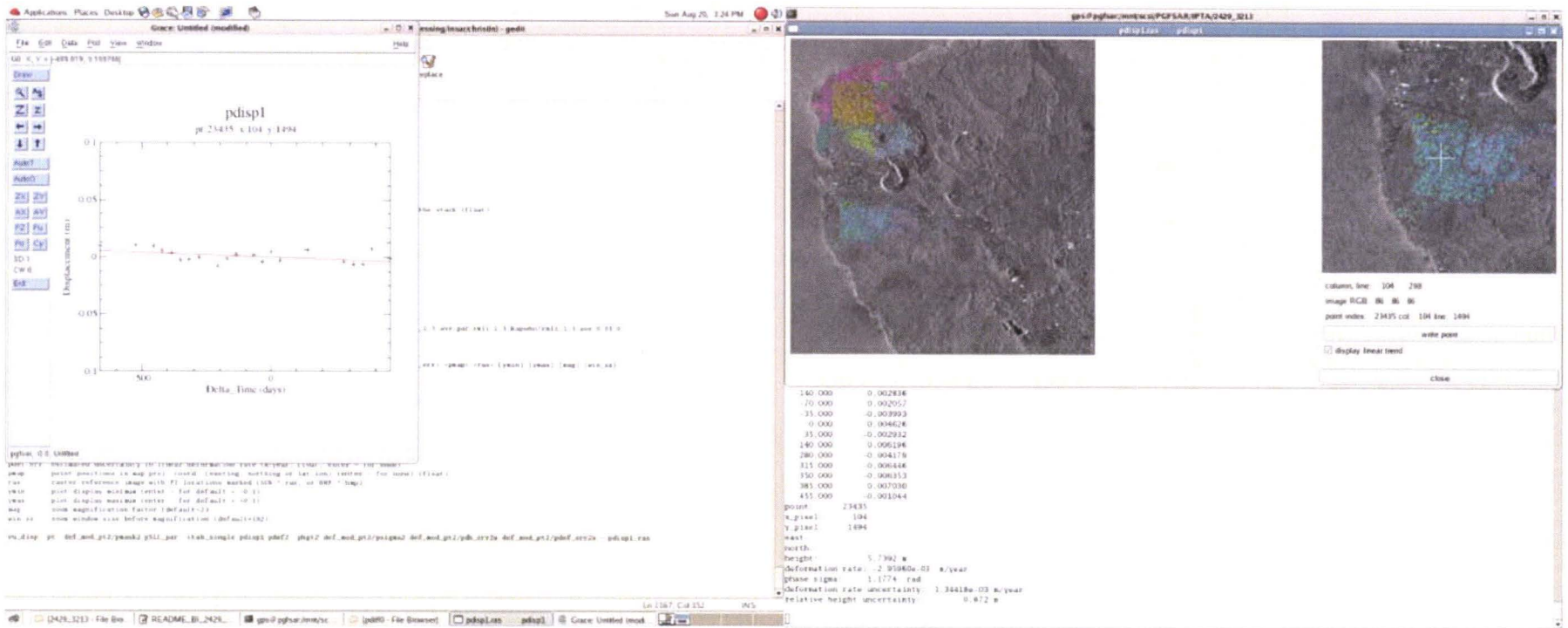


Figure 50: Visual output of GAMMA command `vu_disp` to view the deformation time series for every point. All deformation is with respect to the reference point.

3. References

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Appendices

Appendix A – Glossary of Abbreviations and Acronyms

ALOS	Advanced Land Observing Satellite
ASAR	Advanced Synthetic Aperture Radar
BPERP	Perpendicular baseline
DELFT	University of Delft, Netherlands (precise orbits)
DEM	Digital Elevation Model
DIFF	GAMMA Differential Interferometry software
DISP	GAMMA Display software
DORIS	Doppler Orbitography and Radiopositioning Integrated by Satellite
ERS	European Remote Sensing Satellite
ESA	European Space Agency
FFT	Fast Fourier Transform
FCOMPLEX	Float Complex (SLC format) pairs of 4-byte-float
FIR	Finite Impulse Response (digital filter)
GAMMA	GAMMA Remote Sensing Research and Consulting AG
GCP	Ground Control Point
GEO	GAMMA Geocoding software
HH	Horizontal polarization transmitted, Horizontal polarization received
I/Q	In-phase (I) (real) and quadrature (Q) component (imaginary) of complex radar return
InSAR	Interferometric Synthetic Aperture Radar
IPTA	GAMMA Interferometric Point Target Analysis
IS	Image Swath
ISP	GAMMA Interferometric SAR Processor
JERS	Japanese Earth Resource Satellite
MATLAB	Matrix Laboratory (MathWorks' computing language)
MCF	Minimum Cost Flow (unwrapping method)
MLI	Multi-Look Image
MSP	GAMMA Modular SAR Processor
NASDA	National Space Development Agency of Japan
PRF	Pulse Repetition Frequency
PRI	SAR Precision Image
RMLI	Resampled Multi-Look Image
RSLC	Resampled Single-Look Complex Image
SAR	Synthetic Aperture Radar

SCOMPLEX	Short Complex (SLC format) pairs of 2-byte-short-integer
SLC	Single-Look Complex Image
SNR	Signal to Noise Ratio
SQR	Square root
Std Dev	Standard deviation
UTM	Universal Transverse Mercator
VV	Vertical polarization transmitted, Vertical polarization received
WInSAR	Western North America Interferometric SAR Consortium (online archive)

Appendix B – Satellite and Sensor Overview

ALOS

ALOS (Advanced Land Observing Satellite) has PALSAR (Phased Array type L-band SAR) aboard. Launched in January 2006, it has a high spatial resolution mode (10m) with a 70km swath (HH, VV, HH&HV, VV&VH) and a ScanSAR mode (HH, VV) with a 250-350km swath at 100m resolution.

- [Japan Aerospace Exploration Agency – ALOS](#)
- [NASDA - ALOS](#)

Envisat

ASAR (Advanced Synthetic Aperture Radar) is one of 10 instruments aboard the new ESA satellite Envisat (Environmental Satellite). Envisat was launched in March 2002 and has been operating in repeat cycles of 35 days. ASAR's operational modes allow up to 7 different image swaths with widths ranging from 105km to 56km. The image modes can be HH or VV polarization, incidence angles spanning from 15 to 45 degrees. Spatial resolution is on the order of 30m.

- [ESA – Envisat overview](#)
- [Envisat Instrument description](#)

ERS

ERS-1 (Earth Remote-Sensing Satellite) operates at C-Band (5.6cm), has VV polarization and a fixed incidence angle of 23 degree. The spacecraft provides a swath width of 100km. ESA launched ERS-1 in July 1991 with a variety of repeat cycles such as a 3-day, 35-day and a 336-day cycle. ERS-2 was launched in April 1995, allowing a 5 year tandem mission (1-day apart) until ERS-1 failed in March 2000.

- [European Space Agency ESA – ERS missions](#)
- [ESA – ERS overview](#)

JERS

JERS-1 (Japanese Earth Remote-Sensing Satellite) became operational in May 1992, has a fixed incidence angle of 35 degrees. It is a L-Band (24cm) radar with HH polarization. The SAR has 18m resolution with a 75km swath and its repeat orbit is 44 days. It was discontinued in October 1998. JERS-1 was re-entered in December 2001.

- [National Space Development Agency of Japan \(NASDA\)- Satellites](#)

RADARSAT

RADARSAT-1 (Radar Satellite) is a Canadian Space agency imaging radar that operates at C-Band (5.6cm). Launched in November 1995, RADARSAT-1 has HH polarization and its selectable image swaths range from 45km width at 8m resolution to 500km ScanSAR swath width at about 100m resolution. The incidence angle varies from 10 to 58 degrees and it completes an orbital cycle in 24 days. RADARSAT-2 is under development

and scheduled to launch in March 2007.

- [Alaska Satellite Facility – Data products](#)
- [Canadian Space Agency – Satellites](#)

SIR-C

SIR-C (Shuttle Imaging Radar) Space Shuttle Instrument was flown twice in April 1994. It operates at L-Band (3.5cm) and C-Band (5.8cm) with HH, VV, VH and HV polarization modes. Spatial resolution is about 30m. X-SAR was also on board, operating at X-band (3.1cm).

- [JPL – NASA - SIR-C](#)

Appendix C – Complimentary Information

1. Synthetic Aperture Radar (SAR)

Theory

Radar is an abbreviation for Radio Detection and Ranging. Radar systems send out modulated waveforms using antennas in order to transmit electromagnetic waves. Objects within a certain area will reflect part of the energy (radar returns or echoes) back to the radar. From these radar returns, the radar receiver then extracts information such as velocity and range (distance), angular position, and other identifying characteristics. Synthetic Aperture Radar (SAR) is an active side-looking imaging radar. It can only be utilized by a moving instrument over a relatively immobile target area. Penetrating clouds, smoke, haze and darkness, radar in general is almost completely independent of weather and time of day. This unique feature of all radar makes it an excellent complement to photographic and other optical and passive imaging data sources. Cultural and terrain targets on the ground respond characteristically to certain radar frequencies, enabling monitoring of topographic change.

The satellites being used for InSAR are all orbiting the Earth from north to south in a sun-synchronous pattern to maximize area coverage and minimize the time it takes to get a second acquisition of the same region. The positioning between the satellite and the sun is always the same, which means the area that is covered by the satellite is always being illuminated with the same incidence angle from the sun. The motion of the satellite then matches the motion of the sun across the sky. Satellites in sun-synchronous orbit cross the equator twelve times a day at the same time, making a 98.6° angle with the equator, i.e., 9.8° off north (polar orbits: 90° , aligned with north). Hence, the satellite passes the equator and every latitude at the same time each day.

The direction the satellite is traveling in is known as the along-track direction or azimuth. The across-track direction is called range. Effective across-track resolution is increased by combining accurate range measurements with post processing of the radar data that is done by using the Doppler-effect. The pass of the satellite can either be descending, coming from the northern hemisphere going south or ascending, going north. As mentioned above, the angle between the track and the north direction is about 8.6° . The SAR aboard the satellite is right-looking. Therefore one can get different viewing angles for the same region of interest, once illuminated from the “right” on a descending pass and from the “left” side on a ascending pass providing multiple aspect angles. Envisat, ESA's Environmental Satellite, e.g., can also be operated in different modes, which correspond to swath width, incidence angle and across-track resolution. The most coverage is obtained by operating in the wide swath mode, using ScanSAR (Scanning SAR) techniques, where ASAR (Advanced SAR aboard Envisat) has the capability to illuminate several sub-swaths by scanning it's antenna off-nadir (away from the point directly underneath the satellite) into different positions. Spatial resolution for ScanSAR is rather mediocre, at the order of 150 m at a swath width of 405 km. In image mode, ASAR generates high spatial resolution products (30 m) at swath widths of 56-105 km (Figure 51: [RADARSAT image swaths and ground coverage](#) PDF 880 KB).

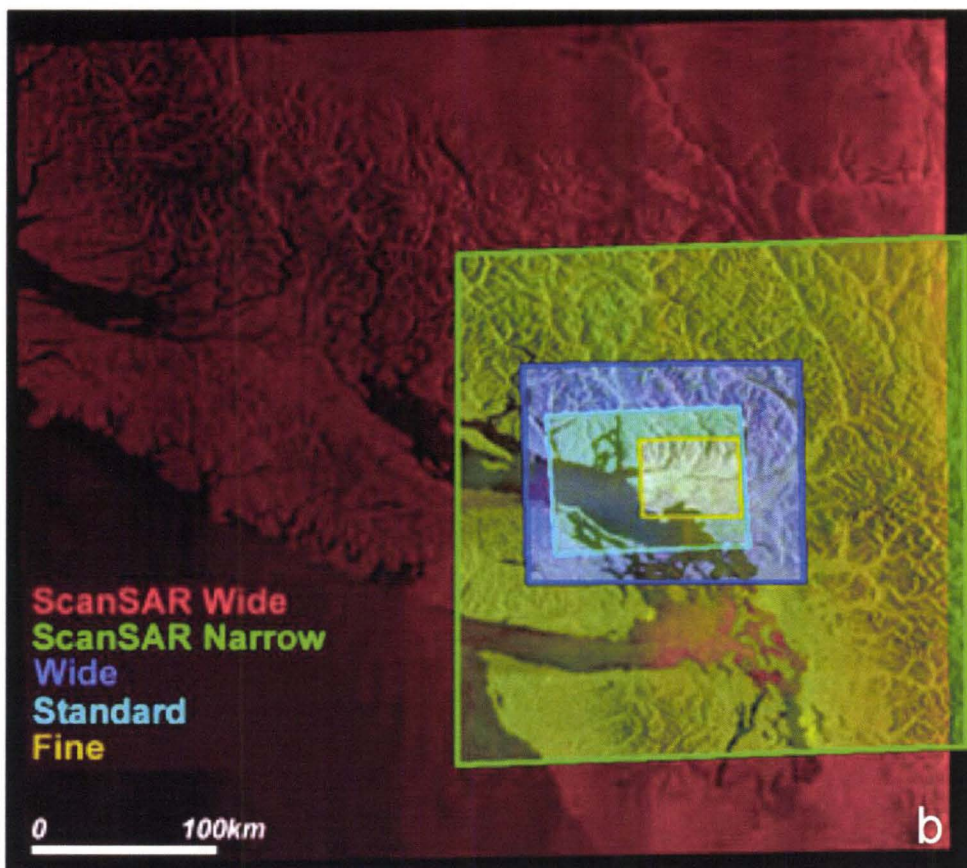
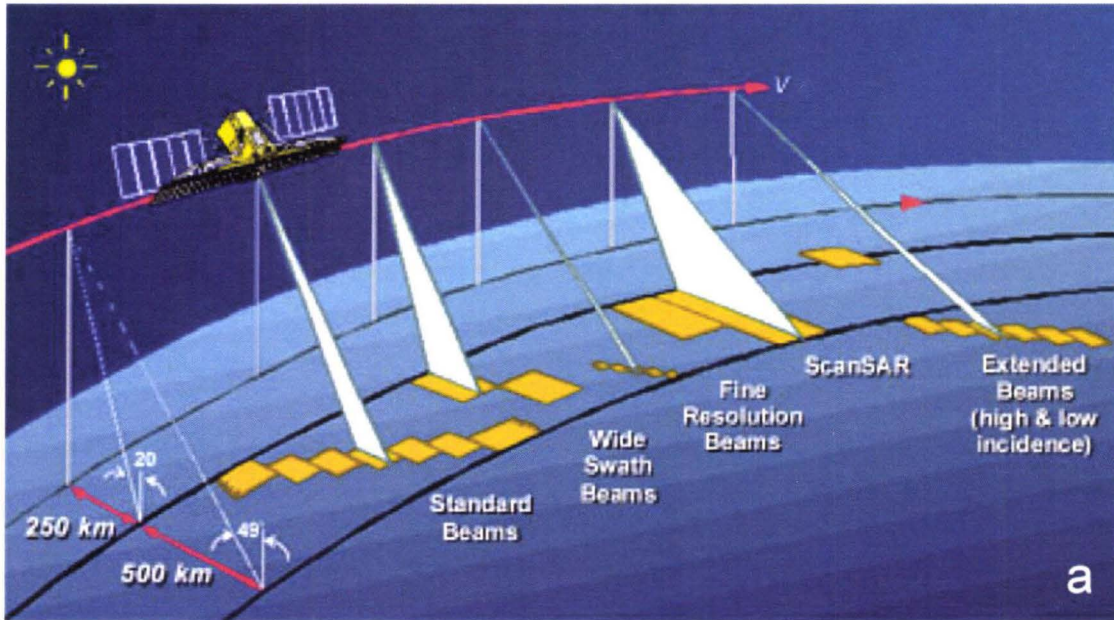


Figure 51: A satellites different operating modes (a) and their ground coverage (b). When operating in image mode, Envisat's seven swaths are similar to the standard beams in (a), where the scene closest to nadir has the largest area coverage (a, b: Canadian Space Agency).

Data Format

The following data description is solely for the ASAR instrument on the ESA satellite Envisat. It may be different for different satellites and sensors. The raw data received by the satellite is a serial data bit-stream that is recorded by a demodulator that recovers information content from the carrier wave. The demodulator output is sent to a band sub-system. After some further processing carried out by ESA, the data is then stored on High-Density Data Tapes. There are a few data formats being provided by ESA depending on the grade of processing. In addition, the data can be either vertically or horizontally polarized.

- ESA's ASAR handbook link: [Product summary by level](#)

Level 0 data is reformatted, time-ordered satellite raw data that is acquired at a high-rate and in a narrow swath and may be from one of seven swaths. The product name will start with *ASA_IM__0C** and can be directly fed into GAMMA. More information about the product type can be found at:

- ESA's ASAR handbook link: [Level 0 image mode](#)

By applying a range-Doppler algorithm and calibration data, Level 0 products are transformed to Level 1B baseline engineering products:

- ESA's ASAR handbook link: [Algorithms for Level B Products](#)

Some of the data in the PGF archive that was directly ordered from ESA is in Image Mode Single Look Complex format:

- ESA's ASAR handbook link: [Level 1 B SLC Products](#)
- [ESA's glossaries](#)

2. Interferometry Technique

The satellite travels in the along track direction which coincides with the velocity-vector direction, looking to the side and recording microwave echoes as a function of slant-range. Slant-range is the line-of-sight of the radar, also called look-vector. Successive echoes are recorded coherently, building the phase of the data stream. The imaged ground surface (3D) is projected into two dimensions: slant range and azimuth.

Traditionally, radar range is measured by timing the transmitted pulse precisely. Depending on the bandwidth of the ranging signal, accuracy is several meters. In InSAR however, the distance to the ground is measured using the phase information of the electromagnetic wave as well as the magnitude, increasing the accuracy to mm - cm resolution corresponding to 1% of the wavelength that can be resolved. The radar signal z is a complex number, in cartesian coordinated expressed: $z \equiv x + i y$ where the magnitude (or complex modulus) is defined as: $|z| = \sqrt{x^2 + y^2}$. The argument of the complex number is the phase $\phi = \arg(z) = \tan^{-1}(y/x)$. For polar coordinates: $z = r e^{i\phi}$, i.e., $r = |z|$. The backscatter phase recorded at the radar includes a propagation phase delay, a function of slant range ρ and wavelength λ : $\phi = -4\pi\rho/\lambda$. The total return from a resolution cell consists of the coherent sum of contributions from all elemental scatterers. The magnitude reflects the backscatter

intensity of many scatterers and is therefore referred to as the brightness. In contrast, the phase consists of propagation time plus the scatter term only known mod 2π .

The range information provides a sphere, a.k.a. the surface of constant range describing the time it takes a radar pulse to propagate to the target and back to the radar [Rosen, *et al.*, 2000]. The surface of constant Doppler is a cone about the velocity-vector (flight direction). The cone angle is proportional to the Doppler frequency. Intersecting the surface of constant range with the cone of constant Doppler leaves a hyperbola on the ground surface. A third measurement is necessary to resolve the 3D location. By imaging the same area from a slightly different viewing angle, hence intersecting two hyperbolas, the point location can exactly be calculated. The second hyperbola has approximately the same orientation, therefore only providing one intersection, rather than two if they are oriented opposite ways. There are two separate interferometry techniques that produce the third measurement: Single-track and repeat-track interferometry. In single-track interferometry the satellite's radar is equipped with two antennas. This allows an absolute precise measurement of the perpendicular baseline, the distance between the two antennas. This kind of interferometry is used to create DEMs, such as SRTM, using the fact that there is negligible temporal separation, i.e., no decorrelation and/or deformation. In contrast, repeat-track interferometry is performed by a single antenna SAR aboard a satellite, which images exactly the same area after completing its orbital cycle. For Envisat the orbital frequency is about 35 days.

To resolve the elevation and/or displacement information (depending on the temporal separation) the two SAR images (i.e., their phases) need to be interfered. This process involves superposing the two waves that are slightly "out-of-phase" received from two different look angles. The path difference is only known mod 2π and is proportional to the distance.

The SAR images must then be processed to SLC images (Single-look complex), co-registering the SAR image to $1/20^{\text{th}}$ of a pixel to prevent the loss of correlation. Once co-registered, an interferogram can be produced.

Appendix D – SAR Data Availability and Acquisition Tools

[EoliSA](#)

EoliSA is ESA's multi-mission catalog service that can be accessed either in form of a web client or as a standalone version ([Figure 52](#) PDF 2 MB). Downloaded on to any operating system, and if connected to Internet, EoliSA has access to the entire ESA data collection, including Envisat, ERS and third party mission products. It displays available scenes in a SLC quick look format along with the tracks and frames plotted on the area of interest. It also allows you to order the data right from ESA.

[Descw](#)

Descw is an off-line catalog and image browser very similar to EoliSA, also provided by ESA ([Figure 53](#) PDF 1.3 MB). It has the advantage to show multiple tracks at the same time for one area. It does not provide a quick look SLC image and its libraries need to be updated regularly. A newly adopted feature includes baselines estimates for Envisat data.

[WInSAR](#)

WInSAR is a web based archive of radar data that is shared by several universities such as Stanford, Scripps and Caltech ([Figure 54](#) PDF 710 KB). A username and password are necessary to download the data. It is easy to use, but some of the Envisat frame/date information can be incomplete. WInSAR also provides Envisat precise orbit files by day (DORIS). WInSAR will be implemented as a standing committee of UNAVCO, according to a vote of Consortium members this past winter.

[PGF Data Archive](#)

The PGF radar data archive can be found on the PGF internal web. An open office calc spreadsheet (similar to Excel) conversion to html was used to create the on-line Envisat archive, in order to easily update the sheets and re-produce the web page ([Figure 55](#) PDF 2.5 MB).

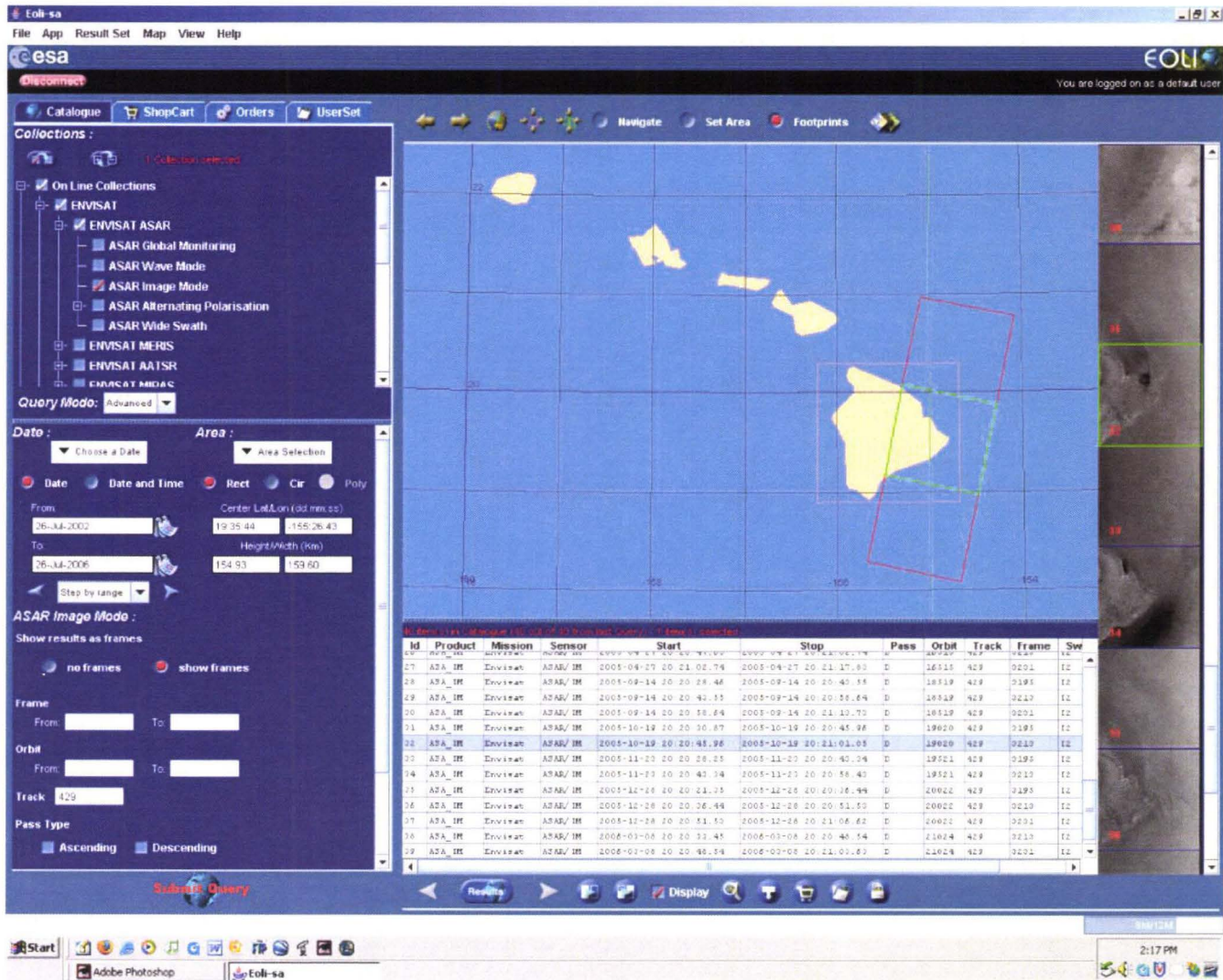


Figure 52: EoliSA software (ESA) used to display Envisat track 429 and frame 3213 (green box).

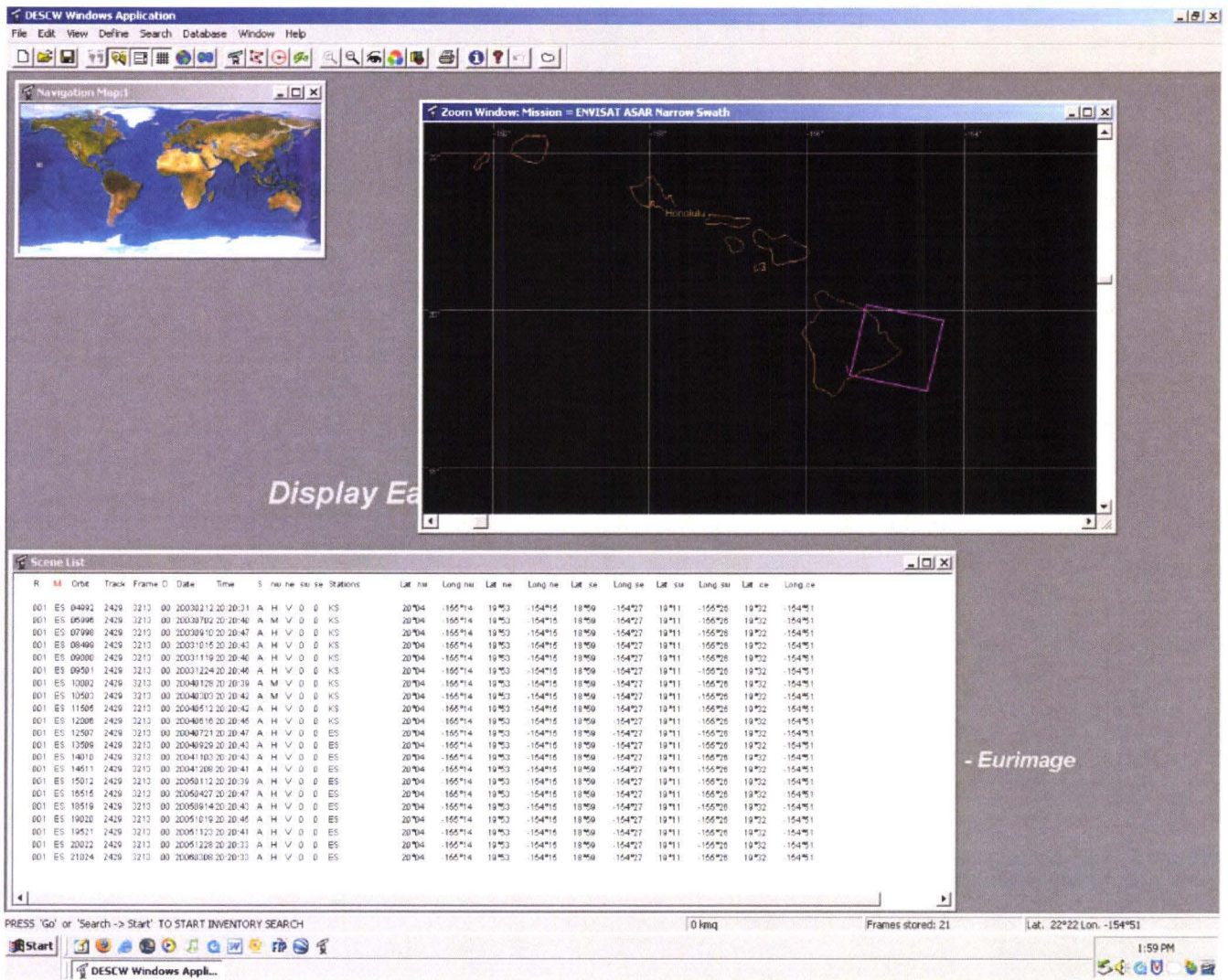



Figure 53: Descw software (ESA) used to display Envisat track 429 and frame 3213 (magenta box).



WinSAR

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Search Results





Found 31 results.
Displaying results 1 to 31.





To download a file, click on the desired location/filetype button in the "location" column.





Site key:





Caltech Scripps Stanford Peer-to-Peer

ERS zip archive    

ERS leader file    

ERS VDF file    

ERS raw data file    

ENV baq data file    













#	preview	sensor	track	frame	orbit	date	latitude	longitude	b_perp	status	location list	comments
1	map	ENV1	429	3231	20022	2005-12-28 20:20:52	19.38	205.02	0	Archived	Std DL:  Fast DL: 	---
2	map	ENV1	429	3231	20022	2005-12-28 20:20:52	19.38	205.02	0	Archived	Std DL:  Fast DL: 	---
3	map	ENV1	429	3231	14511	2004-12-08 20:20:57	19.45	205.04	0	Archived	Std DL:  Fast DL: 	---
4	map	ENV1	429	3231	14511	2004-12-08 20:20:57	19.45	205.04	0	Archived	Std DL:  Fast DL: 	---
5	map	ENV1	429	3231	14511	2004-12-08 20:20:57	19.45	205.04	0	Archived	Std DL:  Fast DL: 	---
6	map	ENV1	429	3231	13509	2004-09-29 20:20:58	19.45	205.04	0	Archived	Std DL:  Fast DL: 	---

Figure 54: Online WinSAR archive. Search results are for Envisat track 429 and frame 3231.

Sheet 8: IS2 T429

Swath	Track	Phase	Pole	Orbit	Start Date/Time	Area coverage	ENVISAT Swath (WENTAR)	ASA_ID (ENVPD) ID (Raw data)
D	429	1195	D	7996	20030930 20:20:33	North of P3211	N/A	
D	429	1195	D	8499	20031015 20:20:29		N/A	
D	429	1195	D	9000	20031119 20:20:27		N/A	
D	429	1195	D	9501	20031224 20:20:30		N/A	
D	429	1195	D	11309	20040929 20:20:21		N/A	
D	429	1195	D	14310	20041103 20:20:31		N/A	
D	429	1195	D	14511	20041209 20:20:28		N/A	
D	429	1195	D	15012	20050112 20:20:23		N/A	
D	429	1195	D	16313	20050427 20:20:31		N/A	
D	429	1195	D	18514	20050914 20:20:24		N/A	
D	429	1195	D	19020	20031019 20:20:31		N/A	
D	429	1195	D	19521	20031123 20:20:21		N/A	
D	429	1195	D	20022	20031228 20:20:21		N/A	
D	429	1210	D	4992	20030212 20:20:33	SLC	2429_3231_4992	IC00030212_202043_000000152013_00429_04992_0020
D	429	1210	D	6996	20030702 20:20:41	SLC	2429_3231_6996	IC00030702_202040_000000152017_00429_06996_0123
D	429	1210	D	7996	20030930 20:20:44	SLC	2429_3231_7996	IC00030930_202047_000000152019_00429_07996_0311
D	429	1210	D	8499	20031015 20:20:44	SLC	2429_3231_8499	IC00031015_202043_000000152020_00429_08499_0372
D	429	1210	D	9000	20031119 20:20:42	SLC	2429_3231_9000	IC00031119_202041_000000152021_00429_09000_0306
D	429	1210	D	9501	20031224 20:20:43	SLC	2429_3231_9501	IC00031224_202044_000000152022_00429_09501_0144
D	429	1210	D	10002	20040128 20:20:41	SLC	2429_3231_10002	IC00040128_202040_000000152023_00429_10002_1913
D	429	1210	D	10303	20040303 20:20:42	SLC	2429_3231_10303	IC00040303_202041_000000152024_00429_10303_2666
D	429	1210	D	11303	20040512 20:20:42	SLC	2429_3231_11303	IC00040512_202041_000000152026_00429_11303_1297
D	429	1210	D	12008	20040616 20:20:48	SLC	2429_3231_12008	IC00040616_202046_000000152027_00429_12008_2062
D	429	1210	D	12507	20040721 20:20:48	SLC	2429_3231_12507	IC00040721_202047_000000152028_00429_12507_0841
D	429	1210	D	13309	20040929 20:20:43	SLC	2429_3231_13309	IC00040929_202047_000000152030_00429_13309_2931
D	429	1210	D	14010	20041103 20:20:46	SLC	2429_3231_14010	IC00041103_202043_000000152031_00429_14010_4033
D	429	1210	D	14511	20041209 20:20:43	SLC	2429_3231_14511	IC00041209_202042_000000152032_00429_14511_4940
D	429	1210	D	15012	20050112 20:20:46	SLC	N/A	IC00050112_202047_000000152033_00429_15012_3832
D	429	1210	D	16513	20050427 20:20:43	SLC	2429_3231_16513	IC00050427_202047_000000152034_00429_16513_9264
D	429	1210	D	18514	20050914 20:20:44	SLC	N/A	IC00050914_202045_000000152040_00429_18514_2347
D	429	1210	D	19020	20031019 20:20:44	SLC	N/A	IC00051019_202046_000000152041_00429_19020_3474
D	429	1210	D	19521	20031123 20:20:43	SLC	N/A	IC00051123_202047_000000152041_00429_19521_4146
D	429	1210	D	20022	20031228 20:20:34	SLC	2429_3231_20022	IC00051228_202037_000000152041_00429_20022_3768
D	429	1210	D	21024	20060308 20:20:33	SLC	N/A	IC00060308_202015_000000152045_00429_21024_0034
D	429	1231	D	15012	20050112 20:20:55	South of P3211	N/A	
D	429	1231	D	16513	20050427 20:21:03		N/A	
D	429	1231	D	18514	20050914 20:20:59		N/A	
D	429	1231	D	20022	20031228 20:20:52		N/A	
D	429	1231	D	21024	20060308 20:20:49		N/A	

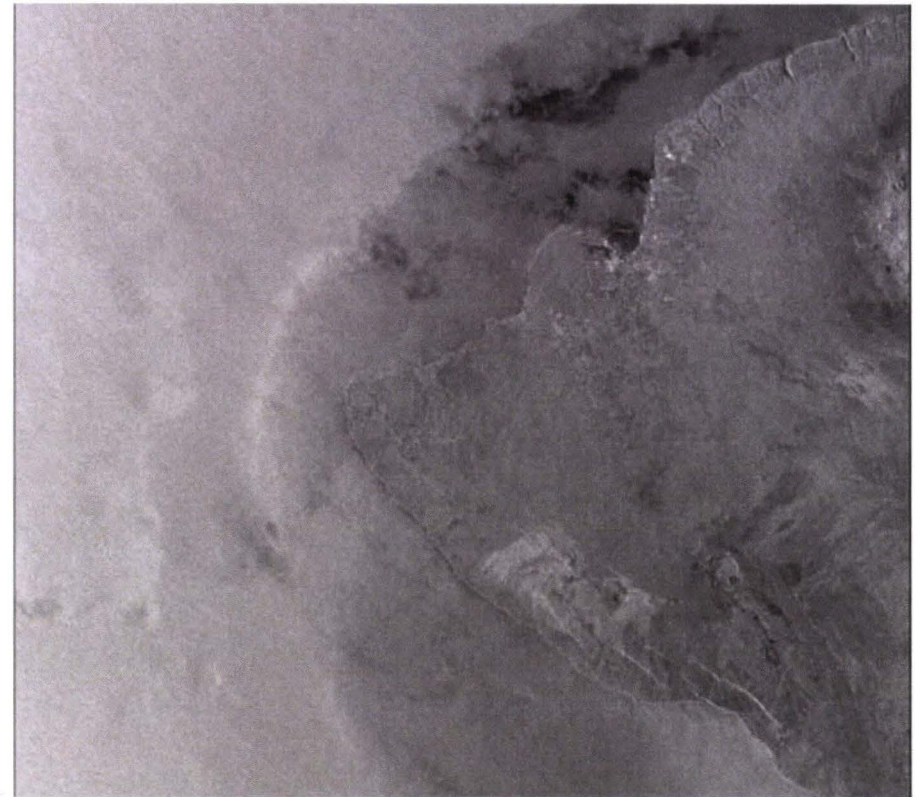


Figure 55: Online PGF archive. Search results are for Envisat track 429 including corresponding SLC image.

Appendix E – [Index of Commands and their Syntax](#)**Appendix F** – **File Access Information*****DORIS precise orbits***

Local: PGFSAR://mnt/scsi/DATA/DORIS/vor
Remote: ESA's EOHelp desk
ftp-ops.fr.envisat.esa.int
IP address: 193.50.83.2
User name: inquire from [Ben Brooks](#)
Password: inquire from [Ben Brooks](#)

WInSAR

Remote: <http://winsar.unavco.org/main.php>
User name: inquire from [Ben Brooks](#)
Password: inquire from [Ben Brooks](#)

External calibration files

Local: PGFSAR://mnt/scsi/DATA/INSAR/ASA_XCA
Remote: http://earth.esa.int/services/auxiliary_data/asar/

Instrument characterization files

Local: PGFSAR://mnt/scsi/DATA/INSAR/ASA_INS
Remote: http://earth.esa.int/services/auxiliary_data/asar/

IPTA scripts

Local: PGFSAR://usr/local/GAMMA_SOFTWARE-20051219/IPTA_v1.2/scripts

MATLAB scripts

Local: PGFSAR://usr/local/matlab/toolbox/pgf_tools

Shuttle Radar Topography Mission (SRTM)

Local: PGFSAR://mnt/scsi/DATA/TOPO
Remote: <http://srtm.usgs.gov/>

Appendix G – [Index of Figures and Linked Files](#)

Appendix E – Index of Commands and their Syntax

Note: This list only contains commands and scripts used in the PGF User Guide, for a complete listing of commands please refer to the “Reference Manual” sections of the GAMMA software documentation.

[MSP](#) – [ISP](#) – [IPTA](#) – [DIFF/GEO](#) – [DISP](#) – [LAT](#)

[adf](#)

<int> <sm> <cc> <width> <alpha> [nfft] [cc_win] [step] [loff] [nlines] [wfrac]

Input parameters

int	interferogram (fcomplex)
width	number of samples/line
alpha	exponent for non-linear filtering (default = 0.5)
nfft	filtering FFT window size (default = 32, 2**N, 8 --> 512)
cc_win	coherence parameter estimation window size (odd, default = 7, max = 15)
step	processing step (default = nfft/8)
loff	offset to starting line to process (default = 0)
nlines	number of lines to process (default = 0: to end of file)
wfrac	minimum fraction of points required to be non-zero in the filter window (default = 0.7)

Output parameters

sm	smoothed interferogram (fcomplex)
cc	coherence derived from smoothed interferogram, (enter - to avoid writing out cc) (float)

[ASAR_IM_proc](#)

<L0> <INS> <SAR_par> <PROC_par> <raw> <ant_gain> [loff] [nl]

Input parameters

L0	Level 0 image mode ASAR data file (ASA_IM_0P)
INS	ASAR instrument characterization file (ASA_INS_AX)
ant_gain	antenna pattern file name (derived using program ASAR_XCA)
loff	offset lines to first line to extract (default = 0, enter - for default)
nl	number of lines to extract (default = to end of input, enter - for default)

Output parameters

SAR_par	MSP SAR sensor parameter file (example)
PROC_par	MSP processing parameter file, (example p<orbit>.slc.par)
raw	byte aligned 8-bit I/Q raw data

[ASAR_pre_proc](#)

<ASAR_list> <DORIS_path> <RAW_dir> <log> <proc_list> <mode> [keyword] [value]

Input parameters

ASAR_list	see above (ASAR_pre_list)
DORIS_path	path to DORIS orbit data (i.e. /mnt/scsi/DATA/DORIS/vor)
RAW_dir	directory for output unpacked raw data files and processing parameter files

	<i>/raw</i> is generated by the script ASAR_pre_proc
mode	processing mode (see above)
keyword	(mode 4 only) search keyword in the MSP processing parameter file (example: doppler_polynomial)
value	(mode 4 only) new value delimited by double quotes (example: "317.0 0. 0. 0.")
<u>Output parameters</u>	
log	ASAR_pre_proc processing log file
proc_list	processing list for use by ASAR_proc_all (the next script to run)

ASAR_proc_all_no_autof

<proc_list> <RAW_dir> <rc_dir> <SLC_dir> <MLI_dir> <rlks> <azlks> <slc_format> [az_patch] [autof_min]

Input parameters

proc_list	processing list generated by ASAR_pre_proc (8 columns): <ol style="list-style-type: none"> 1. scene_identifier (example: 19960816) 2. offset in echoes to start of processed data (enter - for default) 3. number of echoes to process (enter - for default) 4. range offset in samples (enter - for default) 5. number of range samples to process (enter - for default) 6. Doppler centroid for scene (Hz) 7. Doppler slope for scene (Hz/m) 8. azimuth processing bandwidth fraction
RAW_dir	data directory containing fixed ASAR raw data files
rc_dir	directory to temporally store intermediate range compressed data (example: <i>/tmp</i>)
SLC_dir	directory to store output SLC data (example: <i>/slc</i>)
MLI_dir	directory to store multi-look intensity (MLI) files derived from the SLC data (example: <i>/mli_1_5</i>)
rlks	number of range looks to generate MLI images (nominal: 1)
azlks	number of azimuth looks to generate MLI images (nominal: 5 for IS2)
slc_format	desired output SLC image format (0 = fcomplex, 1 = scomplex)
az_patch	azimuth patch size (default = 6144)
autof_min	minimum SNR threshold for autofocus, 0.0 for no autofocus (default = 5) this option does not work

ASAR_XCA

<ASA_XCA> <antenna> [swath] [pol]

Input parameters

ASA_XCA	ASAR external calibration data file (binary)
swath	ASAR image swath (IS1, IS2, ... IS7; SS1, SS2, ... SS5)
pol	polarization (HH, VV, HV, VH)

Output parameters

antenna	1-way antenna gain pattern file or '-' (if not provided)
-------------------------	--

or 'all' to generate all ASAR antenna diagrams

[ave_image](#)

<MLI_tab> <width> <ave> [start] [nlines] [pixav_x] [pixav_y] [zero_flag]

Input parameters

[MLI_tab](#) text file containing list of names of co-registered MLI images in column 1
(cc_list generated by [mk_sp_all](#))

width number of samples/line

start starting line (default = 1)

nlines number of lines to process (default - : entire file)

pixav_x number of pixels to average in width (default = 1)

pixav_y number of pixels to average in height (default = 1)

zero_flag default: zero_flag = 0 => 0.0 interpreted as missing value

Output parameters

ave average of input data files (float)

[az_proc_dop2d](#)

<SAR_par> <PROC_par> <rc_data> <SLC> [az_patch] [SLC_format] [cal_fact] [SLC_type] [kaiser] [npatch]

Input parameters

SAR_par MSP SAR sensor parameter file

PROC_par MSP processing parameter file

rc_data input range compressed data file

az_patch along-track azimuth patch size (range lines): (2**N 2048, 4096, 8192...)

SLC_format output SLC format flag (default=from MSP processing parameter file)

0: FCOMPLEX (pairs of 4-byte float)

1: SCOMPLEX (pairs of 2-byte short integer)

cal_fact radiometric calibration factor [dB] (default = 0.0)

Proposed factors [dB] for absolute calibration:

ERS1 1991-1996: -10.3 dB (49.7 dB for SCOMPLEX output format)

ERS1 1997-2000: -12.5 dB (47.5 dB for SCOMPLEX output format)

ERS2: -2.8 dB (57.2 dB for SCOMPLEX output format)

JERS: -22.1 dB (37.9 dB for SCOMPLEX output format)

SLC_type output parameter type (default = 0)

0: sigma0 ($\sqrt{re^2 + im^2}$) => sigma0)

1: gamma0 ($\sqrt{re^2 + im^2}$) => gamma0) gamma0 = sigma0 / cos(inc)

kaiser Kaiser window parameter beta for the azimuth reference function (default: 2.120)

npatch number of along-track patches to process

Output parameters

[SLC](#) Single-Look Complex image (FCOMPLEX or SCOMPLEX format)

[azsp_IQ](#)

<SAR_par> <PROC_par> <signal_data> <spectrum> [loff] [roff] [nsub] [ambig_flg] [namb]

Input parameters

SAR_par	SAR sensor parameter file
PROC_par	MSP processing parameter file
signal_data	input raw I/Q format SAR data
loff	number of lines offset to start of estimation window (default = 0)
roff	range samples offset to center of estimation window (enter - for default = center_swath)
nsub	number of azimuth sub-apertures for spectrum estimation (default = 12)
ambig_flg	Doppler ambiguity resolution mode 0 = add multiples of PRF specified by the namb command line parameter 1 = use unambiguous Doppler centroid estimate from the PROC_par file (default)
namb	number of multiples of the PRF to add to the ambiguous Doppler estimate (default = 0)

Output parameters

[spectrum](#) azimuth spectrum ([text](#) format for plotting)

base_calc

<SLC_tab> <SLC_par> <gr_file> <bperp_file> <itab> [itab_type] [bperp_min] [bperp_max] [delta_T_min]
[delta_T_max]

Input parameters

SLC_tab	two column list of SLC filenames and SLC parameter filenames (including paths) (ascii)
SLC_par	reference SLC image parameter filename (include path)
itab_type	itab type (0: single reference, 1: all pairs)
bperp_min	minimum magnitude of bperp (m) (default = all, enter - for default)
bperp_max	maximum magnitude of bperp (m) (default = all, enter - for default)
delta_T_min	minimum number of days between passes (default = 0, enter - for default)
delta_T_max	maximum number of days between passes

Output parameters

gr_file	xmgrace plot file name, (to redraw: xmgrace gr_file)
bperp_file	file containing a list of bperp and delta_T for interferogram pairs in the itab
itab	interferogram table

base_ls

<SLC_par> <OFF_par> <gcp_ph> <baseline> [ph_flag] [bc_flag] [bn_flag] [bcdot_flag] [bndot_flag] [bperp_min]
[SLC2R_par]

Input parameters

SLC_par	ISP parameter file of the reference SLC
OFF_par	ISP interferogram/offset parameter file
gcp_ph	ground control point heights + extracted unwrapped phase (text format)
baseline	(input/output) baseline parameter file
ph_flag	restore range phase ramp (default = 0: do not restore 1: restore)
bc_flag	cross-track baseline component estimate

bn_flag	(0: orbit derived, 1: estimate from data, default = 1) normal baseline component estimate
bcdot_flag	(0: orbit derived, 1: estimate from data, default = 1) cross-track baseline rate estimate (0: orbit derived, 1: estimate from data, default = 1)
bndot_flag	normal baseline rate estimate (0: orbit derived, 1: estimate from data, default = 0)
bperp_min	minimum perpendicular baseline required for L.S estimation (m, default = 10.0)
SLC2R_par	parameter file of resampled SLC, required if SLC-2 frequency differs from SLC-1
<u>base_orbit</u>	
<SLC1_par> <SLC2_par> <baseline>	
<u>Input parameters</u>	
SLC1_par	SLC-1 ISP image parameter file
SLC2_par	SLC-2 ISP image parameter file
<u>Output parameters</u>	
baseline	baseline file (text format, enter - for none)
<u>base_orbit_pt</u>	
<pSLC_par> <itab> <rec_num> <pbase>	
<u>Input parameters</u>	
pSLC_par	SLC/MLI parameter stack (binary)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
rec_num	record number to process (default - : all records)
<u>Output parameters</u>	
pbase	baseline parameter stack (binary)
<u>cc_wave</u>	
<interf> <pwr1> <pwr2> <corr> <width> <bx> <by> <wgt_flag> <xmin> <xmax> <ymin> <ymax>	
<u>Input parameters</u>	
interf	complex interferogram (complex coherence) filename
pwr1	intensity image of first scene (or -)
pwr2	intensity image of second scene (or -)
width	number of samples/row
bx	estimator window size (columns) (default = 5.0)
by	estimator window size (rows) (default = 5.0)
wgt_flag	weighting function flag (default: 0 = uniform, 1 = linear, 2 = Gaussian)
xmin	starting range pixel offset (default = 0)
xmax	last range pixel offset (default = width - 1)
ymin	starting azimuth row offset, relative to start (default = 0)
ymax	last azimuth row offset, relative to start (default = nlines - 1)
<u>Output parameters</u>	
corr	estimated degree of coherence filename

NOTICE: Instead of the pwr1 and pwr2 files - may be provided to select alternative coherence estimation method based exclusively on complex interferograms

[create_dem_par](#)

<DEM_par> [SLC_par] [terra_alt] [delta_y] [delta_x]

Input parameters

[DEM_par](#) (input/output) DIFF/GEO DEM parameter file
 SLC_par ISP SLC/MLI image parameter file
 terra_alt nominal terrain altitude, default: 0.0 m
 delta_y DEM y line spacing for new DEM_par file
 geographical coordinates (EQA projection latitude) default: -5.0000e-04 deg
 all other map projections default: -50.0 m (use information from header file)
 delta_x DEM x sample spacing for new DEM_par file
 geographical coordinates (EQA projection) default: 5.0000e-04 deg
 all other map projections default: 50.0 m

[create_diff_par](#)

<PAR1> <PAR2> <DIFF_par> [PAR_type]

Input parameters

PAR1 parameter file 1 (see PAR_type option)
 PAR2 parameter file 2 (or - if not provided)
 PAR_type PAR1 and PAR2 parameter file type:
 0: OFF_par ISP interferogram/offset parameter files (default)
 1: SLC/MLI_par ISP SLC/MLI parameter files
 2: DEM/MAP_par DIFF&GEO DEM/MAP parameter files

Output parameters

DIFF_par differential interferometry parameter file

[create_offset](#)

<SLC1_par> <SLC2_par> <OFF_par> [offset_algorithm] [rlks] [azlks]

Input parameters

SLC1_par SLC-1/MLI-1 ISP image parameter filename (reference)
 SLC2_par SLC-2/MLI-2 ISP image parameter filename
 offset_algorithm offset estimation algorithm:
 1: intensity cross-correlation (default)
 2: fringe visibility
 rlks number of interferogram range looks (default: 1)
 azlks number of interferogram azimuth looks (default: 1)

Output parameters

OFF_par ISP offset/interferogram parameter file

[data2pt](#)

<f_in> <par_in> <plist> <SLC_par> <pdata> <rec_num> <type>

Input parameters

f_in	2-D image data file (various types supported)
par_in	SLC/MLI parameter file of 2-D image data file
plist	point list for sampling (int)
SLC_par	SLC parameter file of point list coordinates
rec_num	record number in output point data stack (starting with 1)
type	data type (0: fcomplex, 1: scomplex, 2: float, 3: int, 4: short, 5: byte, 6: SUN/BMP raster)

Output parameters

pdata	point data stack (same type as f_in)
-------	--------------------------------------

def_mod_pt

<plist> <pmask_in> <pSLC_par> <ppos> <itab> <pbase> <bflag> <pdiff> <pdiff_type> <np_ref> <pres> <pdh>
 <pdef> <punw> <psigma> <pmask_out> [dh_max] [def_max] [sigma_max] [model] [pdh_err] [pdef_err] [ppc_err]
 [bmax] [dtmax]

Input parameters

plist	point list (int)
pmask_in	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters (binary)
ppos	point data stack of interpolated point target positions (fcomplex, enter - to use plist coordinates)

itab_selection

itab_selection	table associating interferogram stack records with pairs of SLC stack records (ascii)
pbase	stack of baseline parameters (binary)
bflag	baseline flag (0: initial baseline, 1: precision baseline)
pdiff	point data stack of differential interferograms (fcomplex or float)
pdiff_type	type of pdiff (0: float (unwrapped), default= 1: fcomplex)
np_ref	phase reference point number (beginning from 0)
dh_max	maximum height correction for initial fit (m, enter - for default: 60.0)
def_max	maximum deformation rate difference for initial fit (m/year, enter - for default: 2.0e-02)
sigma_max	threshold for phase standard deviation to set mask to valid: (enter - for default: 1.200)
model	phase model (see above)
bmax	maximum perpendicular baseline (m) considered, (-1 = all records, default = -1)
dtmax	maximum time interval (days) considered, (-1 = all records, default = -1)

Output parameters

pres	point data stack of residual unmodeled phase (float, = atmospheric phase, non-linear deformation, baseline error)
pdh	point data stack of height correction value (m, float, enter - for none)
pdef	point data stack of linear deformation rate (m/year, float, enter - for none)
punw	point data stack of unwrapped phase of pdiff (float, enter - for none)
psigma	point data stack of phase standard deviation from fit (float, enter - for none)
pmask_out	point data stack of mask values indicating accepted points (uchar, enter - for none)

pdh_err	estimated uncertainty in the height correction (m, float, enter - for none)
pdef_err	estimated uncertainty in linear deformation rate (m/year, float, enter - for none)
ppc_err	estimated uncertainty in phase constant a0 (radians, float, enter - for none)
<u>dem_trans</u>	
<DEM1_par> <DEM1> <DEM2_par> <DEM2> [lat_ovr] [lon_ovr] [Datum_shift] [lookup_table]	
<u>Input parameters</u>	
DEM1_par	DEM parameter file DEM1
DEM1	input DEM (Digital Elevation Model) file
DEM2_par	DEM parameter file of output DEM (transformed coordinates)
lat_ovr	latitude DEM/MAP oversampling factor (default = 2.0)
lon_ovr	longitude DEM/MAP oversampling factor (default = 2.0)
	lat_ovr and lon_ovr are only considered if DEM2_par does not exist
Datum_shift	Datum shift flag: 0: transformed DEM heights not corrected for datum shift 1: transformed DEM heights corrected for datum shift (default = 1)
<u>Output parameters</u>	
DEM2	DEM in transformed coordinates
lookup_table	complex-valued lookup table (DEM2 -> DEM1)
NOTE: if DEM2_par does not exist, it is created using the same projection parameters as DEM1	
<u>dis_ipta</u>	
<plist> <pmask> <pSLC_par> <ppos> <itab> <pbase> <bflag> <pdiff> <pdiff_type> <ras> [dh_max] [def_max] [model] [bmax] [dtmax] [mag] [win_sz]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters (binary)
ppos	point data stack of interpolated point target positions (fcomplex, enter - to use plist coordinates)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
pbase	stack of baseline parameters (binary)
bflag	baseline flag (0: initial baseline, 1: precision baseline)
pdiff	point data stack of differential interferograms (fcomplex or float)
pdiff_type	type of pdiff (0: float (unwrapped), default = 1: fcomplex)
ras	raster reference image with point locations marked (SUN *.ras, or BMP *.bmp)
dh_max	maximum height correction for initial fit (m, enter - for default: 60.0)
def_max	maximum deformation rate difference for initial fit (m/year, enter - for default: 5.0e-03)
model	phase model
max_bl	max. baseline (m), only points with baseline lengths < max_bl are considered
bmax	maximum perpendicular baseline (m) considered, (-1 = all records, default = -1)

dtmax	maximum time interval (days) considered, (-1 = all records, default = -1)
mag	zoom magnification factor (default = 3)
win_sz	zoom window size before magnification (default = 128)
dis2ras	
<ras1> <ras2> [mag] [win_sz]	
<u>Input parameter</u>	
ras1	image 1 SUN raster *.ras, or BMP *.bmp format
ras2	image 2 SUN raster *.ras, or BMP *.bmp format
	image depth (bits/pixel) of image-1 and image-2 must be identical
mag	zoom magnification factor (default = 3)
win_sz	zoom window size before magnification (default = 120)
disdem_par	
<DEM> <DEM_par> [start] [nlines] [exaggerate] [theta0] [phi0]	
<u>Input parameters</u>	
DEM	Digital elevation model (float or short as specified in DEM/MAP_par)
DEM_par	DIFFFF/GEO DEM parameter file
start	starting line of DEM (default = 1)
nlines	number of lines to display (default = 0: to end of file)
exaggerate	relief exaggeration factor to increase contrast of display (default = 2.0)
theta0	illumination elevation angle in degree (default = 45.)
phi0	illumination orientation angle in degree (default = 135.) (0: right, 90: top, 180: left, 270: bottom)
disdt_pwr24	
<data> <pwr> <width> [start_data] [start_pwr] [nlines] [cycle] [scale] [exp]	
<u>Input parameters</u>	
data	real valued image (float, e.g. deformation rate, terrain height)
pwr	intensity image (float)
width	number of samples/row of data and pwr
start_data	starting line of data (default = 1)
start_pwr	starting line of pwr (default = 1)
nlines	number of lines to display (default = 0: to end of file)
cycle	data value per color cycle (default = 1.0000e-02)
scale	pwr display scale factor (default = 1.)
exp	pwr display exponent (default = .35)
dishgt	
<hgt> <pwr> <width> [start_hgt] [start_pwr] [nlines] [m_cycle] [scale] [exp]	
<u>Input parameters</u>	
hgt	height image (float)
pwr	intensity image (float, enter - if not available)

width	samples per row of hgt and pwr
start_hgt	starting line of hgt (default = 1)
start_pwr	starting line of pwr (default = 1)
nlines	number of lines to display (default = 0: to end of file)
m_cycle	meters per color cycle (default = 160.)
scale	display scale factor (default = 1.)
exp	display exponent (default = .35)
dispmap	
<unw> <hgt> <SLC_par> <OFF_par> <disp_map> [mode] [loff] [nlines]	
<u>Input parameters</u>	
unw	unwrapped phase (float)
hgt	height map (float) or - to select constant reference height 0.0
SLC_par	SLC parameter file of reference SLC
OFF_par	ISP/offset parameter file
mode	flag indicating displacement mode: default = 0: along look vector [m] (+ : towards sensor) 1: vertical displacement [m] (+ : increasing height) 2: horizontal displacement [m] (+ : decreasing ground range)
loff	number of lines offset to starting line (default = 0)
nlines	number of lines to process (default: 0 = entire file)
<u>Output parameters</u>	
disp_map	displacement map
dispmap_pt	
<plist> <pmask> <pSLC_par> <itab> <punw> <phgt> <pdisp> [mode]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters (binary)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
punw	point data stack of unwrapped interferometric phases for each point and each layer in the stack (float)
phgt	point data stack of terrain heights (float) (enter - for none)
mode	flag indicating displacement mode: default = 0: along look vector [m] (+ : towards sensor) 1: vertical displacement [m] (+ : increasing height) 2: horizontal displacement [m] (+ : decreasing ground range)
<u>Output parameters</u>	
pdisp	displacement (m) for each point and each layer in the stack(binary)
disras	

<ras> [mag] [win_sz]

Input parameters

ras SUN raster *.ras, or BMP *.bmp format image
 mag zoom magnification factor (default = 3)
 win_sz zoom window size before magnification (default = 120)

disras_dem_par

<ras> <DEM_par> [mag] [win_sz]

Input parameters

ras raster image with same dimensions as in DEM_par (SUN *.ras, or BMP *.bmp format)
 DEM_par DIFF DEM/MAP parameter file
 mag zoom magnification factor (default = 3)
 win_sz zoom window size before magnification (default = 120)

DORIS_proc

<PROC_par> <DOR> [nstate]

Input parameters

PROC_par (input/output) MSP processing parameter file ([example](#) p<orbit>.slc.par)
 DOR ASAR DORIS data file (DOR_VOR_AXVF)
 nstate number of state vectors to extract (enter - for default: 9)

extract_gcp

<DEM_rdc> <OFF_par> <GCP> <nr> <naz> [mask]

Input parameters

DEM_rdc DEM in range-Doppler coordinates
 OFF_par ISP offset/interferogram parameter file
 nr number of GCP points in range (default = 32)
 naz number of GCP points in azimuth (default = 32)
 mask mask image, output set to 0.0 if mask = 0 (SUN raster or BMP format)

Output parameters

GCP GCP height data file (text format)

gc_map

<SLC_par> <OFF_par> <DEM/MAP_par> <DEM> <DEM/MAP_seg_par> <DEM_seg> <lookup_table> [lat_ovr]
 [lon_ovr] [sim_sar] [u] [v] [inc] [psi] [pix] [ls_map] [frame] [ls_mode] [r_ovr]

Input parameters

SLC_par ISP SLC or MLI image parameter file (slant range geometry)
 OFF_par ISP offset/interferogram parameter file (enter - if geocoding SLC or MLI data)
 DEM/MAP_par DEM/MAP parameter file
 DEM DEM data file (or constant height value)
 DEM/MAP_seg_par (input/output) DEM/MAP segment parameter file used for geocoding

NOTE: if this file exists, then the bounds of the DEM segment used for geocoding

are read from the parameter file, otherwise the bounds are estimated using the SLC parameters and state vectors, and written to the new parameter file

lat_ovr	latitude DEM over-sampling factor (default = 2)
lon_ovr	longitude DEM over-sampling factor (default = 2)
frame	number of DEM pixels to add around area covered by SAR image (default = 8)
ls_mode	output lookup table values in regions of layover, shadow, or DEM gaps 0: set to (0., 0.), 1: linear interpolation across these regions (default), 2: actual value, 3: nn-thinned
r_ovr	range over-sampling factor for nn-thinned layover/shadow mode (ls_mode = 3) (default = 2.0)

Output parameters

DEM_seg	segment of DEM used for geocoding interpolated if lat_ovr > 1, or lon_ovr > 1
lookup_table	geocoding lookup table
sim_sar	simulated SAR backscatter image (in DEM geometry)
u	zenith angle of surface normal vector n (angle between z and n)
v	orientation angle of n (between x and projection of n in xy plane)
inc	local incidence angle (between surface normal and look vector)
psi	projection angle (between surface normal and image plane normal)
pix	pixel area normalization factor
ls_map	layover and shadow map (in map projection)

NOTE: enter - as filename to avoid creation of the corresponding output file

gc_map_fine

<gc_in> <width> <DIFF_par> <gc_out> [ref_image]

Input parameters

gc_in	geocoding lookup table
width	width of geocoding lookup table
DIFF_par	DIFF/GEO parameter file containing refinement polynomial coefficient
ref_image	reference image, offsets are measured relative to the reference image 0: actual SAR image 1: simulated SAR image (default)

Output parameters

gc_out	refined geocoding lookup table
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gcp_phase

<unw> <OFF_par> <gcp> <gcp_ph> [win_sz]

Input parameters

unw	unwrapped interferometric phase
OFF_par	ISP interferogram/offset parameter file

gcp	ground control point data (text format)
win_sz	window size for averaging phase for each GCP, must be odd (default = 1)
<u>Output parameters</u>	
gcp_ph	ground control point data + extracted unwrapped phase (text format)
<u>geocode</u>	
<gc_map> <data_in> <width_in> <data_out> <width_out> [nlines_out] [interp_mode] [format_flag] [lr_in] [lr_out] [n_ovr] [rad_max] [nintr]	
<u>Input parameters</u>	
gc_map	lookup table containing pairs of real-valued output data coordinates
data_in	data file (format as specified by format_flag parameter)
width_in	width of input data file and gc_map lookup table
width_out	width of output data file
nlines_out	number of lines for the output data file (enter - for default: all lines)
interp_mode	resampling interpolation mode (default = 0) (0: 1/dist, 1: nearest neighbor, 2: SQR(1/dist), 3: const, 4: Gaussian)
format_flag	input/output data format flag (default = 0) (0: float, 1: fcomplex, 2: SUN raster/BMP format, 3: unsigned char 4: short 5: scomplex)
lr_in	input SUN raster/BMP format left/write flipped (default = 1: not flipped, -1: flipped)
lr_out	output SUN raster/BMP format left/write flipped (default = 1: not flipped, -1: flipped)
n_ovr	interpolation oversampling factor (default = 2)
rad_max	maximum interpolation search radius (default = 4*n_ovr: 8)
nintr	number of points required for interpolation when not nearest neighbor (default = 4)
<u>Output parameters</u>	
data_out	output data file
<u>geocode_back</u>	
<data_in> <width_in> <gc_map> <data_out> <width_out> [nlines_out] [interp_mode] [format_flag] [lr_in] [lr_out]	
<u>Input parameters</u>	
data_in	data file (for SUN raster: *.ras, BMP format: *.bmp)
width_in	width of input data file
gc_map	lookup table containing pairs of real-valued input data coordinates
width_out	width of output data file and gc_map lookup table
nlines_out	number of lines of output data file (default = 0: all lines of gc_map)
interp_mode	interpolation mode (default = 0) (0: nearest-neighbor, 1: spline, 2: spline-log, 3: bilinear, 4: bilinear_log)
format_flag	input/output data format flag (default = 0) (0: float, 1: fcomplex, 2: SUN raster/BMP format, 3: unsigned char, 4: short)
lr_in	input SUN raster/BMP format left/write flipped (default = 1: not flipped, -1: flipped)
<u>Output parameters</u>	
data_out	output data file (for SUN raster: *.ras, BMP format: *.bmp)

lr_out	output SUN raster/BMP format left/write flipped (default = 1: not flipped, -1: flipped)
<u>init_offset</u>	
<SLC-1> <SLC-2> <SLC1_par> <SLC2_par> <OFF_par> [rlks] [azlks] [rpos] [azpos] [offr] [offaz] [thres] [cflag]	
<u>Input parameters</u>	
SLC-1	single-look complex image 1 (reference)
SLC-2	single-look complex image 2
SLC1_par	SLC-1 ISP image parameter file
SLC2_par	SLC-2 ISP image parameter file
OFF_par	(input/output) ISP offset/interferogram parameter file
rlks	number of range looks (default = 1)
azlks	number of azimuth looks (default = 1)
rpos	center of patch in range (samples) (enter - for default: image center)
azpos	center of patch in azimuth (lines) (enter - for default: image center)
offr	initial range offset (samples) (default = 0)
offaz	initial azimuth offset (lines) (default = 0)
thres	correlation SNR threshold (default: 7.000)
cflag	copy offsets to the range and azimuth offset polynomials in OFF_par (default = 0: no, 1: yes)
<u>init_offset_orbit</u>	
<SLC1_par> <SLC2_par> <OFF_par> [rpos] [azpos]	
<u>Input parameters</u>	
SLC1_par	SLC-1 parameter file
SLC2_par	SLC-2 parameter file
OFF_par	(input/output) ISP/offset parameter file
rpos	range position for offset estimation (default=center of SLC-1)
azpos	azimuth position for offset estimation (default=center of SLC-1)
<u>init_offsetm</u>	
<PWR-1> <PWR-2> <DIFF_par> [rlks] [azlks] [rpos] [azpos] [offr] [offaz] [thres] [patch] [cflag]	
<u>Input parameters</u>	
PWR-1	intensity image 1 (float) (reference)
PWR-2	intensity image 2 (float)
DIFF_par	DIFF/GEO parameter file
rlks	number of range looks (enter - for default: 1)
azlks	number of azimuth looks (enter - for default: 1)
rpos	center of region for comparison in range (enter - for default: image center)
azpos	center of region for comparison in azimuth (enter - for default: image center)
offr	initial range offset (enter - for default from DIFF_par)
offaz	initial azimuth offset (enter - for default from DIFF_par)
thres	correlation SNR threshold (default: 7.0)

patch	correlation patch size (128, 256, 512, 1024, enter - for default: 1024)
cflag	copy offsets to the range and azimuth offset polynomials in DIFF_par (default = 0: no, 1 : yes)
<u>intf_pt</u>	
<plist> <pmask> <itab> <rec_num> <pSLC> <pint> <type> [pSLC_par]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
rec_num	record number to process (default - : all records)
pSLC	point data stack of interpolated SLC values (fcomplex or scomplex)
type	pSLC data type (0: fcomplex, 1: scomplex)
pSLC_par	stack of SLC/MLI parameters used for phase correction of ENVISAT/ERS interferograms (binary)
<u>Output parameters</u>	
pint	point data stack of interferograms (fcomplex)
NOTE: if SLC have different carrier frequencies, pSLC_par data are required	
<u>lin_comb_pt</u>	
<plist> <pmask> <pdata1> <rec_num1> <pdata2> <rec_num2> <pdata3> <rec_num3> <constant> <factor1> <factor2> <type> [zero_flag] [pt_num]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar)
pdata1	point data stack 1 (various formats supported)
rec_num1	record number in pdata1 (starting with 1, enter - : all records)
pdata2	point data stack 2 (same format as pdata1)
rec_num2	record number in pdata2 (starting with 1, enter - : same as rec_num1)
rec_num3	record number in pdata3 (starting with 1, enter - : same as rec_num1)
constant	constant value to add to output record elements
factor1	factor to multiply elements of rec_num1 of pdata1
factor2	factor to multiply elements of rec_num2 of pdata2
type	data type (0: fcomplex, 1: scomplex, 2: float, 3: int, 4: short, 5: byte)
zero_flag	zero flag (default = 0: 0.0 is interpreted as NULL, 1: 0.0 is interpreted as valid value)
pt_num	use index of pdata2 point value rather than all points
<u>Output parameters</u>	
pdata3	point data stack (same format as pdata1)
<u>mcf</u>	
<interf> <wgt> <mask> <unw> <width> [tri_mode] [roff] [loff] [nr] [nlines] [npat_r] [npat_az] [ovrlap] [r_init]	

[az_init] [init_flag]

Input parameters

interf interferogram (*.int, *.flt) (fcomplex)
 wgt weight factors (0 -> 1.0) file (float) (enter - for uniform weight)
 mask validity mask (SUN raster or BMP format, value 0 -> pixel not used) (enter - if no mask)
 width number of samples/row
 tri_mode triangulation mode
 0: filled triangular mesh (default)
 1: Delaunay triangulation
 roff offset to starting range of section to unwrap (default: 0)
 loff offset to starting line of section to unwrap (default: 0)
 nr number of range samples of section to unwrap (default (-): width - roff)
 nlines number of lines of section to unwrap (default (-): total number of lines - loff)
 npat_r number of patches in range
 npat_az number of patches in azimuth
 overlap overlap between patches in pixels (overlap >= 7, default (-): 512)
 r_init phase reference point range offset (default (-): roff)
 az_init phase reference point azimuth offset (default (-): loff)
 init_flag flag to set phase at reference point
 0: use initial point phase value (default)
 1: set phase to 0.0 at initial point

Output parameters

unw unwrapped phase image (*.unw) (float)

merge_pt

<plist_tab> <plist_out> <N_min> <r_tol> <az_tol>

Input parameters

plist_tab text file containing names of input point lists and masks
 (each line contains one point list and optionally one mask name)
 N_min occurrence number minimum to include a point to the output list
 r_tol range position tolerance for counting occurrences
 az_tol azimuth line tolerance for counting occurrences

Output parameters

plist_out point list derived from merging input point lists

mk_2d_im

<plist> <pmask> <itab> <SLC_par> <pdata> <src> <nrec> <par_out> <type> <imode> <radius> <np_min>
 <out_dir> <ref_im> <cycle> <sflag> [start] [nlines]

Input parameters

plist point list (int)
 pmask point data stack of mask values

	(uchar, pmask[i] = 0 excludes point i, enter - if not available)
itab	table associating interferogram stack records with pairs of SLC stack records
SLC_par	ISP SLC parameter file of point list coordinates
pdata	point data stack (various types supported)
srec	starting record number in input point data stack (starting with 1)
nrec	number of records to process (enter - for all to end of list)
par_out	SLC/MLI parameter file of output 2-D image data file
type	data type (0: fcomplex, 2: float)
imode	interpolation mode (0: none, 1: 3-pt_bilinear, 2: 6-pt_bilinear, 3: convol, 4: nn_convool, default - : 6-pt_bilinear)
radius	window radius used (only used for imodes (3, 4), nominal = 1.2)
np_min	minimum number of points required for search region (only used for imodes (3, 4), default: 3)
out_dir	directory for output images
ref_im	reference multi-look RMLI 2D image used for display of point data (float)
cycle	data value per color cycle (applies only for float data, enter - or 0 for fcomplex data)
sflag	save data 2-D interpolated data files (0: no, 1: yes)
start	starting line to display (default: 1)
nlines	number of lines to display (default 0: to end of file)
log	not specified, automatically created
mk_adf_2d	
<RSLC_tab> <itab> <rml_i> <diff_dir> [cc_win] [adf_exp] [adf_win]	
<u>Input parameters</u>	
RSLC_tab	two column list of resampled SLC filenames and SLC parameter filenames (including paths) (ascii)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
rml_i	RMLI reference image with same rlls and azlls value as the interferograms (includes path)
diff_dir	differential interferogram directory
cc_win	correlation estimation window size (linear weighting) in pixels (default: 5)
adf_exp	exponent parameter for adf filtering of the interferogram (nominal range .2 --> .6) default = 0.4
adf_win	window size for adf filter (default = 32)
mk_base_2d_cs <RSLC_tab> <itab> <DEM_rdc> <diff_dir> <pbase> [mask] [nr] [naz] [gcp_win] [type]	
<u>Input parameters</u>	
RSLC_tab	two column list of resampled SLC filenames and SLC parameter filenames (including paths) (ascii)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
DEM_rdc	terrain height in radar range-Doppler coordinates (meters, float format)

diff_dir	directory containing unwrapped differential interferograms and baselines
mask	mask for selection of valid GCP (Sun raster or BMP format, enter - for none)
nr	number of GCP selection points in range
naz	number of GCP selection points in azimuth
gcp_win	window size for averaging unwrapped phase, must be odd (default = 3)
type	differential interferogram type: 0: unfiltered data (*.unw) 1: (default) adf filtered data (*.adf.unw)

Output parameters

pbase	baseline parameter stack (enter - for none)
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mk_diff_2d

<RSLC_tab> <itab> <bflag> <DEM_rdc> <def> <delta_t> <rml_i> <rml_i_dir> <diff_dir> <rlks> <azlks> <cc_win>
 [sflag] [mflag]

Input parameters

RSLC_tab	two column list of resampled SLC filenames and SLC parameter filenames (including paths) (text)
itab	table associating interferogram stack records with pairs of SLC stack records (text)
bflag	baseline flag 0: use initial baseline estimate derived from orbit data 1: use precision baseline and do not derive baseline from orbit data
DEM_rdc	terrain height in radar range-Doppler coordinates (m, float, enter - for none)
def	deformation rate (m/year) (enter - for none)
rml_i	reference image with same rlks and azlks value as the interferogram
rml_i_dir	directory containing RMLI images of the resampled SLC
diff_dir	directory for differential interferograms after subtraction of height model phase
rlks	range looks to use for interferogram generation
azlks	azimuth looks to use for interferogram generation
cc_win	correlation estimation window size (linear weighting) in pixels (default: 5)
sflag	apply spectral shift filtering (default = 0: off, 1: on)
mflag	flag to select baseline estimation method (default = 0): 0: orbits, orbits; p1, p2 1: offsets, offsets; p1, p2, off 2: orbits, fft; p1, p2, off, int 3: offsets, fft; p1, p2, off, int 4: fft, fft; p1, off, int

NOTE: mflag values of 1 and 3 only apply to itab files created with base_calc itab_mode = 0 (single reference)

mk_diff_all

<plist> <pmask> <pSLC_par> <itab> <pbase> <bflag> <pint> <int_type> <phgt> <pdef> <patm> <psim_unw>
 <pdiff> [log]

Input parameters

plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters for the resampled SLC data (binary)
itab	table associating interferogram stack records with pairs of SLC stack records (text)
pbase	stack of baseline parameters (binary)
bflag	baseline flag (0: use initial baseline, 1: precision baseline)
pint	point data stack of interferograms (fcomplex or float)
int_type	interferogram stack data type (0: float (unwrapped), default = 1: fcomplex)
phgt	point data stack of terrain height (m, float, single record, enter - for none)
pdef	point data stack of LOS deformation rate (m/yr, float, single record, enter - for none)
patm	point data stack of atmosphere phase estimates (enter - for none)

Output parameters

psim_unw	point data stack of simulated unwrapped interferometric phase (float)
pdiff	point data stack of differential interferograms after subtraction of model phase and atmosphere (enter - for none)
<u>log</u>	mk_diff_all log file

mk_geo_cs

<MLI> <MLI_par> <DEM> <DEM_par> <DEM_seg> <DEM_seg_par> <GEO_dir> <scene_id> <post> <mode>
 [ls_mode] [r_ovr] [n_ovr] [rlks] [azlks] [thres] [rpos] [azpos] [roff] [azoff] [r_patch] [az_patch]

Input parameters

MLI	MLI SAR image (including path)
MLI_par	ISP image parameter file of the MLI image (including path)
DEM	DEM in desired output projection (including path)
DEM_par	DEM parameter file (including path)

Output parameters

DEM_seg	DEM segment for output image products (including path)
DEM_seg_par	(input/output) DEM parameter file for output image products (including path), regenerated each time
GEO_dir	directory for output images, lookup tables and DEM products
scene_id	scene name to identify output files
post	output image sample spacing in meters or degrees for geographic (EQA) projection
mode	processing mode (see above)
ls_mode	algorithm selection in gc_map for regions of layover, shadow, or DEM gaps: <ul style="list-style-type: none"> 0: set to (0., 0.) 1: linear interpolation across these regions 2: use actual value (default) 3: nearest neighbor thinned (nn-thinned)
r_ovr	range over-sampling parameter for ls_mode = 3 (nn-thinned)

	in gc_map (r_ovr default: 2)
n_ovr	interpolation oversampling factor in geocode (default = 2.0)
rad_max	maximum interpolation search radius (default: 4*n_ovr)
NOTE: n_ovr and rad_max are parameters used by the program geocode for transformation of the simulated image and DEM into SAR geometry. The parameters rlks, azlks, thres, rpos, azpos, roff, azoff are used for estimation of the initial offset of the SAR image with respected to the simulated SAR image.	
rlks	number of range looks for the initial offset estimate (default: 1)
azlks	number of azimuth looks for the initial offset estimate (default: 1)
thres	SNR threshold for offset estimates (default: 10)
rpos	range position for initial offset (enter - for default)
azpos	azimuth position for initial offset (enter - for default)
roff	initial range offset estimate (enter - for current value in DIFF_par file)
azoff	initial azimuth offset estimate (enter - for current value in DIFF_par file)
r_patch	range patch size for offset estimation (default: 256 samples)
az_patch	azimuth patch size for offset estimation (default: 256 lines)

mk_int_all

<plist> <pmask> <pSLC_par> <itab> <pSLC> <pbase> <pint> [log]

Input parameters

plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters for the resampled SLC data (binary)
itab	table associating interferogram stack records with pairs of SLC stack records (text)
pSLC	point data stack of interpolated SLC values (enter - for none, fcomplex or scomplex)

Output parameters

pbase	stack of baseline parameters (enter - for none, binary)
pint	point data stack of interferograms (enter - for none, fcomplex)
log	mk_int_all log file

mk_mli_all

<SLC_tab> <MLI_dir> <rlks> <azlks> [sflag] [scale] [exp]

Input parameters

SLC_tab	two column list of slc filenames and slc parameter filenames (including paths) (ascii)
MLI_dir	directory to contain MLI images and MLI image parameter files
rlks	range looks for the MLI images
azlks	azimuth looks for the MLI images
sflag	summation flag, generate sum of MLI images: 0: no 1: yes (default)

scale	relative intensity scale factor (default = 0.9)
exp	display exponent (default = 0.35)
<u>mk_msr_pt</u>	
<SLC_tab> <SLC_par> <MLI_par> <MLI_ras> <MSR_dir> [PWR_min] [mode] [MSR_cal] [PWR_cal] [min_MSR] [num_MSR] [delta_MSR]	
<u>Input parameters</u>	
SLC_tab	two column list of co-registered SLC files and associated SLC parameter files (including paths, text)
SLC_par	reference resampled SLC image parameter file (text)
MLI_par	image parameter file of reference resampled MLI image (text)
MLI_ras	background reference resampled MLI image (SUN raster or BMP format)
MSR_dir	directory to contain mean/sigma ratio images, point lists, and point images
PWR_min	minimum relative intensity threshold for point target selection (default = 0.5)
mode	intensity normalization between SLC scenes (0: none, 1: image average, default = 2: point targets)
MSR_cal	mode 2 mean/sigma ratio for point target selection for relative calibration between scenes (default = 1.5)
PWR_cal	mode 2 intensity threshold ratio for point target selection for relative calibration between scenes (default = 1)
min_MSR	initial mean/sigma ratio threshold for point selection (default = 1.2)
num_MSR	number of MSR thresholds to test (default = 8)
delta_MSR	increment of MSR threshold (default = 0.1)
<u>mk_sp_all</u>	
<SLC_tab> <sp_dir> [rlks] [azlks] [PWR_min] [CC_min] [MSR_min]	
<u>Input parameters</u>	
SLC_tab	two column list of slc filenames and slc parameter filenames (including paths) (ascii)
sp_dir	directory to contain sp_stat generated msr, cc, and point lists
rlks	number of range spectral looks (default - : 4)
azlks	number of azimuth spectral looks (default - : 4)
PWR_min	intensity minimum threshold to accept a point (0.0 to ignore, default - : 0.5)
CC_min	spectral correlation minimum threshold to accept a point (0.4 to ignore, default -)
MSR_min	mean/sigma ratio minimum threshold (relative to spatial average) to accept a point (0.0 to ignore, default - : 1.2)
<u>mk_unw_2d</u>	
<RSLC_tab> <itab> <rml_i> <diff_dir> [cc_thres] [nlks] [npat_r] [npat_az] [mode] [r_init] [az_init]	
<u>Input parameters</u>	
RSLC_tab	two column list of resampled SLC filenames and SLC parameter filenames (including paths) (ascii)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)

rml	RMLI reference image with same rlks and azlks value as the interferograms (includes path)
diff_dir	differential interferogram directory
cc_thres	threshold for cc_mask (0.0 --> 1.0) (default = 0.4)
nlks	number of looks in range and azimuth to scale before unwrapping (default = 1)
npat_r	number of patches in range (default = 1)
npat_az	number of patches in azimuth (default = 1)
mode	processing mode: 0: unwrap unfiltered data (*.diff) 1: (default) unwrap adf filtered data (*.adf.diff) using adf correlation measure
r_init	phase reference range offset (default: 0)
az_init	phase reference azimuth offset (default: 0)
MLI_copy	
<MLI_in> <MLI_in_par> <MLI_out> <MLI_out_par> [roff] [nr] [loff] [nl]	
<u>Input parameters</u>	
MLI_in	multi-look intensity image (float)
MLI_in_par	ISP image parameter file for input MLI
roff	offset to starting range sample (default = 0)
nr	number of range samples (default = 0, to end of line)
loff	offset to starting line (default = 0)
nl	number of lines to copy (default = 0, to end of file)
<u>Output parameters</u>	
MLI_out	selected MLI section (float format)
MLI_out_par	
ISP image parameter file for output MLI	
msk_pt	
<plist1> <pmask1> <mask> <plist2> <pmask2> <rlks> <azlks>	
<u>Input parameters</u>	
plist1	point list 1 (int) (enter - for none)
pmask1	point data stack of mask values (uchar pmask[i] = 0 excludes point i) (enter - for none)
mask	raster image file (8-bit/pixel Sun raster or BMP format) points located in regions set to 0 are masked out
rlks	mask file range looks relative to the point coordinates
azlks	mask file azimuth looks relative to the point coordinates
<u>Output parameters</u>	
plist2	output list of points within the mask (int) (enter - for none)
pmask2	output mask, matches length of plist1 if plist2 is not specified
multi_cpx	
<cpx_input> <OFF_par_in> <cpx_output> <OFF_par_out> [rlks] [azlks] [loff] [nlines] [roff] [nsamp]	
<u>Input parameters</u>	

cpx_input	input complex image file
OFF_par_in	interferogram/offset parameter file for input image
rlks	number of range looks values < -1 interpreted as an image magnification factor (default = 1)
azlks	number azimuth looks values < -1 interpreted as an image magnification factor (default = 1)
loff	line offset to starting line (default = 0)
nlines	number of lines (default = 0, to end of file)
roff	offset to starting range sample (default = 0)
nsamp	number of range samples to extract (default = 0, to end of line)
<u>Output parameters</u>	
cpx_output	output multi-look or interpolated complex data file
OFF_par_out	interferogram/offset parameter file for output, use existing parameter file if available

multi_def_pt

<plist> <pmask_in> <pSLC_par> <ppos> <itab> <pbase> <bflag> <pdiff> <pdiff_type> <np_ref> <pres> <pdh>
 <pdef> <punw> <psigma> <pmask_out> [dh_max] [def_max] [rpatch] [sigma_max] [sigma_max2] [model]
 [noise_min] [bmax] [dtmax]

Input parameters

plist	point list (int)
pmask_in	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters (binary)
ppos	point data stack of interpolated point target positions (fcomplex, enter - to use plist coordinates)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
pbase	stack of baseline parameters (binary)
bflag	baseline flag (0: initial baseline, 1: precision baseline)
pdiff	point data stack of differential interferograms (fcomplex or float)
pdiff_type	type of pdiff (0: float (unwrapped), default = 1: fcomplex)
np_ref	phase reference point number (beginning from 0)
dh_max	maximum height correction for initial fit (m, enter - for default: 60.0)
def_max	maximum deformation rate difference for initial fit (m/year, enter - for default: 2.0e-02)
rpatch	patch size in range pixels (enter - for default: 100)
sigma_max	threshold for phase std. deviation to set mask to valid: (enter - for default: 1.2)
sigma_max2	threshold for phase std. deviation for patch to patch fit: (enter - for default: 0.75)
model	phase model (see above)
noise_min	noise minimization for patch to patch processing (default=0:off 1:on)
bmax	maximum perpendicular baseline (m) considered, (-1 = all records, default = -1)
dtmax	maximum time interval (days) considered, (-1 = all records, default = -1)

Output parameters

pres	point data stack of residual unmodeled phase (float, enter - for none)
pdh	point data stack of height correction value (m, float, enter - for none)
pdef	point data stack of linear deformation rate (m/year, float, enter - for none)
punw	point data stack of unwrapped phase of pdiff (float, enter - for none)
psigma	point data stack of phase standard deviation from fit (float, enter - for none)
pmask_out	point data stack of mask values indicating accepted points (uchar, enter - for none)
multi_look	
<SLC> <SLC_par> <MLI> <MLI_par> <rlks> <azlks> [loff] [nlines] [scale] [image_format]	
<u>Input parameters</u>	
SLC	single-look complex image
SLC_par	SLC ISP image parameter file
rlks	number of range looks
azlks	number of azimuth looks
loff	offset to starting line (default = 0)
nlines	number of SLC lines to process (default = entire file)
scale	scale factor for output MLI (default = 1.0)
<u>Output parameters</u>	
MLI	multi-look intensity image
MLI_par	MLI ISP image parameter file
multi_real	
<real_input> <OFF_par_in> <real_output> <OFF_par_out> [rlks] [azlks] [loff] [nlines] [roff] [nsamp]	
<u>Input parameters</u>	
real_input	input real image file
OFF_par_in	interferogram/offset parameter file for input image
rlks	number of range looks values < -1 interpreted as an image magnification factor (default = 1)
azlks	number azimuth looks values < -1 interpreted as an image magnification factor (default = 1)
loff	line offset to starting line (default = 0)
nlines	number of lines (default = 0, to end of file)
roff	offset to starting range sample (default = 0)
nsamp	number of range samples to extract (default = 0, to end of line)
<u>Output parameters</u>	
real_output	output multi-look or interpolated real data file
OFF_par_out	interferogram/offset parameter file for output, use existing parameter file if available
multi_SLC	
<SLC_PROC_par> <MLI_PROC_par> <SLC> <MLI> <rlks> <azlks> [slc_format]	
<u>Input parameters</u>	
SLC_PROC_par	SLC MSP processing parameter file

SLC	single-look complex image
rlks	number of range looks
azlks	number of azimuth looks
slc_format	input SLC format flag (default: from SLC_PROC_par) 0: FCOMPLEX (pairs of 4-byte float) 1: SCOMPLEX (pairs of 2-byte short integer)
<u>Output parameters</u>	
MLI_PROC_par	MLI MSP processing parameter file
MLI	multi-look intensity image derived from SLC
npt	
<plist> [pmask]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar pmask [i] = 0 excludes point i)
offset_fit	
<offs> <snr> <OFF_par> [coffs] [coffsets] [thres] [npoly] [interact_flag]	
<u>Input parameters</u>	
offs	range and azimuth offset estimates (fcomplex)
snr	SNR values of offset estimates (float)
OFF_par	ISP offset/interferogram parameter file
thres	SNR threshold (enter - for default from OFF_par)
npoly	number of polynomial model parameters (enter - for default, 1, 3, 4, default = 4)
interact_flag	interactive culling of input data (1 = YES, 0 = NO, default = NO)
<u>Output parameters</u>	
coffs	culled range and azimuth offset estimates (fcomplex, enter - for none)
coffsets	culled offset estimates and SNR values (text format, enter - for none)
offset_fitm	
<offs> <snr> <DIFF_par> [coffs] [coffsets] [thres] [npoly] [interact_flag]	
<u>Input parameters</u>	
offs	range and azimuth offset estimates (fcomplex)
snr	SNR values of offset estimates (float)
DIFF_par	DIFF/GEO parameter file for the scene
thres	SNR threshold (enter - for default from DIFF_par)
npoly	number of polynomial model parameters (enter - for default, 1, 3, 4, default = 4)
interact_flag	interactive culling of input data (1 = YES, 0 = NO, default = NO)
<u>Output parameters</u>	
coffs	culled range and azimuth offset estimates (fcomplex, enter - for none)
coffsets	culled offset estimates and SNR values (text format, enter - for none)
offset_pwr	

<SLC-1> <SLC-2> <SLC1_par> <SLC2_par> <OFF_par> <offs> <snr> [rwin] [azwin] [offsets] [n_ovr] [nr] [naz]
[thres] [pflag]

Input parameters

SLC-1 single-look complex image 1 (reference)
SLC-2 single-look complex image 2
SLC1_par SLC-1 ISP image parameter file
SLC2_par SLC-2 ISP image parameter file
OFF_par ISP offset/interferogram parameter file
rwin search window size (range pixels, (enter - for default from offset parameter file))
azwin search window size (azimuth pixels, (enter - for default from offset parameter file))
n_ovr SLC oversampling factor (integer 2**N (1, 2, 4) default = 2)
nr number of offset estimates in range direction
 (enter - for default from offset parameter file)
naz number of offset estimates in azimuth direction
 (enter - for default from offset parameter file)
thres offset estimation quality threshold (enter - for default from offset parameter file)
pflag print flag (0: print offset summary (default), 1: print all offset data)

Output parameters

offs offset estimates (fcomplex)
snr offset estimation SNR (float)
offsets range and azimuth offsets and SNR data in text format, enter - for no output

offset_pwr

<PWR-1> <PWR-2> <DIFF_par> <offs> <snr> [rwin] [azwin] [offsets] [n_ovr] [nr] [naz] [thres] [pflag]

Input parameters

PWR-1 real valued intensity image 1 (reference)
PWR-2 real valued intensity image 2
DIFF_par DIFF/GEO parameter file
rwin search window size (range pixels, (enter - for default from offset parameter file))
azwin search window size (azimuth pixels, (enter - for default from offset parameter file))
n_ovr image oversampling factor (integer 2**N (1, 2, 4) default = 2)
nr number of offset estimates in range direction
 (enter - for default from offset parameter file)
naz number of offset estimates in azimuth direction
 (enter - for default from offset parameter file)
thres offset estimation quality threshold (enter - for default from offset parameter file)
pflag print flag (0: print offset summary, 1: print all offset data (default))

Output parameters

offs offset estimates (fcomplex)
snr offset estimation SNR (float)

offsets	range and azimuth offsets and SNR data in text format, enter - for no output
ORB_prop	
<PROC_par> [nstate] [interval] [extra]	
<u>Input parameters</u>	
PROC_par	(input/output) MSP processing parameter file
nstate	number of state vectors to calculate for the MSP image parameter file, enter - for the default value determined from the duration of the state vectors)
interval	time interval between state vectors (default: input state vector time interval)
extra	extra time for state vector coverage at start and end of image (default= 30.0 s.)
par_MSP	
<SAR_par> <PROC_par> <SLC/MLI_par> [image_format]	
<u>Input parameters</u>	
SAR_par	MSP SAR sensor parameter file
PROC_par	MSP processing parameter file
image_format	image format flag (default: from MSP processing parameter file) 0: fcomplex (pairs of 4-byte float) 1: scomplex (pairs of 2-byte short integer) 2: float (4-bytes/value)
<u>Output parameters</u>	
SLC/MLI_par	ISP SLC/MLI image parameter file
pdisdt_pwr24, prasdt_pwr24	
<plist> <pmask> <SLC_par> <pdata> <rec_num> <par_out> <mli> <cycle> [radius]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, pmask[i] = 0 excludes point i, enter - if not available)
SLC_par	SLC parameter file of point list coordinates
pdata	point data stack (float)
rec_num	record number to process (default - : all records)
par_out	SLC/MLI parameter file of 2-D output image data file
mli	2-D image data file used as output image intensity
cycle	data value per color cycle
radius	interpolation window radius used default: 4.0)
phase_sim	
<SLC1_par> <OFF_par> <baseline> <hgt> <sim_unw> [ph_flag] [bflag] [def] [delta_t] [SLC2R_par]	
<u>Input parameters</u>	
SLC1_par	parameter file of reference SLC-1
OFF_par	ISP offset/interferogram parameter file

baseline	baseline parameter file
hgt	height map in the same geometry as the interferogram (m, float, enter - for none)
ph_flag	range phase trend selection: 0: unflattened interferogram (default) 1: flattened interferogram
bflag	baseline selection: 0: initial baseline (default) 1: precision baseline
def	LOS deformation rate map (m/yr, float, enter - for none)
delta_t	interferogram time interval (days, required for deformation modeling, enter - for none)
SLC2R_par	parameter file of resampled SLC, required if SLC-2 frequency differs from SLC-1
<u>Output parameters</u>	
sim_unw	simulated unwrapped interferometric phase
<u>phase_sim_pt</u>	
<plist> <pmask> <pSLC_par> <ppos> <itab> <rec_num> <pbase> <phgt> <psim_unw> [pdef] [ph_flag] [bflag]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters (binary)
ppos	point data stack of interpolated point target positions (fcomplex, enter - to use plist coordinates)
itab	table associating interferogram stack records with pairs of SLC stack records (ascii)
rec_num	itab record number to process (default - : all records)
pbase	stack of baseline parameters (binary)
phgt	point data stack of terrain height (m, float, single record, enter - for none)
pdef	point data stack of LOS deformation rate (m/yr, float, single record, enter - for none)
ph_flag	phase model flag 0: unflattened (default) 1: flattened 2: height + deformation phase, relative to ref. layer
bflag	baseline flag 0: use initial baseline (default) 1: precision baseline
<u>Output parameters</u>	
psim_unw	point data stack of simulated unwrapped interferometric phase (float)
<u>pre_rc</u>	
<SAR_par> <PROC_par> <signal_data> <rc_data> [prefilt_dec] [loff] [nl] [nr_samp] [kaiser] [filt_lm] [nr_ext] [fr_ext]	
<u>Input parameters</u>	

SAR_par	MSP SAR sensor parameter file
PROC_par	MSP processing parameter file
signal_data	uncompressed raw SAR signal data filename
prefilt_dec	prefilter decimation factor (enter - for default from PROC_par)
loff	number of lines offset from start of file (enter - for default from PROC_par)
nl	number of lines to range compress (enter - for default from PROC_par)
nr_samp	number of range samples (enter - for default from PROC_par)
kaiser	range chirp Kaiser window parameter beta (default = 2.120, -30 dB sidelobes)
filt_lm	filter length multiplier, FIR length = FIR_lm * prefilt_dec + 1 (default = 8)
nr_ext	near range swath extension in samples (default from PROC_par)
fr_ext	far range swath extension in samples (default from PROC_par)
<u>Output parameters</u>	
rc_data	range compressed data filename
pt2data	
<plist> <pmask> <SLC_par> <pdata> <rec_num> <f_out> <par_out> [type] [imode] [radius] [np_min]	
<u>Input parameters:</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, pmask[i] = 0 excludes point i, enter - if not available)
SLC_par	SLC parameter file of point list coordinates
pdata	point data stack (various types supported)
rec_num	record number in input point data stack (starting with 1)
par_out	SLC/MLI parameter file of 2-D output image data file
type	data type (0: fcomplex, 1: scomplex, 2: float, 3: int, 4: short, 5: byte, default - : float)
imode	interpolation mode (0: none, 1: 3-pt_bilinear, 2: 6-pt_bilinear, 3: convol, 4: nn_convol, default - : 6-pt_bilinear)
radius	window radius used (only used for imodes (3, 4), default: 4.0)
np_min	minimum number of points required for search region (only used for imodes (3, 4), default: 3)
<u>Output parameters</u>	
f_out	2-D image data file (same type as input point data stack)
pwr_stat	
<SLC_tab> <SLC_par> <MSR> <plist> [MSR_min] [PWR_min] [roff] [loff] [nr] [nl] [type] [mode] [MSR_cal] [PWR_cal]	
<u>Input parameters</u>	
SLC_tab	text file containing list of names of co-registered SLC in column 1
SLC_par	SLC parameter file of point list coordinates (i.e. for the reference SLC)
MSR_min	mean/sigma ratio minimum threshold to accept a point (0.0 to ignore, default - : 1.5)
PWR_min	intensity minimum threshold (relative to spatial average) to accept a point

	(0.0 to ignore, default - : 1.0)
roff	offset to starting range of section to analyze (default - : 0)
loff	offset to starting line of section to analyze (default - : 0)
nr	number of range pixels to analyze (default - : to end of line)
nl	number of azimuth lines to analyze (default - : to end of file)
type	SLC type (default 0: fcomplex, 1: scomplex)
mode	intensity normalization between SLC scenes (0: none, 1: image average, default = 2: point targets)
MSR_cal	mode 2 mean/sigma ratio for point target selection for relative calibration between scenes: 1.5
PWR_cal	mode 2 intensity threshold ratio for point target selection for relative calibration between scenes: 1.0
<u>Output parameters</u>	
MSR	mean/sigma ratio image (float, enter - for none)
plist	point list (int, enter - for none)
<u>ras_linear</u>	
<pwr> <width> [start] [nlines] [pixavr] [pixavaz] [min] [max] [LR] [rasf] [inverse] [channel]	
<u>Input parameters</u>	
pwr	intensity image (float)
width	samples per row of pwr
start	starting line of pwr (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
min	minimum value for display (default= 0.0)
max	maximum value for display (default = 1.0)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
inverse	inverse flag (default = 1: float_to_raster, -1: raster_to_float)
channel	RGB channel flag (default = 1: red, 2: green, 3: blue)
<u>rasf</u>	(output/input) raster image (default = *.ras: SUN raster format, *.bmp: BMP format)
<u>ras_pt</u>	
<plist> <pmask> <ras_in> <ras_out> [rlks] [azlks] [r] [g] [b] [xs] [zflg] [mflg]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
ras_in	raster image (SUN raster or BMP format)
rlks	range looks of raster image relative to point coordinates (default = 1.0)
azlks	azimuth looks of raster image relative to point coordinates (default = 1.0)
r	line color value red (0 --> 255) default: 255

g	line color value green (0 --> 255) default: 255
b	line color value blue (0 --> 255) default: 0
xs	size of cross in pixels, set to 1 for single points (default: 1)
zflg	zero image flag (default = 0: retain image values, 1: set all image values to 0 except crosses)
mflg	mask flag (0: display all points default = 0: display points only where pmask[i] = 1, 1: display points where pmask[i] = 0)
<u>Output parameters</u>	
ras_out	raster image with crosses drawn at points (SUN raster or BMP format)
<u>rascc</u>	
<cc> <pwr> <width> [start_cc] [start_pwr] [nlines] [pixavr] [pixavaz] [cmin] [cmax] [scale] [exp] [LR] [rasf]	
<u>Input parameters</u>	
cc	correlation coefficient image (float)
pwr	intensity data (float, enter - if not available)
width	samples per row of cc and pwr
start_cc	starting line of cc (default = 1)
start_pwr	starting line of pwr (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
cmin	minimum correlation value used for linear cc display (default = .1)
cmax	maximum correlation value used for linear cc display (default = .9)
scale	pwr display scale factor (default = 1.)
exp	pwr display exponent (default = .35)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
<u>Output parameters</u>	
rasf	raster image (default = *.ras: SUN raster format, *.bmp: BMP format)
<u>rascc_mask</u>	
<cc> <pwr> <width> [start_cc] [start_pwr] [nlines] [pixavr] [pixavaz] [cc_thres] [cc_min] [cc_max] [scale] [exp] [LR] [rasf]	
<u>Input parameters</u>	
cc	coherence image (float)
pwr	intensity image (float, enter - if not available)
width	number of samples/row
start_cc	starting line of coherence image (default = 1)
start_pwr	starting line of intensity image (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)

cc_thres	coherence threshold for masking, all pixels with coherence below cc_thres are set to (0, 0, 0)
cc_min	minimum coherence value used for color display (default = .1)
cc_max	maximum coherence value used for color display (default = .9)
scale	intensity display scale factor (default = 1.)
exp	intensity display exponent (default = .35)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
<u>Output parameters</u>	
rasf	validity mask (default = *.ras: SUN raster format, *.bmp: BMP format)
<u>rashgt</u>	
<hgt> <pwr> <width> [start_hgt] [start_pwr] [nlines] [pixavr] [pixavaz] [m/cycle] [scale] [exp] [LR] [rasf]	
<u>Input parameters</u>	
hgt	height data (float)
pwr	intensity data (float, enter - if not available)
width	samples per row of hgt and pwr
start_hgt	starting line of hgt (default = 1)
start_pwr	starting line of pwr (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
m/cycle	meters per color cycle (default = 160.)
scale	display scale factor (default = 1.)
exp	display exponent (default = .35)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
<u>Output parameters</u>	
rasf	raster image (default = *.ras: SUN raster format, *.bmp: BMP format)
<u>rasmph_pwr</u>	
<cpx> <pwr> <width> [start_cpx] [start_pwr] [nlines] [pixavr] [pixavaz] [scale] [exp] [LR] [rasf] [cc] [start_cc] [cc_min]	
<u>Input parameters</u>	
cpx	complex image (fcomplex, e.g. interferogram)
pwr	intensity image (float)
width	number of samples/row of cpx and pwr
start_cpx	starting line of cpx (default = 1)
start_pwr	starting line of pwr (default = 1)
nlines	number of lines to display (default=0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
scale	pwr display scale factor (default = 1.)

exp	pwr display exponent (default = .35)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
cc	display threshold data file (float, e.g. correlation)
start_cc	starting line of cc data file (default = 1)
cc_min	pixels with cc values below cc_min are displayed using grayscale (default = .2)
<u>Output parameters</u>	
rasf	raster file (enter - for default=*.ras: SUN raster format, *.bmp: BMP format)
raspwr	
<pwr> <width> [start] [nlines] [pixavr] [pixavaz] [scale] [exp] [LR] [rasf] [data_type] [hdrsz]	
<u>Input parameters</u>	
pwr	intensity image (float or short integer data type)
width	samples per row
start	starting line to display (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
scale	display scale factor (default = 1.)
exp	display exponent (default = .35)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image) for descending track
data_type	input data type (default = 0) 0: float 1: short integer
hdrsz	line header size in bytes (default=0, ESA PAF for PRI products=12)
<u>Output parameters</u>	
rasf	raster image (enter - for default = *.ras: SUN raster format, *.bmp: BMP format example)
rasmg	
<unw> <pwr> <width> [start_unw] [start_pwr] [nlines] [pixavr] [pixavaz] [ph_scale] [scale] [exp] [ph_offset] [LR] [rasf][rasf] [cc] [start_cc] [cc_min]	
<u>Input parameters</u>	
unw	unwrapped phase data (float)
pwr	intensity data (float, enter - if not available)
width	samples per row of unwrapped phase and intensity files
start_unw	starting line of unwrapped phase file (default = 1)
start_pwr	starting line of intensity file (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
ph_scale	phase display scale factor (default = .33333)
scale	pwr display scale factor (default = 1.)

exp	pwr display exponent (default = .35)
ph_offset	phase offset in radians subtracted from unw (default = .0)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
cc	correlation threshold data file (float): e.g. correlation
start_cc	starting line of threshold data file (default = 1)
cc_min	pixels with cc values below cc_min are displayed using grayscale (default = .2)
<u>Output parameters</u>	
rasf	raster image (enter - for default = *.ras: SUN raster format, *.bmp: BMP format)
rasshd	
<DEM> <width> <col_post> [row_post] [start] [nlines] [pixavr] [pixavaz] [theta0] [phi0] [LR] [rasf] [format] [zero_flag]	
<u>Input parameters</u>	
DEM	digital elevation model (float)
width	samples per row of DEM
col_post	posting between cols (same unit as DEM values)
row_post	posting between rows (same unit as DEM, default = col_post)
start	starting line of DEM (default = 1)
nlines	number of lines to display (default = 0: to end of file)
pixavr	number of pixels to average in range (default = 1)
pixavaz	number of pixels to average in azimuth (default = 1)
theta0	illumination elevation angle in deg. (default = 45.)
phi0	illumination orientation angle in deg. (default = 135.) (0: right, 90: top, 180: left, 270: bottom)
LR	left/right flipping flag, (default = 1: normal, -1: mirror image)
format	data format (default = 0: float, 1: short integer)
zero_flag	zero_flag default = 0: 0.0 interpreted as missing value, 1: interpreted as valid value
<u>Output parameters</u>	
rasf	raster image (default = *.ras: SUN raster format *.bmp: BMP format)
<u>set_value</u>	
<PAR_in> <PAR_out> <keyword> <value> [new_key]	
<u>Input parameters</u>	
PAR_in	keyword:value based parameter file
keyword	search keyword of keyword:value pair
value	new value (note: delimit value with double quotes if it contains spaces or punctuation)
new_key	0 = default: new keyword:value pair not permitted, 1: insert new keyword:value pair
<u>Output parameters</u>	
PAR_out	keyword:value based parameter file (can be the same file as PAR_in)
<u>SLC_copy</u>	
<SLC_in> <SLC_par_in> <SLC_out> <SLC_par_out> [fcase] [sc] [roff] [nr] [loff] [nl] [swap] [header_lines]	

<u>Input parameters</u>	
SLC_in	SLC (FCOMPLEX or SCOMPLEX format)
SLC_par_in	ISP SLC parameter file for input SLC
fcase	data format conversion (enter - for default: output format = input format) 1: fcomplex --> fcomplex (default sc = 1.0) 2: fcomplex --> scomplex (default sc = 10000.0) 3: scomplex --> fcomplex (default sc = 0.0001) 4: scomplex --> scomplex (default sc = 1.0)
sc	scale factor for input SLC data (enter - for default)
roff	offset to starting range sample (enter - for default: 0)
nr	number of range samples (enter - for default: to end of line)
loff	offset to starting line (enter - for default: 0)
nl	number of lines to copy (enter - for default: to end of file)
swap	swap data (enter - for default) 0: normal (default) 1: swap real/imaginary part of complex data 2: swap left/right (near/far range)
header_lines	number of input file header lines (enter - for default: 0)
<u>Output parameters</u>	
SLC_out	selected SLC section (fcomplex or scomplex format)
SLC_par_out	ISP SLC parameter file of output SLC
<u>SLC_interp</u>	
<SLC-2> <SLC1_par> <SLC2_par> <OFF_par> <SLC-2R> <SLC2R_par> [loff] [nlines]	
<u>Input parameters</u>	
SLC-2	SLC-2 image to be resampled to the geometry of the SLC-1 reference image
SLC1_par	SLC-1 ISP image parameter file
SLC2_par	SLC-2 ISP image parameter file
OFF_par	ISP offset/interferogram parameter file
loff	offset to first valid output line (in SLC-1 lines) (default = 0)
nlines	number of valid output lines (default = 0: to end of file)
<u>Output parameters</u>	
SLC-2R	single-look complex image 2 co-registered to SLC-1
SLC2R_par	SLC-2R ISP image parameter file for co-registered image
<u>SLC_intf</u>	
<SLC-1> <SLC-2R> <SLC1_par> <SLC2R_par> <OFF_par> <interf> <rlks> <azlks> [loff] [nlines] [sps_flg] [azf_flg]	
<u>Input parameters</u>	
SLC-1	single-look complex image 1 (reference)
SLC-2R	single-look complex image 2 co-registered to SLC-1

SLC1_par	SLC-1 ISP image parameter file
SLC2R_par	SLC-2R ISP image parameter file for the co-registered image
OFF_par	ISP offset/interferogram parameter file
rlks	number of range looks
azlks	number of azimuth looks
loff	offset to starting line relative to SLC-1 for interferogram (default = 0)
nlines	number of SLC lines to process (enter - for default: to end of file)
sps_flg	range spectral shift flag: 1: apply spectral shift filter (default) 0: do not apply spectral shift filter
azf_flg	azimuth common band filter flag: 1: apply azimuth common band filter (default) 0: do not apply azimuth common band filter
<u>Output parameter</u>	
interf	interferogram from SLC-1 and SLC-2R
<u>SLC_par_pt</u>	
<SLC_par> <pSLC_par> <SLC_rec_num> <rw_flag>	
<u>Input parameters</u>	
SLC_par	SLC/MLI parameter file (enter - for output to std out)
SLC_rec_num	record number in the stack of SLC/MLI image parameters (starts with 1)
rw_flag	read/write flag (0: read from pSLC_par, 1: write to pSLC_par)
<u>Output parameters</u>	
pSLC_par	stack of SLC/MLI parameter files (binary)
<u>SLC_resamp_all_cs</u>	
<SLC_tab> <ref_SLC> <ref_par> <rslc_dir> <RSLC_tab> <mode> [rflag] [rlks] [azlks] [rpos] [azpos]	
<u>Input parameters</u>	
SLC_tab	two column list of SLC filenames and SLC parameter filenames (including paths) (ascii)
ref_SLC	reference SLC (including path)
ref_par	ISP image parameter file of the reference SLC (including path)
rslc_dir	directory to receive the resampled SLC and ISP image parameter files
mode	processing mode
rflag	measure offsets to the resampled SLC to confirm the offset model (default = 0:off, 1: on)
rlks	number of range looks for initial offset estimation (enter - for default)
azlks	number of azimuth looks for initial offset estimation (enter - for default)
rpos	center of patch in range (samples) (enter - for default: image center)
azpos	center of patch in azimuth (lines) (enter - for default: image center)
<u>Output parameters</u>	
<u>RSLC_tab</u>	RSLC_tab file for the resampled SLC files
<u>SLC2pt</u>	

<SLC_tab> <plist> <pmask> <pSLC_par> <pSLC> <SLC_rec_num>

Input parameters

SLC_tab two column list of SLC filenames and SLC parameter filenames (including paths) (ascii)
 plist point list (enter - for none) (int)
 pmask point data stack of mask values (uchar, set to - to accept all points)

Output parameters

pSLC_par stack of SLC/MLI parameters for the resampled SLC data (enter - for none) (binary)
 pSLC point data stack of interpolated SLC values (enter - for none) (fcomplex or scomplex)
 SLC_rec_num record number in the output point data stack (starting with 1, default: - for all)
 log not specified, automatically created

sp_stat

<SLC> <pwr> <cc> <MSR> <plist> <width> [PWR_min] [CC_min] [MSR_min] [rlks] [azlks] [roff] [loff] [nr] [nl] [bx]
 [by] [type]

Input parameters

SLC SLC file (fcomplex or scomplex)
 width number of samples/row of the SLC data
 PWR_min intensity minimum threshold to accept a point (0.0 to ignore, default - : 10.0)
 CC_min spectral correlation minimum threshold to accept a point (0.0 to ignore, default - : 0.4)
 MSR_min mean/sigma ratio minimum threshold (relative to spatial average) to accept a point
 (0.0 to ignore, default - : 1.2)
 rlks number of range looks to use (default - : 4)
 azlks number of azimuth looks to use (default - : 4)
 roff offset to starting range of section to analyze (default - : 0)
 loff offset to starting line of section to analyze (default - : 0)
 nr number of range pixels to analyze (default - : to end of line)
 nl number of azimuth lines to analyze (default - : to end of file)
 bx window width in SLC pixels used for spectral statistics (default - : rlks)
 by window number of SLC lines used for spectral statistics (default - : azlks)
 type SLC type (default = 0: fcomplex, 1: scomplex)

Output parameters

pwr intensity image (float, enter - for none)
 cc spectral correlation image (float, enter - for none)
 MSR mean/sigma ratio image (float, enter - for none)
 plist point list of point target candidates (int)

spf_pt

<plist> <pmask> <SLC_par> <pdata_in> <pdata_out> [rec_num] [type] [r_max] [spf_type] [np_max] [pt_num]
 [mode]

Input parameters

plist point list (int)

pmask	point data stack of mask values (uchar, enter - if not available)
SLC_par	SLC parameter file of point list coordinates
pdata_in	point data stack (various types supported)
rec_num	record number in pdata_in and pdata_out to process (default - : all records)
type	data type (0: fcomplex, 1: scomplex, 2: float, default - : float)
r_max	maximum radius (range samples, default: 64)
spf_type	spatial filter type 0: uniform average (default for fcomplex and scomplex) 1: triangular weighted average 2: linear least-squares (default for float data)
np_max	maximum number of neighbor points in the filter window (default - : all points within search radius)
pt_num	point index for local filtering about a single point (first point = 0)
mode	processing mode (default = 0: spatial filtering, 1: reference point phase bias removal)
<u>Output parameters</u>	
pdata_out	spatially filtered point data stack (same type as pdata_in)
<u>stacking</u>	
<diff_tab> <width> <ph_rate> <sig_ph_rate> <sig_ph> <roff> <loff> [nr] [nl] [np_min]	
<u>Input parameters</u>	
diff_tab	2 column list of unwrapped diff. interferograms and delta_T values in days (text)
width	number of samples/line of the interferograms in the stack
roff	range pixel offset to center of the phase reference region
loff	line offset to center of the phase reference region
nr	number of range pixels to average in the phase reference region (enter - for default = 16)
nl	number of lines average in the phase reference region (enter - for default= 16)
np_min	min. number of phase values required to accept phase rate estimate (enter - for default = nfiles)
<u>Output parameters</u>	
ph_rate	average phase rate determined from a weighted sum of phases (radians/year, float)
sig_ph_rate	standard deviation of the estimated phase rate (radians/year, float)
sig_ph	standard deviation of the residual phases (enter - for none, radians, float)
<u>sub_phase</u>	
<int-1> <unw-2> <DIFF_par> <diff_int> <int-1_type> [sub/add_flag]	
<u>Input parameters</u>	
int-1	complex or unwrapped interferogram 1 file
unw-2	unwrapped interferogram 2 file
DIFF_par	differential interferogram parameter file
int-1_type	int-1 type: 0 = unwrapped phase, 1 = complex interferogram

sub/add_flag	0 = subtract phase of unw-2; 1 = add phase of unw-2 (default = 0)
<u>Output parameters</u>	
diff_int	differential interferogram output file
<u>sub_phase_pt</u>	
<plist> <pmask> <pint1> <rec_num> <punw2> <pdiff> <int_type> [inverse]	
<u>Input parameters</u>	
plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pint1	point data stack of interferograms (fcomplex or float)
rec_num	record number to process (default - : all records)
punw2	point data stack of unwrapped phases (float)
int_type	data type of pint1 and pdiff 0: float (unwrapped phase) 1: fcomplex
inverse	subtract/add flag (default = 0: subtract punw2, 1: add punw2)
<u>Output parameters</u>	
pdiff	point data stack of differential interferograms (punw2 subtracted/added, same type as pint1)
<u>swap_bytes</u>	
<infile> <outfile> <swap_type>	
<u>Input parameters</u>	
infile	input data file
swap_type	swap type flag (bytes/value) 2: (1, 2, 3, 4, 5, 6, 7, 8 ...) --> (2, 1, 4, 3, 6, 5, 8, 7 ...) (short, scomplex) 4: (1, 2, 3, 4, 5, 6, 7, 8 ...) --> (4, 3, 2, 1, 8, 7, 6, 5 ...) (int, float, fcomplex) 8: (1, 2, 3, 4, 5, 6, 7, 8 ...) --> (8, 7, 6, 5, 4, 3, 2, 1 ...) (double)
<u>Output parameters</u>	
outfile	output data file
<u>thres_im_pt</u>	
<f_in> <width> <plist> <t_min> <t_max> <rlks> <azlks>	
<u>Input parameters</u>	
f_in	2-D image data file (float)
width	number of samples/row of f_in
t_min	minimum threshold value (enter - for none)
t_max	maximum threshold value (enter - for none)
rlks	number of range looks in the sampled data geometry
azlks	number of azimuth looks in the sampled data geometry
<u>Output parameters</u>	
plist	point list of coordinates of all non-zero points (int)

Note: It is recommended to try a few different thresholds in order to get a variety point densities. Save the according number of points.

unw_model

<interf> <unw_model> <unw> <width> [xinit] [yinit] [ref_ph]

Input parameters

interf	interferogram (*.int, *.flt) (fcomplex)
unw_model	approximate unwrapped phase image (= model, float)
width	number of samples/row
xinit	offset to phase reference location range pixel (col)
yinit	offset to phase reference location azimuth pixel (row)
ref_ph	reference phase to use other than value at (xinit, yinit)

Output parameters

unw	unwrapped phase (float)
-----	-------------------------

vu_disp

<plist> <pmask> <pSLC_par> <itab> <pdisp> <pdef> <phgt> <psigma> <pdef_err> <pdh_err> <pmap> <ras>
[ymin] [ymax] [mag] [win_sz]

Input parameters

plist	point list (int)
pmask	point data stack of mask values (uchar, set to - to accept all points)
pSLC_par	stack of SLC/MLI parameters (binary)
itab	table associating interferogram stack records with pairs of SLC (enter - for none) (ascii)
pdisp	displacement (m) for each point and each layer in the stack (float)
pdef	point data stack of linear deformation rate (m/year, float)
phgt	point data stack of height (m, float)
psigma	point data stack of phase standard deviation from fit (float)
pdh_err	estimated uncertainty in the height (m, float, enter - for none)
pdef_err	estimated uncertainty in linear deformation rate (m/year, float, enter - for none)
pmap	point positions in map projection coordinates (easting, northing or lat, lon) (enter - for none) (float)
ras	raster reference image with PT locations marked (SUN *.ras, or BMP *.bmp)
ymin	plot display minimum (enter - for default = -0.1)
ymax	plot display maximum (enter - for default = +0.1)
mag	zoom magnification factor (default = 2)
win_sz	zoom window size before magnification (default = 192)

Appendix G – List of Figures and Linked Files

Note: Does not include www links

2. Introduction

Figure 1: [Deformation_fig.pdf](#)

Figure 2: [Kapoho_fig.pdf](#)

3. GAMMA Processing work flow for the Kapoho subsidence project

[README](#)

[html GAMMA manual](#)

MSP – Modular SAR Processor

[Modular SAR Processor](#)

MSP2 ASAR raw data preprocessing

[ASAR_pre_proc](#)

MSP2.1 ASAR_pre_list

[ASAR_pre_list](#)

MSP2.2 ASAR_pre_proc

[ASAR_proc_all](#)

Mode 1 [log](#)

[antenna](#)

sensor parameter: [SAR_par example](#)

processing parameter: [PROC_par example](#)

[raw](#)

Mode 2 [log](#)

Mode 3 [log](#)

[GAMMA manual](#)

[spectrum: text](#)

Mode 4 [log](#)

Mode 5 [log](#)

[ASAR_all_list](#)

MSP3.1 Detect looks

Figure 3a: [Envisat_range_fig.pdf](#)

Figure 3b: [Envisat_range_fig.pdf](#)

Figure 4a: [SLC_MLI.pdf](#)

Figure 3c: [Envisat_range_fig.pdf](#)

MSP3.2 ASAR_proc_all

[ASAR_proc_all_no_autof](#)

[rc_data](#)

Figure 4a: [SLC_MLI.pdf](#)

[MLI](#)

[MLI PROC par](#)

Figure 4b: [SLC_MLI.pdf](#)

[ISP image parameter](#)

[MLI ISP image parameter](#)

ISP – Interferometric SAR processor – Resampling

[Interferometric SAR processor](#)

[ISP_README](#)

[SLC tab](#)

ISP1 Calculate baselines and generate itab file

[base calc](#)

[itab: example](#)

[SLC tab](#)

ISP2 Reference scene: Selection-criteria and algorithms

[scenelocation.m](#)

Figure 5a, b: [Matlab_MLIs_fig.pdf](#)

ISP3 Resample set of SLC to a common reference SLC

[SLC_resamp_all](#)

ISP3.3 Generate raster image of RMLI

Figure 5c: [Matlab_MLIs_fig.pdf](#)

ISP3.4 Resample SLC to reference RSLC

[RSLC tab](#)

[SLC_resamp_all_cs](#)

Mode 0 [log](#)

[create_offset.in](#)

Mode 1 [log](#)

Mode 2 [log](#)

Mode 3 [log](#)

Mode 4 [log](#)

ISP4 Create multi-look images for individual RSLC and calculate average MLI

Figure 5d: [Matlab_MLIs_fig.pdf](#)

Figure 5e: [Matlab_MLIs_fig.pdf](#)

GEO – SAR geocoding and image registration

[SAR geocoding and image registration](#)

GEO1 Type and location of raw DEM data

[meta](#)

[header](#)

[*.blw](#)

GEO3 Byte order verification

Figure 6a: [DEM_EQA_UTM.pdf](#)

GEO4 Conversion from EQA to UTM

[UTM DEM parameter file](#)
[mk_geo_cs](#)

Figure 6b: [DEM_EQA_UTM.pdf](#)

Figure 5e: [Matlab_MLIs_fig.pdf](#)

GEO5 Transforming DEM map coordinates to RDC geometry

[mk_geo](#)
Mode 0 [log](#)

Figure 7a: [SIM_SAR_0_1.pdf](#)

Figure 8: [dis2ras_0.pdf](#)

Mode 2 [log](#)
Mode 3 [log](#)
Mode 4 [log](#)

Figure 9: [dis2ras_1.pdf](#)

Figure 10: [DEM_RDC_MLI.pdf](#)

Figure 30b: [PDEM_RDC.pdf](#)

Figure 11: [disras_dem_par_RMLI_MAP.pdf](#)

ISP/DIFF – Interferograms and differential interferometry

[ISP](#)
[DIFF](#)

ISP/DIFF1 Generate itab with all possible pair combinations

[itab_all](#)

ISP/DIFF2 Generate differential interferograms

[mk_diff_2d](#)
[log](#)
[off_par.in](#)

Figure 12: [INT.pdf](#)

[diff_par.in](#)

Figure 13a, b: [TOPO_DIFF.pdf](#)

Figure 13c, d: [TOPO_DIFF.pdf](#)

Figure 14: [CC.pdf](#)

ISP/DIFF4 Apply adf filter to the stack of differential interferograms

[log](#)

Figure 15: [ADF.pdf](#)

ISP/DIFF5 Unwrap the differential interferograms

[log](#)

Figure 16a: [CC_MASK.pdf](#)

Figure 17: [INI_REF_BASELINES.pdf](#)

ISP/DIFF6 Baseline refinement

Figure 18: [Baseline.pdf](#)

[mk_base_2d](#)
[mk_diff_2d_cs](#)

ISP/DIFF6.1 Create mask

Figure 16b: [CC_MASK.pdf](#)

ISP/DIFF6.2 Estimate improved baselines from unwrapped phase and DEM in RDC

[log](#)
[GCP](#)
[gcp_ph](#)
[baseline example](#)

ISP/DIFF8 Stack the data

[diff_tab](#)

ISP/DIFF9 Create displacement map in RDC

Figure 19: [Displacement_close.pdf](#)

Figure 20a: [DISPL_RDC_UTM.pdf](#)

ISP/DIFF10 Geocode displacements to transform to UTM

Figure 21: [DISPL_UTM.pdf](#)

Figure 20b: [DISPL_RDC_UTM.pdf](#)

Figure 22: [DISPL_rasrmg.pdf](#)

IPTA – Interferometric Point Target Analysis

[Interferometric Point Target Analysis](#)
[GAMMA IPTA Processing Example Luxemburg](#)

IPTA1 Focus on smaller area of interest (optional)

Figure 23: [new_area.pdf](#)

IPTA1.1 Crop the RSLC to desired region

[SLC_par_in](#)
[SLC_par_out](#)

IPTA1.2 Create new RSLC_tab

[RSLC_tab GAMMA description](#)
[existing RSLC_tab \(example\)](#)
[RSLC_tab kapoho](#)

IPTA1.3 Generate new RMLI, average RMLI and parameter file
[parameter file](#)

Figure 24: [CROP_MLI_AVE.pdf](#)

IPTA1.4 Crop DEM (MLI format) to desired region
[MLI_out_par](#)

Figure 25: [disras_DEM_RDC.pdf](#)

IPTA2 Create IPTA interferogram table

[IPTA interferogram table GAMMA description](#)

[base_calc](#)

[itab_single](#): [full standard output](#)

IPTA3 Generate point list pt

[point list GAMMA description](#)

IPTA3.1 Generation of point target candidate list based on spectral properties of individual SLC

[MLI_tab](#)

IPTA3.2 Use spectral correlation average and different thresholds to create point list

Figure 27: [PT_CC.pdf](#)

IPTA3.3 Generation of point target candidate list based on low intensity variability

[mk_msr_pt](#)

Figure 28: [PT_MSR.pdf](#)

IPTA3.4 Merging of point target candidate lists to create the point list

[list](#)

Figure 29a: [MERGED_MASKED.pdf](#)

IPTA3.5 Create a mask raster file

Figure 29a: [MERGED_MASKED.pdf](#)

IPTA4 Generate pSLC_par and SLC point data stacks pSLC

[log](#)

IPTA5.1 Generate DEM point data file pdem

Figure 30a: [PDEM_RDC.pdf](#)

IPTA5.2.1 Create interferogram and baseline data stacks

[log](#)

[mk_int_all](#)

IPTA5.2.2 Display the layers of pint

[log](#)

Figure 31: [PINT0.pdf](#)

IPTA5.3.1 Simulate topographic phase and subtract it from interferogram

[log](#)
[mk_diff_all](#)

IPTA5.3.2 Create raster images of the differential interferograms

Figure 32: [PDIFF0.pdf](#)

IPTA6.1 Interactive point-wise phase regression analysis

Figure 33: [dis_ipta.pdf](#)

IPTA6.2.1 Multi-patch estimation of linear deformation, height corrections and residual phase

Figure 34: [MULTI_DEF_PT1.pdf](#)

[prasdt_pwr24](#)
[pdisdt_pwr24](#)

Figure 35: [PRES1.pdf](#)

Figure 36: [PRES1_126.pdf](#)

Figure 37: [PUNW1.pdf](#)

IPTA6.2.2 Estimation of linear deformation rate, height correction and residual phase over the entire scene

[itab_selection](#)

Figure 38: [DEF_MOD_PT.pdf](#)

Figure 39: [PRES2.pdf](#)

Figure 40: [PRES2_126.pdf](#)

Figure 35: [PRES1.pdf](#)

Figure 41: [PUNW2.pdf](#)

IPTA8 Calculate updated differential interferograms

[mk_diff_all](#) [log](#)
[mk_2d_im](#) [log](#)

Figure 42: [PDIFF0.pdf](#)

IPTA9 Regression analysis on updated differential interferograms

Figure 43: [MULTI_DEF_PT2.pdf](#)

Figure 44: [PRES1.pdf](#)

Figure 45: [PRES1_126.pdf](#)

Figure 46: [PUNW1.pdf](#)

IPTA11.1 Estimate and apply atmospheric phase

Figure 47: [PRES1_SPF_25.pdf](#)

Figure 48: [PUNW1_NOATM.pdf](#)

Figure 49: [PUNW1_NOATM_SPF.pdf](#)

IPTA11.3 Visualization of deformation histories

Figure 50: [vu_disp.pdf](#)

References

[Caccamise](#)

[Delaney](#)

[Ferretti](#)

[Fiedler](#)

[Hannsen](#)

[Hooper](#)

[Li](#)

[Madson](#)

[Rosen](#)

[Wegmueller Luxemburg](#)

[Wegmueller GAMMA](#)

[Wegmueller Unwrapping](#)

[Werner](#)

[Wiesmann](#)

Appendix C – Complimentary Information

Figure 51: [radar.pdf](#)

Appendix D – Data availability and acquisition tools

Figure 52: [EoliSA.pdf](#)

Figure: 53: [Descw.pdf](#)

Figure 54: [winsar.pdf](#)

Figure 55: [PGF.pdf](#)

Appendix H – Processing Examples

The processing examples give command summaries for each chapter. Using the information throughout the according chapter as well as the example data on the provided CD (also available as zipped download) and assuming a UNIX system with GAMMA (version 20051219) installed, the user is bale to re-generate all files and images. Please also refer to the [README](#) for additional information.

1) MSP

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_1.log
ASAR_all_list 1
```

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_2.log
ASAR_all_list 2
```

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_3.log
ASAR_all_list 3
```

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_4.log
ASAR_all_list 4 doppler_polynomial "-150.3790 0. 0. 0."
```

```
ASAR_pre_proc ASAR_pre_list /mnt/scsi/DATA/DORIS/vor raw ASAR_pre_proc_5.log
ASAR_all_list 5
```

```
mkdir tmp
mkdir slc
mkdir mli_1_5
```

```
ASAR_proc_all_no_autof ASAR_all_list raw tmp slc mli_1_5 1 5 0
```

```
rasSLC slc/20030212.slc 5158 1 0 1 1 1. .5 1 0 0 slc/20030212.slc.ras
convert -scale 20% -quality 50 slc/20030212.slc.ras slc/20030212.slc.jpg
```

Generated Files	Compare to Figures
slc/20030212.slc	4a) SLC, 4b) MLI

2) ISP - Resampling

```

mkdir rslc

SLC_copy slc/20041208.slc slc/20041208.slc.par rslc/20041208.rslc
rslc/20041208.rslc.par 1

mkdir rml_i_1_5

multi_look rslc/20041208.rslc rslc/20041208.rslc.par rml_i_1_5/20041208.rml_i
rml_i_1_5/20041208.rml_i.par 1 5

raspwr rml_i_1_5/20041208.rml_i 5174

xv rml_i_1_5/20041208.rml_i.ras

SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 0
1 1 5

SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 1
1 1 5

SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 2
1 1 5

SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 3
1 1 5

offset_pwr rslc/20041208.rslc slc/20060308.slc rslc/20041208.rslc.par
slc/20060308.slc.par rslc/20041208_20060308.off rslc/20041208_20060308.off
rslc/20041208_20060308.snr - - - 2 500 500

offset_fit rslc/20041208_20030212.off rslc/20041208_20030212.snr
rslc/20041208_20030212.off rslc/20041208_20030212.coffs - 6.5 4 1

offset_fit rslc/20041208_20030702.off rslc/20041208_20030702.snr
rslc/20041208_20030702.off rslc/20041208_20030702.coffs - 6.5 4 1
etc.

SLC_resamp_all_cs SLC_tab rslc/20041208.rslc rslc/20041208.rslc.par rslc RSLC_tab 4
1 1 5

mk_mli_all RSLC_tab rml_i_1_5 1 5 1
cp rml_i_1_5/20041208.rml_i.par rml_i_1_5/rml_i_1_5.ave.par

dis2ras rml_i_1_5/20030212.rml_i.ras rml_i_1_5/20041208.mli.ras

```

Generated Files	Compare to Figures
rml_i_1_5/20041208.rml_i	5c) RMLI of reference scene
rml_i_1_5/20030212.rml_i	5d) RMLI of a co-registered scene
rml_i_1_5/rml_i_1_5.ave	5e) RMLI of average intensity image

3) GEO

```

create_dem_par SRTM_BI_larc_eqa.dem_par

swap_bytes SRTM_BI_larc_eqa.dem SRTM_BI_larc_eqa_swp.dem 2

disdem_par SRTM_BI_larc_eqa_swp.dem SRTM_BI_larc_eqa.dem_par

create_dem_par SRTM_BI_15m_utm.dem_par rml_i_1_5/rml_i_1_5.ave.par

dem_trans SRTM_BI_larc_eqa.dem_par SRTM_BI_larc_eqa_swp.dem SRTM_BI_15m_utm.dem_par
SRTM_BI_15m_utm.dem

disdem_par SRTM_BI_15m_utm.dem SRTM_BI_15m_utm.dem_par
gimp rml_i_1_5/rml_i_1_5.ave.ras

mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem
SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI
15 0 2 2 2 8 1 5 6.5 - - - - 256 256

mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem
SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI
15 2 2 2 2 8 1 5 6.5 - - - - 256 256

mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem
SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI
15 2 2 2 2 8 1 5 6.5 3000 3000 - - 256 256

mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem
SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI
15 3 2 2 2 8 1 5 6.5 - - - - 256 256

offset_fitm geo/BI.off_s geo/BI.sn_r geo/BI.diff geo/BI.coffs - 6.5 4 1

mk_geo_cs rml_i_1_5/rml_i_1_5.ave rml_i_1_5/rml_i_1_5.ave.par SRTM_BI_15m_utm.dem
SRTM_BI_15m_utm.dem_par SRTM_BI_15m_utm_seg.dem SRTM_BI_15m_utm_seg.dem_par geo BI
15 4 2 2 2 8 1 5 6.5 - - - - 256 256

```

Generated Files	Compare to Figures
SRTM_BI_larc_eqa_swp.dem	6a) DEM in EQA
SRTM_BI_15m_utm.dem	6b) DEM in UTM
geo/BI_0.sim_rdc	7a) Initial simulated SAR, 8a) dis2ras
geo/BI_1.sim_rdc	7b) Refine simulated SAR, 9a) dis2ras
geo/BI_dem.rdc	10a), 30b) DEM in RDC
geo/BI_map.mli	11) RMLI in UTM

4) ISP/DIFF

```
base_calc SLC_tab slc/20041208.slc.par 20041208.bperp.gr 20041208.bperp itab_all 1
mk_diff_2d RSLC_tab itab_all 0 geo/BI_dem.rdc - rml_1_5/rml_1_5.ave rml_1_5
diff0_2d 1 5 3 1 0

xv diff0_2d/*.diff.ras&

base_calc SLC_tab slc/20041208.slc.par 20041208.bperp.gr 20041208.bperp itab 1 0
750 -
mk_adf_2d RSLC_tab itab rml_1_5/rml_1_5.ave diff0_2d 5 .3 32

mk_unw_2d RSLC_tab itab rml_1_5/rml_1_5.ave diff0_2d .4 2 1 1 1

rascc_mask diff0_2d/20030212_20030702.adf.cc rml_1_5/rml_1_5.ave 5174 1 1 0 1 1
.8 .3 1.0 1. .35 1 diff0_2d/20030212_20030702.adf.cc_mask.8.ras

gimp diff0_2d/20030212_20030702.adf.cc_mask.8.ras

mk_base_2d_cs RSLC_tab itab geo/BI_dem.rdc diff0_2d -
diff0_2d/20030212_20030702.adf.cc_mask.8.ras 128 128 5 1

cat diff0_2d/20030212_20030702.base

mkdir diff1_2d
cp diff0_2d/*.base diff1_2d/
mv diff0_2d/*.unw_int diff1_2d/
cp diff0_2d/*.off diff1_2d/
mv /mnt/scsi/PGFSAR/IPTA/2429_3213/diff0_2d/*adf.cc diff1_2d/

mk_diff_2d_cs RSLC_tab itab 1 geo/BI_dem.rdc - rml_1_5/rml_1_5.ave rml_1_5
diff1_2d 1 5 3 1 0

dis2ras diff0_2d/20030212_20030702.adf.unw1.ras
diff1_2d/20030212_20030702.adf.diff_unw1.ras &

stacking diff_tab 5174 ph_rate sig_ph_rate sig_ph 2041 2413 - - -

dispmaph ph_rate - slc/20041208.slc.par diff1_2d/20051228_20060308.off ph_rate_disp

dishgt ph_rate_disp rml_1_5/rml_1_5.ave 5174 1 1 0 0.01 1 .35

rashgt ph_rate_disp rml_1_5/rml_1_5.ave 5174 1 1 0 1 1 .01 1. .35 -1
ph_rate_disp.ras

geocode_back ph_rate_disp 5174 geo/BI_1.map_to_rdc ph_rate_disp.utm 8334 0 0 0

geocode_back rml_1_5/rml_1_5.ave 5174 geo/BI_1.map_to_rdc geo/BI_map.mli 8334
7825 1 0

geocode_back diff1_2d/20031119_20040303.adf.cc 5174 geo/BI_1.map_to_rdc disp.cc
8334 7825 1 0

dishgt ph_rate_disp.utm geo/BI_map.mli 8334 1 1 0 0.01 1 .35
```

```

rashgt ph_rate_disp.utm geo/BI_map.mli 8334 1 1 0 1 1 .01 1. .35 1
ph_rate_disp.utm.ras

rasrmg ph_rate_disp.utm geo/BI_map.mli 8334 1 1 0 1 1 670 1. .35 0 1
ph_rate_disp.utml.ras disp.cc 1 .2

gimp ph_rate_disp.utml.ras

convert -crop 5040x4460+60+240 ph_rate_disp.ras ph_rate_disp.jpg
convert -scale 70% -quality 60 ph_rate_disp.jpg ph_rate_disp.jpg

convert -crop 4900x6060+0+440 ph_rate_disp.utml.ras ph_rate_disp.utml.jpg
convert -scale 70% -quality 60 ph_rate_disp.utml.jpg ph_rate_disp.utm2.jpg

```

Generated Files	Compare to Figures
diff0_2d_1_5/20030212_20030702.int	12) Interferogram
diff0_2d_1_5/20030212_20030702.sim_unw	13a,b) Simulated topography
diff0_2d_1_5/20030212_20030702.diff	13c,d) Differential interferogram
diff0_2d_1_5/20030212_20030702.cc	14) Interferometric correlation
diff0_2d_1_5/20030212_20030702.adf.diff	15c,d) Filtered differential interferogram
diff0_2d_1_5/20030212_20030702.adf.cc_mask	16a) Coherence (CC) mask
diff0_2d_1_5/20030212_20030702.adf.unw	17a,b) Unwrapped differential interferogram
diff0_2d_1_5/20030212_20030702.adf.cc_mask8	16b) CC mask with very high threshold
diff1_2d_1_5/20030212_20030702.diff_unw	17c,d) Refined unwrapped differential interferogram
ph_rate_disp	19) Displacement in RDC (dishgt), 20a) Displacement in RDC (rashgt)
ph_rate_disp.utm	21) Displacement in UTM (dishgt), 20b) Displacement in UTM (rashgt), 22) Displacement in UTM (rasrmg)

5) IPTA

```
SLC_copy rslc/20030212.rslc rslc/20030212.rslc.par rslc_kapoho/20030212.rslc
rslc_kapoho/20030212.rslc.par 1 - 1912 500 11500 2600
```

```
SLC_copy rslc/20030702.rslc rslc/20030702.rslc.par rslc_kapoho/20030702.rslc
rslc_kapoho/20030702.rslc.par 1 - 1912 500 11500 2600
```

etc.

```
mk_mli_all RSLC_tab_kapoho rml_i_1_5_kapoho 1 5 1
```

```
cp rml_i_1_5_kapoho/20041208.rml_i.par rml_i_1_5_kapoho/rml_i_1_5.ave.par
```

```
MLI_copy geo/BI_dem.rdc rml_i_1_5/rml_i_1_5.ave.par geo/BI_dem_kapoho.rdc
geo/BI_dem_kapoho.rdc.par 1912 500 11500 2600
```

```
rasshd geo/BI_dem_kapoho.rdc 500 15 15 1 0 1 1
```

```
disras geo/BI_dem_kapoho.rdc.ras&
```

```
base_calc RSLC_tab_kapoho rslc_kapoho/20041208.rslc.par 20041208.bperp.gr
20041208.bperp itab_single 0
```

```
mk_sp_all RSLC_tab_kapoho sp 4 4 0.5 0.4 1.2
```

```
thres_im_pt sp/ave.sp_cc 500 sp/pt_cc_318 0.318 - 1 1
ras_pt sp/pt_cc_318 - rml_i_1_5_kapoho/rml_i_1_5.ave.ras sp/pt_cc_318.ras 1 5
```

```
thres_im_pt sp/ave.sp_cc 500 sp/pt_cc_35 0.35 - 1 1
ras_pt sp/pt_cc_35 - rml_i_1_5_kapoho/rml_i_1_5.ave.ras sp/pt_cc_35.ras 1 5
```

```
thres_im_pt sp/ave.sp_cc 500 sp/pt_cc_3 0.3 - 1 1
ras_pt sp/pt_cc_3 - rml_i_1_5_kapoho/rml_i_1_5.ave.ras sp/pt_cc_3.ras 1 5
```

```
mk_msr_pt RSLC_tab_kapoho rslc_kapoho/20041208.rslc.par
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave.ras msr .5 2 1.5 1
1.2 10 .1
```

```
ls -l sp/pt_cc_318 msr/pt_1.7 > plist_tab
```

```
merge_pt plist_tab pt_merged 1 0 0
```

```
ras_pt pt - rml_i_1_5_kapoho/rml_i_1_5.ave.ras pt_merged.ras 1 5 255 0 0 3
```

```
xv pt.ras
```

```
msh_pt pt_merged - rml_i_1_5_kapoho/mask.ras pt_masked pmask 1 5
```

```
cp pt_merged pt_nomask
mv pt_masked pt
```

```
ras_pt pt - rml_i_1_5_kapoho/rml_i_1_5.ave.ras pt.ras 1 5
```

```
xv pt.ras
```

```
SLC2pt RSLC_tab_kapoho pt - pSLC_par pSLC -

SLC_par_pt rslc_kapoho/20060308.rslc.par pSLC_par 21 1
data2pt rslc_kapoho/20060308.rslc rslc_kapoho/20060308.rslc.par pt
rslc_kapoho/20060308.rslc.par pSLC 21 0

data2pt geo_kapoho/BI_dem.rdc rml_i_1_5_kapoho/rml_i_1_5.ave.par pt
rslc_kapoho/20041208.rslc.par pdem 1 2

data2pt geo/BI_dem_kapoho.rdc geo/BI_dem_kapoho.rdc.par pt
rslc_kapoho/20041208.rslc.par pdem2 1 2

pdisdt_pwr24 pt - rslc_kapoho/20041208.rslc.par pdem 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 256. 1

pdisdt_pwr24 pt - rslc_kapoho/20041208.rslc.par pdem2 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 256. 1

mk_int_all pt - pSLC_par itab_single pSLC pbase pint mk_int_all.log

mk_2d_im pt - itab_single rslc_kapoho/20041208.rslc.par pint 1 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par 0 3 1.2 1 int0 rml_i_1_5_kapoho/rml_i_1_5.ave - 0

xv int0/*ras

mk_diff_all pt - pSLC_par itab_single pbase 0 pint 1 pdem2 - - psim_unw0 pdiff0
mk_diff_all0.log

mk_2d_im pt - itab_single rslc_kapoho/20041208.rslc.par pdiff0 1 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par 0 3 1.2 1 diff0 rml_i_1_5_kapoho/rml_i_1_5.ave - 0

xv diff0/*ras

cp pdiff0 pdiff0_no_phgt
cp -r diff0 diff0_no_phgt

dis_ipta pt - pSLC_par - itab_single pbase 0 pdiff0 1 diff0/pdiff0_008.ras 30 0.01
2 - - 3 128

Choose this pt as reference

ref.:23558 x:104 y:1502 point:13003 x:149 y: 714

mkdir multi_def_pt1

multi_def_pt pt - pSLC_par - itab_single pbase 0 pdiff0 1 23558 multi_def_pt1/pres1
multi_def_pt1/pdh1 multi_def_pt1/pddef1 multi_def_pt1/punw1 multi_def_pt1/psigma1
multi_def_pt1/pmask1 30. 0.01 70 1.2 1.0 2 0 -1 -1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pdh1 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 30.0 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pddef1 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 0.01 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/punw1 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 12.6 1
```

```
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/psigma1 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 1.5 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pres1 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 12.6 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par multi_def_pt1/pres1 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 2.5 1

take out 2,3,7,12

mkdir def_mod_pt

def_mod_pt pt multi_def_pt1/pmask1 pSLC_par - itab_selection pbase 0
multi_def_pt1/punw1 0 23558 def_mod_pt/pres2 def_mod_pt/pdh2 def_mod_pt/pddef2
def_mod_pt/punw2 def_mod_pt/psigma2 def_mod_pt/pmask2 25. 0.03 3.0 2 - - -

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pdh2 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 30.0 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pddef2 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 0.01 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/punw2 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 12.6 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/psigma2 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 1.5 1

prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pres2 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 12.6 1
prasdt_pwr24 pt multi_def_pt1/pmask1 rslc/20040303.rslc.par def_mod_pt/pres2 -
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 2.5 1

lin_comb_pt pt def_mod_pt/pmask2 pdem2 1 def_mod_pt/pdh2 1 phgt1 1 -0. 1. 1. 2 0

cp def_mod_pt/pddef2 pdef1

mk_diff_all pt def_mod_pt/pmask2 pSLC_par itab_single pbase 0 pint 1 phgt1 pdef1 -
def_mod_pt/psim_unw0 def_mod_pt/pdiff0 def_mod_pt/mk_diff_all0.log

mk_2d_im pt def_mod_pt/pmask2 itab_single rslc_kapoho/20041208.rslc.par
def_mod_pt/pdiff0 1 - rml_i_1_5_kapoho/rml_i_1_5.ave.par 0 3 1.2 1 def_mod_pt/diff0
rml_i_1_5_kapoho/rml_i_1_5.ave - 0

mkdir multi_def_pt2

multi_def_pt pt def_mod_pt/pmask2 pSLC_par - itab_single pbase 0 def_mod_pt/pdiff0
1 23558 multi_def_pt2/pres1 multi_def_pt2/pdh1 multi_def_pt2/pddef1
multi_def_pt2/punw1 multi_def_pt2/psigma1 multi_def_pt2/pmask1 3. 0.003 70 1.3 0.9
2 1 -1 -1

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pdh1 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 30.0 1

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pddef1 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 0.01 1
```

```
prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/punw1 -
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 12.6 1

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/psigma1 1
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 1.5 1

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pres1 -
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 12.6 1

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20040303.rslc.par multi_def_pt2/pres1 -
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 2.5 1

spf_pt pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/pres1
atm/pres1_spf_10_1 - 2 10 1 -

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20041208.rslc.par atm/pres1_spf_10_1 -
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 4 1

spf_pt pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/pres1
atm/pres1_spf_25_1 - 2 25 1 -

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20041208.rslc.par atm/pres1_spf_25_1 -
rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave 4 1

cp atm/pres1_spf_25_1 atm/patm1

lin_comb_pt pt multi_def_pt2/pmask1 atm/patm1 - atm/patm1 14 atm/patmlx - 0. 1. -1.
2 1

sub_phase_pt pt multi_def_pt2/pmask1 multi_def_pt2/punw1 - atm/patmlx
multi_def_pt2/punw1_noatm 0 0

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20041208.rslc.par
multi_def_pt2/punw1_noatm - rml1_1_5_kapoho/rml1_1_5.ave.par
rml1_1_5_kapoho/rml1_1_5.ave 12.6 1

spf_pt pt multi_def_pt2/pmask1 rslc/20041208.rslc.par multi_def_pt2/punw1_noatm
multi_def_pt2/punw1_noatm_spf - 2 25 0 - 23558 0

prasdt_pwr24 pt multi_def_pt2/pmask1 rslc/20041208.rslc.par
multi_def_pt2/punw1_noatm_spf - rml1_1_5_kapoho/rml1_1_5.ave.par
rml1_1_5_kapoho/rml1_1_5.ave 12.6 1

mkdir def_mod_pt2

def_mod_pt pt multi_def_pt2/pmask1 pSLC_par - itab_single pbase 0
multi_def_pt2/punw1_noatm_spf 0 23558 def_mod_pt2/pres2 def_mod_pt2/pdh2
def_mod_pt2/pddef2 def_mod_pt2/punw2 def_mod_pt2/psigma2 def_mod_pt2/pmask2 25.
0.03 2.0 2 def_mod_pt2/pdh_err2a def_mod_pt2/pdef_err2a def_mod_pt2/ppc_err2a

prasdt_pwr24 pt def_mod_pt2/pmask2 rslc_kapoho/20041208.rslc.par
def_mod_pt2/psigma2 1 rml1_1_5_kapoho/rml1_1_5.ave.par rml1_1_5_kapoho/rml1_1_5.ave
1.5 1

lin_comb_pt pt def_mod_pt2/pmask2 phgt1 1 def_mod_pt2/pdh2 1 phgt2 1 -0. 1. 1. 2 0

lin_comb_pt pt def_mod_pt2/pmask2 pdef1 1 def_mod_pt2/pddef2 1 pdef2 1 -0. 1. 1. 2
0
```

```

phase_sim_pt pt def_mod_pt2/pmask2 pSLC_par - itab_single - pbase - ptmp1 pdef2 1 0

lin_comb_pt pt def_mod_pt2/pmask2 ptmp1 - def_mod_pt2/pres2 - pdef_phase1 - 0.0 1.
1. 2 1

dispmap_pt pt def_mod_pt2/pmask2 pSLC_par itab_single pdef_phase1 phgt2 pdispl 0

prasdt_pwr24 pt def_mod_pt2/pmask2 rslc_kapoho/20041208.rslc.par pdispl 1
rml_i_1_5_kapoho/rml_i_1_5.ave.par rml_i_1_5_kapoho/rml_i_1_5.ave 0.05 0

vu_disp pt def_mod_pt2/pmask2 pSLC_par itab_single pdispl pdef2 phgt2
def_mod_pt2/psigma2 def_mod_pt2/pdh_err2a def_mod_pt2/pdef_err2a - pdispl.ras

```

Generated Files	Compare to Figures
rml_i_1_5_kapoho/20040812.rml_i	24a) Cropped RMLI of reference scene
rml_i_1_5_kapoho/rml_i_1_5.ave	24b) Cropped RMLI of average intensity image
geo/BI_dem_kapoho.rdc	25) Cropped DEM in RDC
sp/pt_cc *	27) Spectral correlation point locations
msr/pt_*	28) Intensity variability point locations
pt	29a) Merged points
pt_masked	29b) Masked points
pdem2	30a) Point heights
pint	31) Point interferograms
pdiff0	32) Point differential interferograms
multi_def_pt1 (pdh1,pddef1,psigma1)	34) Results from 1st multi-patch regression analysis
multi_def_pt1 (pres1)	35,36) Residual phases
multi_def_pt1 (punw1)	37) Unwrapped phases
def_mod_pt (pdh2,pddef2,psigma2)	38) Results from 1st single-patch regression analysis (correct unw)
def_mod_pt (pres2)	39,40) Residual phases
def_mod_pt (punw2)	41) Unwrapped phases
def_mod_pt/pdiff0	42) Updated differential interferograms
multi_def_pt2 (pdh1,pddef1,psigma1)	43) Results from 2nd multi-patch regression analysis
multi_def_pt2 (pres1)	44,45) Residual phases
multi_def_pt2 (punw1)	46) Unwrapped phases
atm/pres1_spf_25_1	47) Filtered residual phase (atmosphere)
multi_def_pt2/punw1_noatm	48) Atmosphere w/o auto-interferogram
multi_def_pt2/punw1_noatm_spf	49) Noise-free atmosphere
pdispl	50) Deformation time series